Method of sepiction of variable:
Q Obtain the solution of wave of

$$
\frac{\partial^{2} y}{\partial t^{2}}=c^{2} \frac{\partial^{2} y}{\partial x^{2}}
$$

Such that $y=P_{0} \operatorname{cospt}\left(P_{0}\right.$ is constant) isle $x=l$
and $y=0$ sha $x=0$
and $y=0$ shoa $x=0$

$$
\begin{equation*}
\text { ST L } \frac{\partial^{2} y}{\partial t^{2}}=c^{2} \frac{\partial^{2} y}{\partial x^{2}} \tag{1}
\end{equation*}
$$

Let $y=X T$ where $X$ is the faction of $x$ only and
$T$ is he function of $t$ only

$$
\frac{\partial y}{\partial t}=x \frac{\partial T}{d t} \text { and } \frac{\partial y}{\partial x^{2}}=T \frac{d^{2} x}{d x^{2}}
$$

Substituting these values in the given equation, beget

$$
x \frac{d^{2} T}{d t^{2}}=c^{2} T \frac{d^{2} x}{d x^{2}}
$$

By sepreting the variables, we get

$$
\frac{1}{c^{2} T} \frac{d^{2} T}{d t^{2}}=\frac{1}{x} \frac{d^{2} x}{d x^{2}}=k
$$

Auxillery equation are

$$
D^{2}-k c^{2} T=0 \text { and } \frac{d^{2} x}{d x^{2}}-k x=0
$$

Method of sepiction of variable:
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Auxillery equation are

$$
D^{2}-k c^{2} T=0 \text { and } \frac{d^{2} x}{d x^{2}}-k x=0
$$

Auxiliary equations are

according to sandal condition $k<0$

$$
\begin{align*}
& y=T x \\
& y=\left(C_{5} \cos c \sqrt{k} 6+C_{6} \sin c \sqrt{k} t\right)\left(C_{7} \cos \sqrt{k} x+\right. \\
& \left.C_{8} \sin \sqrt{k} x\right) \tag{2}
\end{align*}
$$

Noes putting $y=0$ and $x=0$ in ep (2), we get

$$
\begin{aligned}
0 & =\left(c_{1} \cos c \sqrt{k} t+c_{2} \sin c \sqrt{k} t\right) c_{3} \\
\Rightarrow c_{3} & =0
\end{aligned}
$$

On putting $C_{3}=0$ in (2), we get

Auxiliary equations are

according to sandal condition $k<0$

$$
\begin{align*}
& y=T x \\
& y=\left(C_{5} \cos c \sqrt{k} 6+C_{6} \sin c \sqrt{k} t\right)\left(C_{7} \cos \sqrt{k} x+\right. \\
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\Rightarrow c_{3} & =0
\end{aligned}
$$

On putting $C_{3}=0$ in (2), we get

## The Inverse Laplace Transform

1. If $\mathcal{L}\{f(t)\}=F(s)$, then the inverse Laplace transform of $F(s)$ is

$$
\begin{equation*}
\mathcal{L}^{-1}\{F(s)\}=f(t) \tag{1}
\end{equation*}
$$

The inverse transform $\mathcal{L}^{-1}$ is a linear operator:

$$
\begin{equation*}
\mathcal{L}^{-1}\{F(s)+G(s)\}=\mathcal{L}^{-1}\{F(s)\}+\mathcal{L}^{-1}\{G(s)\}, \tag{2}
\end{equation*}
$$

and

$$
\begin{equation*}
\mathcal{L}^{-1}\{c F(s)\}=c \mathcal{L}^{-1}\{F(s)\} \tag{3}
\end{equation*}
$$

for any constant $c$.
2. Example: The inverse Laplace transform of

$$
U(s)=\frac{1}{s^{3}}+\frac{6}{s^{2}+4}
$$

is

$$
\begin{align*}
u(t) & =\mathcal{L}^{-1}\{U(s)\} \\
& =\frac{1}{2} \mathcal{L}^{-1}\left\{\frac{2}{s^{3}}\right\}+3 \mathcal{L}^{-1}\left\{\frac{2}{s^{2}+4}\right\} \\
& =\frac{s^{2}}{2}+3 \sin 2 t . \tag{4}
\end{align*}
$$

3. Example: Suppose you want to find the inverse Laplace transform $x(t)$ of

$$
X(s)=\frac{1}{(s+1)^{4}}+\frac{s-3}{(s-3)^{2}+6} .
$$

Just use the shift property (paragraph 11 from the previous set of notes):

$$
\begin{aligned}
x(t) & =\mathcal{L}^{-1}\left\{\frac{1}{(s+1)^{4}}\right\}+\mathcal{L}^{-1}\left\{\frac{s-3}{(s-3)^{2}+6}\right\} \\
& =\frac{e^{-t} t^{3}}{6}+e^{3 t} \cos \sqrt{6} t
\end{aligned}
$$

4. Example: Let $y(t)$ be the inverse Laplace transform of

$$
Y(s)=\frac{e^{-3 s} s}{s^{2}+4}
$$

Don't worry about the exponential term. Since the inverse transform of $s /\left(s^{2}+4\right)$ is $\cos 2 t$, we have by the switchig property (paragraph 12 from the previous notes):

$$
\begin{aligned}
y(t) & =\mathcal{L}^{-1}\left\{\frac{e^{-3 s} s}{s^{2}+4}\right\} \\
& =H(t-3) \cos 2(t-3)
\end{aligned}
$$

5. Example: Let $G(s)=s\left(s^{2}+4 s+5\right)^{-1}$. The inverse transform of $G(s)$ is

$$
\begin{align*}
g(t) & =\mathcal{L}^{-1}\left\{\frac{s}{s^{2}+4 s+5}\right\} \\
& =\mathcal{L}^{-1}\left\{\frac{s}{(s+2)^{2}+1}\right\} \\
& =\mathcal{L}^{-1}\left\{\frac{s+2}{(s+2)^{2}+1}\right\}-\mathcal{L}^{-1}\left\{\frac{2}{(s+2)^{2}+1}\right\} \\
& =e^{-2 t} \cos t-2 e^{-2 t} \sin t . \tag{5}
\end{align*}
$$

6. There is usually more than one way to invert the Laplace transform. For example, let $F(s)=\left(s^{2}+4 s\right)^{-1}$. You could compute the inverse transform of this function by completing the square:

$$
\begin{align*}
f(t) & =\mathcal{L}^{-1}\left\{\frac{1}{s^{2}+4 s}\right\} \\
& =\mathcal{L}^{-1}\left\{\frac{1}{(s+2)^{2}-4}\right\} \\
& =\frac{1}{2} \mathcal{L}^{-1}\left\{\frac{2}{(s+2)^{2}-4}\right\} \\
& =\frac{1}{2} e^{-2 t} \sinh 2 t \tag{6}
\end{align*}
$$

You could also use the partial fraction decomposition (PFD) of $F(s)$ :

$$
F(s)=\frac{1}{s(s+4)}=\frac{1}{4 s}-\frac{1}{4(s+4)}
$$

Therefore,

$$
\begin{align*}
f(t)=\mathcal{L}^{-1}\{F(s)\} & \\
& =\mathcal{L}^{-1}\left\{\frac{1}{4 s}\right\}-\mathcal{L}^{-1}\left\{\frac{1}{4(s+4)}\right\} \\
& =\frac{1}{4}-\frac{1}{4} e^{-4 t} \\
& =\frac{1}{2} e^{-2 t} \sinh 2 t \tag{7}
\end{align*}
$$

7. Example: Compute the inverse Laplace transform $q(t)$ of

$$
Q(s)=\frac{3 s}{\left(s^{2}+1\right)^{2}}
$$

You could compute $q(t)$ by partial fractions, but there's a less tedious way. Note that

$$
Q(s)=-\frac{3}{2} \frac{d}{d s} \frac{1}{s^{2}+1}
$$

Hence,

$$
\begin{align*}
q(t) & =\mathcal{L}^{-1}\{Q(s)\} \\
& =-\frac{3}{2} \mathcal{L}^{-1}\left\{\frac{d}{d s} \frac{1}{s^{2}+1}\right\} \\
& =\frac{3}{2} t \sin t \tag{8}
\end{align*}
$$

8. Definition: The convolution of functions $f(t)$ and $g(t)$ is

$$
\begin{equation*}
(f * g)(t)=\int_{0}^{t} f(t-v) g(v) d v \tag{9}
\end{equation*}
$$

As we showed in class, the convolution is commutative:

$$
\begin{equation*}
(f * g)(t)=\int_{0}^{t} f(t-v) g(v) d v=\int_{0}^{t} g(t-v) f(v) d v=(g * f)(t) \tag{10}
\end{equation*}
$$

9. Example: Let $f(t)=t$ and $g(t)=e^{t}$. The convolution of $f$ and $g$ is

$$
\begin{align*}
(f * g)(t) & =\int_{0}^{t}(t-v) e^{v} d v \\
& =t \int_{0}^{t} e^{v} d v-\int_{0}^{t} v e^{v} d v \\
& =e^{t}-t-1 \tag{11}
\end{align*}
$$

10. Proposition: (The Convolution Theorem) If the Laplace transforms of $f(t)$ and $g(t)$ are $F(s)$ and $G(s)$ respectively, then

$$
\begin{equation*}
\mathcal{L}\{(f * g)(t)\}=F(s) G(s) \tag{12}
\end{equation*}
$$

that is,

$$
\begin{equation*}
\mathcal{L}^{-1}\{F(s) G(s)\}=(f * g)(t) . \tag{13}
\end{equation*}
$$

11. Suppose that you want to find the inverse transform $x(t)$ of $X(s)$. If you can write $X(s)$ as a product $F(s) G(s)$ where $f(t)$ and $g(t)$ are known, then by the above result, $x(t)=(f * g)(t)$.
12. Example: Consider the previous example: Find the inverse transform $q(s)$ of

$$
Q(s)=\frac{3 s}{\left(s^{2}+1\right)^{2}}
$$

Write $Q(s)=F(s) G(s)$, where

$$
F(s)=\frac{3}{s^{2}+1}
$$

and

$$
G(s)=\frac{s}{s^{2}+1}
$$

The inverse transforms are of $F(s)$ and $G(s)$ are $f(t)=3 \sin t$ and $g(t)=\cos t$. Therefore

$$
\begin{align*}
q(s) & =\mathcal{L}^{-1}\{Q(s)\} \\
& =\mathcal{L}^{-1}\{F(s) G(s)\} \\
& =(f * g)(t) \\
& =3 \int_{0}^{t} \sin (t-v) \cos v d v \tag{14}
\end{align*}
$$

Even if you stop here, you at least have a fairly simple, compact expression for $q(s)$. To do the integral (14), use the trigonometric identity

$$
\sin A \cos B=\frac{\sin (A+B)+\sin (A-B)}{2}
$$

With this, (14) becomes

$$
\begin{align*}
q(s) & =\frac{3}{2} \int_{0}^{t} \sin t d v+\int_{0}^{t} \sin (t-2 v) d v \\
& =\frac{3}{2} t \sin t \tag{15}
\end{align*}
$$

13. Example: Find the inverse Laplace transform $x(t)$ of the function

$$
X(s)=\frac{1}{s\left(s^{2}+4\right)}
$$

If you want to use the convolution theorem, write $X(s)$ as a product:

$$
X(s)=\frac{1}{s} \frac{1}{s^{2}+4}
$$

Since

$$
\mathcal{L}^{-1}\left\{\frac{1}{s}\right\}=1
$$

and

$$
\mathcal{L}^{-1}\left\{\frac{1}{s^{2}+4}\right\}=\frac{1}{2} \sin 2 t
$$

we have

$$
\begin{aligned}
x(t) & =\frac{1}{2} \int_{0}^{t} \sin 2 v d v \\
& =\frac{1}{4}(1-\cos 2 t)
\end{aligned}
$$

You could also use the PFD:

$$
X(s)=\frac{1}{4 s}-\frac{s}{4\left(s^{2}+4\right)}
$$

Therefore,

$$
\begin{aligned}
x(t) & =\mathcal{L}^{-1}\left\{\frac{1}{4 s}\right\}-\mathcal{L}^{-1}\left\{\frac{s}{4\left(s^{2}+4\right)}\right\} \\
& =\frac{1}{4}(1-\cos 2 t)
\end{aligned}
$$

# Ordinary Differential Equations-Lecture Notes 

Eugen J. Ionascu

(c) Draft date April 25, 2006

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## Preface

These lecture notes were written during the two semesters I have taught at the Georgia Institute of Technology, Atlanta, GA between fall of 2005 and spring of 2006. I have used the well known book of Edwards and Penny [4]. Some additional proofs are introduced in order to make the presentation as comprehensible as possible. Even that the audience was mostly engineering major students I have tried to teach this course as for mathematics majors.

I have used the book of F. Diacu [3] when I taught the Ordinary Differential Equation class at Columbus State University, Columbus, GA in the Spring of 2005. This work determined me to have a closer interest in this area of mathematics and it influenced a lot my teaching style.

## Chapter 1

## Solving various types of differential equations

### 1.1 Lecture I

Quotation: "The mind once expanded to the dimensions of larger ideas, never returns to its original size." Oliver Wendell Holms

Notions, concepts, definitions, and theorems: Definition of a differential equations, definition of a classical solution of a differential equation, classification of differential equations, an example of a real world problem modeled by a differential equations, definition of an initial value problem.

If we would like to start with some examples of differential equations, before we give a formal definition, let us think in terms of the main classes of functions that we studied in Calculus such as polynomial, rational, power functions, exponential, logarithmic, trigonometric, and inverse of trigonometric functions, what will be some equations that will be satisfied by these classes of functions or at least some of these type of functions?

For polynomials we can think of a differential equation of the type:

$$
\begin{equation*}
\frac{d^{n} y}{d x^{n}}(x)=0 \text { for all } x \text { in some interval, } \tag{1.1}
\end{equation*}
$$

(with $n \in \mathbb{N}$ ) whose "solutions" would obviously include any arbitrary polynomial function $y$ of $x$ with degree at most $n-1$. In other words $y(x)=a_{1} x^{n-1}+a_{2} x^{n-2}+$ $\ldots+a_{n}$ is a polynomial function that satisfies (1.1). Let us notice that there are $n$ constants that we can choose as we like in the expression of $y$.

## 4CHAPTER 1. SOLVING VARIOUS TYPES OF DIFFERENTIAL EQUATIONS

Let us say we consider a power function whose rule is given by $y(x)=x^{\alpha}$ with $\alpha \in \mathbb{R}$. Then by taking its derivative we get $\frac{d y}{d x}(x)=\alpha x^{\alpha-1}$, we see that we can make up a differential equation, in terms of only the function itself, that this function will satisfy

$$
\begin{equation*}
\frac{d y}{d x}(x)=\frac{\alpha y(x)}{x}, \text { for } x \text { in some interval contained in }(0, \infty) \tag{1.2}
\end{equation*}
$$

For a rational function, lets say $y(x)=\frac{x+1}{2 x+1}, x \in \mathbb{R} \backslash\left\{-\frac{1}{2}\right\}$, if we take the derivative of $y(x)$, we get $\frac{d y}{d x}(x)=-\frac{1}{(2 x+1)^{2}}$ and since $y(x)=\frac{1}{2}+\frac{1}{2(2 x+1)}$ a relatively natural way to involve the derivative and the function will be:

$$
\begin{equation*}
\frac{d y}{d x}(x)=-(2 y(x)-1)^{2} \tag{1.3}
\end{equation*}
$$

For a general rational function it is not going to be easy to find a corresponding differential equation that will be of the same type as before. These equations will be called later separable equations. Most of the time the independent variable is dropped from the writing and so a differential equation as (1.3) can be rewritten as $y^{\prime}=-(2 y-1)^{2}$.

Suppose we are interested in finding a similar differential equation satisfied by an exponential function. It is easily seen that such a candidate is:

$$
\begin{equation*}
\frac{d y}{d x}(x)=k y(x) \tag{1.4}
\end{equation*}
$$

where $k$ is some non-zero real valued constant.
If we take $f(x)=\sin x$ and $g(x)=\cos x$ then we see that these two functions satisfy the following system of differential equations:

$$
\left\{\begin{array}{l}
\frac{d f}{d x}(x)=g(x)  \tag{1.5}\\
\frac{d g}{d x}(x)=-f(x)
\end{array}\right.
$$

Now we are going to consider $f(x)=\arctan x, x \in\left(-\frac{\pi}{2}, \frac{\pi}{2}\right)$. Because the derivative of $f$ is $f^{\prime}(x)=\frac{1}{1+x^{2}}$ we can build a differential equation that $f$ will satisfy:

$$
\begin{equation*}
f^{\prime}(x)=\frac{1}{1+(\tan f(x))^{2}} \tag{1.6}
\end{equation*}
$$

Finally a function of two variables such as $f(x, y)=x^{2}-y^{2}, x, y \in \mathbb{R}^{2}$ satisfies:

$$
\begin{equation*}
\frac{\partial f^{2}}{\partial x^{2}}+\frac{\partial f^{2}}{\partial y^{2}}=0 \tag{1.7}
\end{equation*}
$$

Hoping that we have enough examples we will give a formal definition:
Definition 1.1.1. A differential equation, shortly $D E$, is a relationship between a finite set of functions and its derivatives.

Depending upon the domain of the functions involved we have ordinary differential equations, or shortly ODE, when only one variable appears (as in equations (1.1)-(1.6)) or partial differential equations, shortly PDE, (as in (1.7)).

From the point of view of the number of functions involved we may have one function, in which case the equation is called simple, or we may have several functions, as in (1.5), in which case we say we have a system of differential equations.

Taking in account the structure of the equation we may have linear differential equation when the simple DE in question could be written in the form:

$$
\begin{equation*}
a_{0}(x) y^{(n)}(x)+a_{1}(x) y^{(n-1)}(x)+\ldots+a_{n}(x)=F(x), \tag{1.8}
\end{equation*}
$$

or if we are dealing with a system of DE or PDE , each equation should be linear as before in all the unknown functions and their derivatives. In case such representations are not possible we are saying that the DE is non-linear. If the function $F$ above is zero the linear equation is called homogenous. Otherwise, we are dealing with a non-homogeneous linear DE. If the differential equation does not contain (depend) explicitly of the independent variable or variables we call it an autonomous DE . As a consequence, the $\mathrm{DE}(1.2)$, is non-autonomous. As a result of these definitions the DE's (1.1), (1.2), (1.4), (1.5) and (1.7) are homogeneous linear differential equations.

The highest derivative that appears in the DE gives the order. For instance the equation (1.1) has order $n$ and (1.7) has order two.


Figure 1.1: The man and his dog

Definition 1.1.2. We say that a function or a set of functions is a solution of a differential equation if the derivatives that appear in the DE exist on a certain domain and the DE is satisfied for all all the values of the independent variables in that domain.

This concept is usually called a classical solution of a differential equation. The domain for ODE is usually an interval or a union of intervals.

Next we are going to deal with an example of DE that has rather a more real world flavor than a theoretical one as the ones we have encountered so far.

Problem 1.1.3. We have a man and his dog running on a straight beach. At a given point in time the dog is 12 m from his owner who starts running in a direction perpendicular to the beach with a certain constant speed. The dog runs twice as fast and always toward his owner. The question is where are they going to meet?

Solution: Let us assume that the dog runs on a path given by the graph of the function $f$ as in the figure above. Suppose that after a certain time $t$ the dog is at a position $(x, f(x))$ and the man is on the $y$-axis at $(0, v t)$ where $v$ is his speed in meters per second (assumed constant). The fact that the dog is running toward the man at every time is going to give us a DE . This condition can be translated into the fact that the tangent line to the graph of $f$ at $(x, f(x))$ passes through $(0, v t)$.

The equation of the tangent line is $Y-f(x)=f^{\prime}(x)(X-x)$ and so the intersection with the $y$ axis is $v t=f(x)-f^{\prime}(x) x$. Let us assume the distance between the dog and the man is originally $a$ ( $a=12$ in this problem). The dog is running the distance $\int_{x}^{a} \sqrt{1+f^{\prime}(s)^{2}} d s$ which is suppose to be twice as big as $v t$. Hence we get
the equation $\int_{x}^{a} \sqrt{1+f^{\prime}(s)^{2}} d s=2\left(f(x)-f^{\prime}(x) x\right)$ in $x$, for every $x$ in the interval $(0, a)$. Differentiating with respect to $x$ we obtain: $-\sqrt{1+f^{\prime}(x)^{2}}=2\left(-x f^{\prime \prime}(x)\right)$ or

$$
\frac{f^{\prime \prime}(x)}{\sqrt{1+f^{\prime}(x)^{2}}}=\frac{1}{2 x}, x>0
$$

Integrating with respect to $x$ will give

$$
\ln \left(f^{\prime}(x)+\sqrt{1+f^{\prime}(x)^{2}}\right)=\ln k \sqrt{x}
$$

for some constant $k$. Since $f^{\prime}(a)=0$ we determine $k$ right away to be $k=\frac{1}{\sqrt{a}}$. Solving for $f^{\prime}(x)$ will give $f^{\prime}(x)=\left(\frac{\sqrt{x}}{\sqrt{a}}-\frac{\sqrt{a}}{\sqrt{x}}\right) / 2$. Integrating again with respect to $x$ we obtain $f(x)=\frac{x \sqrt{x}}{3 \sqrt{a}}-\sqrt{a x}+C$ for another constant $C$. Since $f(a)=0$ we get $C=2 a / 3$. Therefore $f(0)=\frac{2 a}{3}$. So, the dog and its owner are going to meet at 8 meters from the point the "race" began.

In general we like to know whether or not, of course under certain circumstances, a DE has a unique solution so that we may talk about the solution of the DE. This thing may happen, but in the general situation, this is hardly the case without some extra conditions such as initial conditions. In order to accomplish such a thing we usually consider the so called initial value problem which takes the following form when we are dealing with a single, first order ODE:

$$
\left\{\begin{array}{l}
\frac{d y}{d x}(x)=f(x, y(x)), x \in I  \tag{1.9}\\
y\left(x_{0}\right)=y_{0}, x_{0} \in I, y_{0} \in J, I \times J \subset \operatorname{Domain}(f),
\end{array}\right.
$$

where $I$ and $J$ are open intervals. For a system of ODE or a higher order ODE the initial value problem associated to it takes a slightly different form. We are going to see those at the appropriate time.

Homework: Problems 1-12, 27-31, 34, 37-43, 47 and 48, pages 8-9.

### 1.2 Lecture II

Quotation: " An idea which can be used once is a trick. If it can be used more than once it becomes a method." George Polya and Gabor Szego

## 8CHAPTER 1. SOLVING VARIOUS TYPES OF DIFFERENTIAL EQUATIONS

Notions, concepts, definitions, and theorems: Methods of study for differential equations, a/the general solution of a differential equation, particular solution, velocity and acceleration example, slope field and solution curves, existence theorem and an existence and uniqueness theorem.

We say that differential equations are studied by quantitative or exact methods when they can be solved completely (i.e. all the solutions are known and could be written in closed form in terms of elementary functions or sometime special functions (or inverses of these type of functions). This reduces the study of DE to the study of functions of one or more real variables given in an explicit or implicit way.

As an example let us consider the equation in Exercise 4, page 16

$$
\begin{equation*}
\frac{d y}{d x}=\frac{1}{x^{2}} \tag{1.10}
\end{equation*}
$$

If we rewrite the equation as $\frac{d}{d x}\left(y(x)+\frac{1}{x}\right)=0$ we see that we are dealing with a function whose derivative is zero. If we talk about solutions defined on a interval, the Mean Value Theorem from Calculus, tells us that $y(x)+\frac{1}{x}=C$ for some constant $C$ and for all $x \in I, I$ an interval not containing zero. Therefore any solution (as long as we consider the domains of solutions intervals like $I$ ) of the DE in (1.10) is of the form $y(x)=C-\frac{1}{x}$ for $x \in I$. So, we were able to solve the equation (1.10) exactly. To finish the Exercise 4, page 16, we determine $C$ such that the initial value condition, $y(1)=5$, is satisfied too. This gives $C=6$ and $y(x)=\frac{6 x-1}{x}$ for all $x \in I$.

There are also some other type of methods, called analytical methods or qualitative methods in which one can describe the behavior of a DE's solution such as, existence, uniqueness, stability, chaotic or asymptotic character, boundlessness, periodicity, etc. without actually solving it exactly. This is an important and relatively new step in the theory of DE. Important because most of the differential equations cannot be solved exactly and relatively new because it has all started mainly at the end of the 19th century. One of the mathematicians that pioneered in this area was Henri Poincaré.

We can add to the list another type of methods for studying DE to which are numerical methods. These methods mainly involve the use of a computer, a special designed software following the procedure given by an approximation algorithm. In this part of mathematics one studies the algorithms and the error analysis involved in approximating the solution of a DE that in general cannot be studied with exact methods. Very good approximations could be obtained most of the time only locally (not too far from the initial value point).

Definition 1.2.1. A general solution of a DE of order $n$ is a solution which is given in terms of $n$ independent parameters. A particular solution of a DE (relative to a
general solution) is a solution which could be obtained from that general solution by simply choosing specific values of the parameters involved.

If all the solutions of DE are particular solutions obtained from a general solution then this is referred to as the general solution.

As an example, we are going to show later that the general solution of the second order linear equation $y^{\prime \prime}+4 y^{\prime}+4=0$ is $y(x)=\left(C_{1}+C_{2} x\right) e^{-2 x}$ for all $x \in I$.

Another example is the particular case of the movement of a body under the action of a constant force according to Newton's second law mechanics: $m \vec{a}=\vec{F}$. This implies that if we denote the position of the body relative to a fixed point in space by $x(t)$ (the dependent variable here being the time t , and units are fixed but not specified). Integrating twice the equation

$$
\begin{equation*}
\frac{d^{2} x}{d t^{2}}(t)=a \tag{1.11}
\end{equation*}
$$

we get

$$
\begin{equation*}
x(t)=a t^{2} / 2+v_{0} t+x_{0}, t \in \mathbb{R} \tag{1.12}
\end{equation*}
$$

where $a$ is the constant acceleration, $v_{0}$ is the initial velocity and $x_{0}$ is the initial position. We can look at this as the general solution of the equation (1.11).

As an application let us work the following problem from the book (No. 36, page 17).

Problem 1.2.2. If a woman has enough "spring" in her legs to jump vertically to a hight of 2.25 ft on the earth, how high could she jump on the moon, where the surface gravitational acceleration is (approximately) $5.3 \frac{\mathrm{ft}}{\mathrm{s}^{2}}$ ?

Solution: From the equation (1.12) we see that whatever her speed is initially, say $v_{0}$, on earth, she is going to get to a maximum height $h=v_{0} t-g t^{2} / 2$ where $t$ is given by the condition that $d x / d t=0$ or $v_{0}-g t=0$. Hence, we get $h=$ $v_{0} \frac{v_{0}}{g}-g\left(\frac{v_{0}}{g}\right)^{2} / 2$ or $h=\frac{v_{0}^{2}}{2 g}$. (Notice that, so far, this is basically solving Problem 35, page 17). From this we can solve for $v_{0}$ and obtain $v_{0}=\sqrt{2 g h}$. On the moon she is going to use the same initial velocity (this is saying that the energy is the same). Hence $h_{\text {max }}=\frac{v_{0}^{2}}{2 g_{m}}=\frac{2 g h}{2 g_{m}}=\frac{g h}{g_{m}}$ or $h_{\text {max }}=\frac{32 \times 2.25}{5.3}=13.58 \mathrm{ft}$.

From now on in this Chapter we are going to concentrate on first order, single, ODE of the form:

$$
\begin{equation*}
y^{\prime}=f(x, y) \quad \text { or } \quad \frac{d y}{d x}(x)=f(x, y(x)) \tag{1.13}
\end{equation*}
$$

We are trying to solve for $y$ as a function of $x$. The best thing here is to look at an example. Let us take the example from the book, page 18, i.e. $y^{\prime}=x^{2}+y^{2}$ whose solution is not expressible in terms of simple functions. If we try Maple on this we get

$$
y(x)=-x \frac{\operatorname{Bessel} J\left(-3 / 4, x^{2} / 2\right) C+\operatorname{Bessel} Y\left(-3 / 4, x^{2} / 2\right)}{\operatorname{Bessel} J\left(1 / 4, x^{2} / 2\right) C+\operatorname{Bessel} Y\left(1 / 4, x^{2} / 2\right)}
$$

We will learn later about Bessel functions which appear in the above expression of the general solution. This expression is useful if we want to do numerical calculations since Bessel functions can be expressed in terms of power series.

On the other hand if we imagine that at each point of coordinates $(x, y)$ in the xy-plane we draw a little unit vector of slope $f(x, y)=x^{2}+y^{2}$ then we get the picture below:


Vector / slope field
and we kind of see how the solution curves should look like. We are drawing next (of course, using a special tool like Maple) the solution curve passing through ( $0,-1$ ) for instance.


Figure 3

It seems like the vector field in Figure 2 defines uniquely the solutions curves. We are asking then the two fundamental questions in most of the mathematics when dealing with equations:

- When do we have at least a solution for (1.13)?
- If there exist a solution of (1.13) is that the only one?

The first problem is usually referred as existence problem and the second as the uniqueness problem. In general, in order to obtain existence for the DE (1.13) we only need continuity for the function $f$ :

Theorem 1.2.3. (Peano) If the function $f(x, y)$ is continuous on a rectangle $\mathcal{R}=\{(x, y) \mid a<x<b, c<y<d\}$, and if $\left(x_{0}, y_{0}\right)$ in $\mathcal{R}$, then the initial value problem

$$
\left\{\begin{array}{l}
\frac{d y}{d x}(x)=f(x, y(x))  \tag{1.14}\\
y\left(x_{0}\right)=y_{0},
\end{array}\right.
$$

has a solution in the neighborhood of $x_{0}$.
We need more than continuity in order to obtain uniqueness:
Theorem 1.2.4. (Cauchy) Let $f(x, y)$ be continuous such that the derivative $\frac{\partial f}{\partial y}(x, y)$ exists and it is continuous on a rectangle $\mathcal{R}=\{(x, y) \mid a<x<b, c<y<d\}$, and if $\left(x_{0}, y_{0}\right)$ in $\mathcal{R}$, then the initial value problem (1.14) has a solution which is unique on an interval around $x_{0}$.

As an example we will look at Problem 30, page 28.
Problem 1.2.5. Verify that if $c$ is a constant, then the function defined piecewise by

$$
y(x)=\left\{\begin{array}{l}
1 \text { if } x \leq c  \tag{1.15}\\
\cos (x-c) \text { if } c<x<c+\pi \\
-1 \text { if } x \geq c+\pi
\end{array}\right.
$$

satisfies the differential equation $y^{\prime}=-\sqrt{1-y^{2}}$ for all $x \in \mathbb{R}$. Determine how many different solutions (in terms of $a$ and b) the initial value problem

$$
\left\{\begin{array}{l}
y^{\prime}=-\sqrt{1-y^{2}} \\
y(a)=b
\end{array}\right.
$$

has.

Solution: It is not hard to see that the function $y$ given in (1.15) is differentiable at each point and its derivative is actually

$$
y^{\prime}(x)=\left\{\begin{array}{l}
0 \text { if } x \leq c  \tag{1.16}\\
-\sin (x-c) \text { if } c<x<c+\pi \\
0 \text { if } x \geq c+\pi
\end{array}\right.
$$

Hence if $x \leq c$ or $x \geq c+\pi$ then the equation $y^{\prime}=-\sqrt{1-y^{2}}$ is satisfied because $y^{\prime}=-\sqrt{1-y^{2}}=0$. If $c<x<c+\pi$ then $0<x-c<\pi$ and then $\sin (x-c)$ is positive, which implies $\sqrt{1-\cos (x-c)^{2}}=\sin (x-c)$ and so the equation is satisfied in this case also.

For the second part of this problem, it is clear that if $|b|>1$ we do not have any solution because $\sqrt{1-y(a)^{2}}$ is not a real number. If $b=1$, we have infinitely many solutions, by just taking $c>a$, then the $y(x)$ defined by (1.15) is a solution of the initial value problem in discussion. Similarly we get infinitely many solutions if $b=-1$, in which case we have to take $c+\pi<a$ or $c<a-\pi$. If $-1<b<1$ we have a unique solution around the point $a$ by Cauchy's Theorem but not on $\mathbb{R}$.

## Homework:

Section 1.2 pages 16-18: 1-5, 11-15, 35 and 36;
Section 1.3 pages 26-29: 11-15, 27-33.

### 1.3 Lecture III

Quotation: " Hardy, Godfrey H. (1877-1947) I believe that mathematical reality lies outside us, that our function is to discover or observe it, and that the theorems which we prove, and which we describe grandiloquently as our "creations," are simply the notes of our observations. A Mathematician's Apology, London, Cambridge University Press, 1941. "
Type of equations which can be solved with exact methods, notions, real world applications: Separable equations, implicit solution, singular solution, natural growth or decay equation and general solution, Newton's law of cooling or heating and its general solution, Torricelli's law, liner first-order equations and the general solution, mixture problems.

One of the simplest cases in which the general solution could be found is the so called separable differential equations. This is an equation of the form

$$
\begin{equation*}
y^{\prime}=f(x) g(y) \tag{1.17}
\end{equation*}
$$

where $f$ and $g$ are, let us say continuous functions on a their domains that each contains an interval. Let us assume that $g$ is not a constant. Then the function $g$ is not zero for a set containing an interval too, say $I$. Then the equation (1.17) can be written equivalently as $\frac{y^{\prime}}{g(y)}=f(x)$ if we assume that $y \in I$. We are going to treat the situation $g(y)=0$ separately. Suppose $G(u)$ is an antiderivative of $\frac{1}{g(u)}$ on $I$, and $F$ and antiderivative of $f$. Then the equation in question is equivalent to $\frac{d}{d x}(G(y(x))-F(x))=0$ which means that the general solution should be

$$
\begin{equation*}
G(y(x))-F(x)=C . \tag{1.18}
\end{equation*}
$$

Most of the time, this equation cannot be solved in terms of $y(x)$ and we just say in that case that the solution, $y(x)$, is given implicitly.

The case $g\left(y_{0}\right)=0$, will give solutions $y(x)=y_{0}$ which are usually called singular solutions unless (1.18) gives this solution for some value of the constant (parameter) $C$.

As an example let us take a look at Newton's law of cooling or heating: the time rate of change of the temperature $T(t)$ of a body immersed in a medium of constant temperature $M$ is proportional to the difference $M-T(t)$.

This translates into

$$
\begin{equation*}
T^{\prime}(t)=k(M-T(t)) \tag{1.19}
\end{equation*}
$$

for some positive constant, which is a separable equation. Equivalently, this can be written as $\frac{T^{\prime}(t)}{T(t)-M}=-k$ assuming that $T(t) \neq M$ at any time $t$. Integrating we obtain $\ln |T(t)-M|=-k t+C$ which implies $|T(t)-M|=e^{-k t} e^{C}$. If we make $t=0$ we get that $e^{C}= \pm\left(T_{0}-M\right)$ where $T_{0}$ is the initial temperature of the body. Then the expression of $T(t)$ becomes

$$
\begin{equation*}
T(t)=M+\left(T_{0}-M\right) e^{-k t} \tag{1.20}
\end{equation*}
$$

Let us observe that the equation (1.19) admits only one other solution, namely the constant function $T(t)=M, t \in \mathbb{R}$, and that this solution is actually contained in (1.20) by simply taking $T_{0}=M$. The equality above then is the general solution of (1.19) As an application of (1.20), let us take and solve Problem 43, page 42.

Problem 1.3.1. A pitcher of buttermilk initially at $25^{\circ} C$ is to be cooled by setting it on the front porch, where the temperature is $0^{\circ}$ C. Suppose that the temperature of the buttermilk has dropped to $15^{\circ}$ after 20 minutes. When will it be at $5^{\circ}$ ?

Solution: Using the formula (1.20), twice, we get $T(20)=25 e^{-20 k}=15$ which gives $k=\frac{1}{20} \ln (5 / 3)$ and so $T(t)=25 e^{-k t}=5$. This last equation can then be solved for $t$ to obtain $t=\frac{\ln 5}{k}=20 \frac{\ln 5}{\ln 5 / 3} \approx 63$ minutes.

Another application of separable DE is Torricelli's law: suppose that a water tank has a hole with area a at its bottom and cross sectional area $A(y)$ for each height $y$, then the water flows in such a way the following $D E$ is satisfied:

$$
\begin{equation*}
A(y) \frac{d y}{d t}=-k \sqrt{y} \tag{1.21}
\end{equation*}
$$

where $k=a \sqrt{2 g}$ and $g$ is the the gravitational acceleration.
As an example of this situation let's take problem 62, page 43.
Problem 1.3.2. Suppose that an initially full hemispherical water tank of radius 1 $m$ has its flat side as its bottom. It has a bottom hole of radius 1 cm . If this bottom hole is opened at 1 P.M., when will the tank be empty?

Solution: On the figure below we see that in order to calculate the cross sectional area $A(y)$ corresponding to height $y$ we need to apply the Pythagorean theorem: $A(y)=\pi\left(1-y^{2}\right)$. Hence the equation that we get is $\pi \frac{1-y^{2}}{\sqrt{y}} \frac{d y}{d t}=-\pi \frac{1}{10000} \sqrt{2 g}$.


Figure 4
Integrating with respect to $t$ we get $\frac{2}{5} y^{\frac{5}{2}}-2 y^{\frac{1}{2}}=\frac{t \sqrt{2 g}}{10000}+C$. Since at $t=0$ we have $y=1$ the constant $C$ is determined: $C=\frac{-8}{5}$. We are interested to see when is $y=0$. This give $t=\frac{16000}{\sqrt{2 g}} \approx=3614$ second since $g$ is measured here in $\mathrm{m} / \mathrm{s}^{2}$. This is approximately 1 hours so the tank will be empty around 2 P.M. ( 14 seconds after).

### 1.3.1 Linear First Order DE

These equation are equations of the type:

$$
\begin{equation*}
y^{\prime}+P(x) y=Q(x), y\left(x_{0}\right)=y_{0} \tag{1.22}
\end{equation*}
$$

where $P$ and $Q$ are continuous on a given interval $I\left(x_{0} \in I\right)$. In order to solve (1.22), the trick is to multiply both side by $e^{R(x)}$ where $R(x)$ is an antiderivative of $P(x)$ on $I$. This way the equation becomes $\frac{d y}{d x}\left(y(x) e^{R(x)}\right)=Q(x) e^{R(x)}$ which after integration gives $y e^{R(x)}=\int Q(x) e^{R(x)} d x$. So the general solution of (1.22) is

$$
\begin{equation*}
y(x)=e^{-R(x)} \int Q(x) e^{R(x)} d x \tag{1.23}
\end{equation*}
$$

Let us observe that if we have an initial value problem

$$
\left\{\begin{array}{l}
y^{\prime}+P(x) y=Q(x),  \tag{1.24}\\
y\left(x_{0}\right)=y_{0}
\end{array}\right.
$$

where $x_{0} \in I$, then we can take explicitly $R(x)=\int_{x_{0}}^{x} P(t) d t$ and (1.23) becomes

$$
\begin{equation*}
y(x)=y_{0} e^{-R(x)}+\int_{x_{0}}^{x} Q(t) e^{R(t)-R(x)} d t, \quad x \in I \tag{1.25}
\end{equation*}
$$

This proves the following theorem.
Theorem 1.3.3. Given that $P$ and $Q$ are continuous functions on an interval $I$, the initial value problem (1.22) has a uniques solution on I given by (1.25).

As an example let's solve the problem 12, page 54:

$$
\left\{\begin{array}{l}
x y^{\prime}+3 y=2 x^{5}  \tag{1.26}\\
y(2)=1
\end{array}\right.
$$

The equation becomes $y^{\prime}+\frac{3}{x} y^{\prime}=2 x^{4}$. Since $R(x)=\int_{1}^{x} \frac{3}{t} d t=3 \ln x$ we obtain that we need to multiply the equation $\left(y^{\prime}+\frac{3}{x} y^{\prime}=2 x^{4}\right)$ by $e^{R(x)}=x^{3}$. So, $x^{3} y^{\prime}+3 x^{2} y=$ $2 x^{7}$. The left hand side is $\frac{d y}{d x}\left(x^{3} y(x)\right)$, so if we integrate from 2 to $a$ we obtain $a^{3} y(a)-8 y(2)=\int_{2}^{a} 2 x^{7} d x$. Equivalently, $a^{3} y(a)-8=2\left(\frac{a^{8}}{8}-\frac{2^{8}}{8}\right)$. So, the solution of this equation on $I=(0, \infty)$ is $y(a)=\frac{a^{5}}{4}-\frac{56}{a^{3}}$ for $a \in I$.

We are going to work out, as another application, the mixture problem 37 on page 54.

Problem 1.3.4. A 400-gal tank initially contains 100 gal of brine containing 50 lb of salt. Brine containing 1 lb of salt per gallon enters the tank at the rate of $5 \mathrm{gal} / \mathrm{s}$, and the well mixed brine in the tank flows at the rate of $3 \mathrm{gal} / \mathrm{s}$. How much salt will the tank contain when it is full of brine?

Solution: The tank is filling up at a speed of $2 \mathrm{gal} / \mathrm{s}$ and it is needed 300 gallons more to be full. So that is going to happen after 150 seconds. The volume of the brine in the tank after $t$ seconds is $V(t)=100+2 t$. Let us do an analysis similar to that in the book at page 51. Denote the amount of salt in the tank at time $t$ by $y(t)$. We balance the change in salt $y(t+h)-y(t)$ during a small interval of time $h$ in the following way: the difference comes from the amount of salt that is getting in the tank minus the amount that is getting out. The amount that is getting in is $5 h \mathrm{lb} / \mathrm{s}$, if we measure $h$ in seconds. Then the amount that is getting out assuming perfect mixture (instantaneous) is approximately $\frac{y(t)}{V(t)} 3 h \mathrm{lb}$ of salt. So, the balance is $\frac{y(t+h)-y(t)}{h} \approx 5-\frac{3 y(t)}{100+2 t}, t \in[0,150]$. Letting $h$ go to zero, we obtain the initial value problem

$$
\left\{\begin{array}{l}
\frac{d y}{d t}=5-\frac{3 y(t)}{100+2 t}  \tag{1.27}\\
y(0)=50 .
\end{array}\right.
$$

This is a linear equation with initial condition that we solve using the same method as above. We have $R(t)=\int_{0}^{t} \frac{3}{100+2 s} d s=\frac{3}{2} \ln \left(\frac{50+t}{50}\right)$. This means that we need to
multiply by $\left(\frac{50+t}{50}\right)^{\frac{3}{2}}$ both sides of $\frac{d y}{d t}+\frac{3 y(t)}{100+2 t}=5$. We get

$$
\left(\frac{50+t}{50}\right)^{\frac{3}{2}} \frac{d y}{d t}+\frac{3}{100}\left(\frac{50+t}{50}\right)^{\frac{1}{2}} y(t)=5\left(\frac{50+t}{50}\right)^{\frac{3}{2}}
$$

Integrating this last equation with respect to $t$ from 0 to $s$, we obtain:

$$
\left(\frac{50+s}{50}\right)^{\frac{3}{2}} y(s)-50=\frac{2}{50 \sqrt{50}}\left((50+s)^{\frac{5}{2}}-50^{2} \sqrt{50}\right)
$$

Substituting $s=150$ in this last equality, we obtain $y(150)=\frac{800000 \sqrt{2}-12500 \sqrt{2}}{2000 \sqrt{2}} \approx$ 393.75 lb .

## Homework:

Section 1.4 pages 41-44: 1-28, 32, 43, 48, 61, 62 and 64;
Section 1.5 pages 54-56: 11-15, 26-33.

### 1.4 Lecture IV

Quotation: "The highest form of pure thought is in mathematics", Plato

Other type of equations which can be solved with exact methods, notions, real world applications: homogeneous (in terms of variables) differential equations, Bernoulli differential equations, exact differential equations, characterization of exactness, reducible to secondorder DE (dependent variable missing, independent variable missing).

### 1.4.1 Substitution methods

We begin with an example by solving the Problem 55, page 72 .
Problem 1.4.1. Show that the substitution $v=a x+b y+c$ transforms the differential equation $\frac{d y}{d x}=F(a x+b y+c)$ into a separable equation.

Solution: Let us assume that $b \neq 0$. Differentiating the substitution $v=a x+b y+c$ with respect to $x$, we get $\frac{d v}{d x}=a+b \frac{d y}{d x}$ and then the equation becomes $\frac{d v}{d x}=a+b F(v)$ which is a separable equation indeed. If $b=0$ then the equation $\frac{d y}{d x}=F(a x+c)$ is already separable.

As an application lets work problem 18 on page 71: find a general solution of $(x+y) y^{\prime}=1$. We make the substitution $v=x+y$. Then $\frac{d v}{d x}=1+\frac{d y}{d x}$ which turns the original equation into $v\left(v^{\prime}-1\right)=1$ or $v v^{\prime}=v+1$. Observe that one singular solution of this equation is $v(x)=-1$ for all $x \in \mathbb{R}$ which corresponds to $y(x)=-x-1$ for all $x \in \mathbb{R}$.
If $v(x) \neq-1$ for $x$ in some interval we can write the equation as $\frac{v v^{\prime}}{v+1}=1$. Equivalently, in order to integrate let us write this equation as

$$
v^{\prime}-\frac{v^{\prime}}{v+1}=1
$$

Integrating with respect to $x$ we obtain $v-\ln |v+1|=x+C$. Getting back to the original variable this can be written as $x+y-\ln |x+y+1|=x+C$. Notice that this could be simplified to $y=\ln |x+y+1|+C$ which gives $y$ only implicitly. If we want to get rid of the logarithmic function and also of the absolute value function as well, we can exponentiate the last equality to turn it into $e^{y}=k(x+y+1)$ where $k$ is a real constant which is not zero. In order to include the singular solution we can move the constant to the other side and allow it to be zero: $x+y+1=k_{1} e^{y}$, $k_{1} \in \mathbb{R}$. Some solutions curves can be drawn with Maple and we include some here ( $k=1,2$ and $1 / 10$ ).


### 1.5 Homogeneous DE

The homogeneous property referes here to the function $f$ when writing the DE as $y^{\prime}=f(x, y)$. The prototype is actually

$$
y^{\prime}=f(y / x)
$$

The recommended substitution here is $v=y / x$. This implies $y(x)=x v(x)$ so, differentiating with respect to $x$ we obtain $y^{\prime}=v+x v^{\prime}$. Then the original equation becomes $v^{\prime}=(f(v)-v) / x$ which is a separable DE.

In order to check that a differential equation is homogeneous we could substitute $y=v x$ in the expression of $f(x, y)$ and see if the resulting function can be written just in terms of $v$. As an example let's take Exercise 14, page 71.
Problem 1.5.1. Find all solutions of the equation $y y^{\prime}+x=\sqrt{x^{2}+y^{2}}$.
Solution: One can check that this DE is homogeneous. Let $y=x v$, where $v$ is a function of $x$. Then $y^{\prime}=v+x v^{\prime}$ and so our equation becomes $x v\left(v+x v^{\prime}\right)+x=$ $\sqrt{x^{2}+x^{2} v^{2}}$. Assume first that we are working on an interval $I \subset(0, \infty)$. Then the DE simplifies to $x v v^{\prime}=\sqrt{1+v^{2}}-1-v^{2}$. Let us observe that $v(x)=0$ for all $x \in I$ is a solution of this equation. We will see that this is not a singular solution. So, if we assume $v$ is not zero on $I$, say $v>0$, the DE is equivalent to $\frac{v v^{\prime}}{1+v^{2}-\sqrt{1+v^{2}}}=-\frac{1}{x}$ or $\frac{v v^{\prime}}{\sqrt{1+v^{2}}\left(\sqrt{1+v^{2}}-1\right)}=-\frac{1}{x}$.

Using the conjugate we can modify the left hand side to $\frac{v\left(\sqrt{1+v^{2}}+1\right) v^{\prime}}{v^{2} \sqrt{1+v^{2}}}=-\frac{1}{x}$ or

$$
\begin{equation*}
\frac{v^{\prime}}{v}+\frac{v^{\prime}}{v \sqrt{1+v^{2}}}=-\frac{1}{x} . \tag{1.28}
\end{equation*}
$$

If we integrate both sides with respect to $x$ we obtain $\ln (v)+\int \frac{d v}{v \sqrt{1+v^{2}}}=-\ln x+$ $C$. In the indefinite integral we've got, let us do the change of variables: $v=\frac{1}{u}$. This integral becomes $\int \frac{d v}{v \sqrt{1+v^{2}}}=-\int \frac{d u}{\sqrt{1+u^{2}}}=-\ln \left|u+\sqrt{1+u^{2}}\right|+C^{\prime}=\ln |v|-$ $\ln \left(\sqrt{1+v^{2}}+1\right)+C^{\prime}$. Hence the DE (1.28) leads to

$$
\begin{equation*}
\frac{x v^{2}}{\sqrt{1+v^{2}}+1}=k \tag{1.29}
\end{equation*}
$$

where the constant $k$ can take any non-negative value. Using the conjugate again (2) changes into $x\left(\sqrt{1+v^{2}}-1\right)=k$ or $\sqrt{x^{2}+y^{2}}-x=k$. If we solve this for $y$ we obtain

$$
\begin{equation*}
y(x)=\sqrt{(2 x+k) k} \tag{1.30}
\end{equation*}
$$

Let us observe that this function is actually defined on $(-k / 2, \infty)$ if $k \geq 0$. One can go back and check that $v<0$ leads to the choice of

$$
\begin{equation*}
y(x)=-\sqrt{(2 x+k) k}, x \in(-k / 2, \infty) \tag{1.31}
\end{equation*}
$$

The two general solutions (1.30) and (1.31) are all the solutions of the original equation (the singular solution $y(x)=0$ is included in (1.30) for $k=0$.

### 1.6 Bernoulli DE

The general form of these equations is very close to that of linear DE:

$$
\begin{equation*}
y^{\prime}+P(x) y=Q(x) y^{n} \tag{1.32}
\end{equation*}
$$

So, we may assume that $n$ is different of 0 or 1 since these cases lead to DE that we have already studied. The recommended substitution is $v=y^{1-n}$. This implies $y=v^{\frac{1}{1-n}}$ and $y^{\prime}=\frac{1}{1-n} v^{n /(1-n)} v^{\prime}$. This changes the original equation to

$$
\frac{1}{1-n} v^{\frac{n}{1-n}} v^{\prime}+P(x) v^{\frac{1}{1-n}}=Q(x) v^{\frac{n}{1-n}}
$$

which after dividing by $v^{\frac{n}{1-n}}$ (assuming is not zero)

$$
\frac{1}{1-n} v^{\prime}+P(x) v=Q(x)
$$

which is a linear DE.
Let us see how this method works with the exercise 26, page 71 .
Problem 1.6.1. Find all solutions of the equation $3 y^{2} y^{\prime}+y^{3}=e^{-x}$.
Solution: If we put the given equation in the form (1.32) we get $y^{\prime}+\frac{1}{3} y=\frac{e^{-x}}{3} y^{-2}$. This says that $n=-2$ and so we substitute $v=y^{3}$. This could have been observed from the start. The equation becomes $v^{\prime}+v=e^{-x}$ or $\left(e^{x} v(x)\right)^{\prime}=1$. Hence $e^{x} v(x)=x+C$ and then $v(x)=(C+x) e^{-x}$. This gives the general solution $y(x)=(C+x)^{1 / 3} e^{-x / 3}$ for all $x \in \mathbb{R}$.

### 1.6.1 Exact equations

We say the equation

$$
\begin{equation*}
y^{\prime}=f(x, y) \tag{1.33}
\end{equation*}
$$

is an exact equation if $f(x, y)=-\frac{M(x, y)}{N(x, y)}$ and for some function $F(x, y)$ we have $M(x, y)=\frac{\partial F}{\partial x}(x, y)$ and $N(x, y)=\frac{\partial F}{\partial y}(x, y)$. Let us observe that if an equation is
exact then $F(x, y(x))=C$ is a general solution giving $y$ implicitly: indeed, differentiating this with respect to $x$ we get $\frac{\partial F}{\partial x}(x, y)+\frac{\partial F}{\partial y}(x, y) y^{\prime}(x)=0$. This is nothing but the original equation. There is a condition on $M$ and $N$ that tells us if the equation is exact or not.

Theorem 1.6.2. Let $M$ and $N$ as before, defined and continuously differentiable on a rectangle $\mathcal{R}=\{(x, y): a<x<b, c<y<d\}$. Then the equation (1.33) is exact if and only if $\frac{\partial M}{\partial y}(x, y)=\frac{\partial N}{\partial x}(x, y)$ for each $(x, y) \in \mathcal{R}$.

Proof: First suppose that the equation (1.33) is exact. Then there exists $F$ differentiable such that $M(x, y)=\frac{\partial F}{\partial x}(x, y)$ and $N(x, y)=\frac{\partial F}{\partial y}(x, y)$ with $f(x, y)=$ $-\frac{M(x, y)}{N(x, y)}$. Then $\frac{\partial M}{\partial y}(x, y)=\frac{\partial^{2} F}{\partial y \partial x}(x, y)$ and $\frac{\partial N}{\partial x}(x, y)=\frac{\partial^{2} F}{\partial x \partial y}(x, y)$. For a function which twice continuously differentiable the mixed derivatives are equal: $\frac{\partial^{2} F}{\partial y \partial x}(x, y)=\frac{\partial^{2} F}{\partial x \partial y}(x, y)$. This is called Schwartz (or Clairaut)'s theorem.
For the other implication, we define $F(x, y)=\int_{y_{0}}^{y} N\left(x_{0}, s\right) d s+\int_{x_{0}}^{x} M(s, y) d s$. We need to check that $M(x, y)=\frac{\partial F}{\partial x}(x, y)$ and $N(x, y)=\frac{\partial F}{\partial y}(x, y)$. The first equality is a consequence of the Second Fundamental Theorem of Calculus since the first part in the definition of $F$ is constant with respect to $x$ and the second gives $\frac{\partial F}{\partial x}=\frac{d}{d x}\left[\int_{x_{0}}^{x} M(s, y) d s\right]=M(x, y)$. To check the second equality we differentiate with respect to $y$ the definition of $F: \frac{\partial F}{\partial y}=\frac{d}{d y}\left[\int_{y_{0}}^{y} M\left(y_{0}, s\right) d s\right]+\int_{x_{0}}^{x} \frac{\partial M}{\partial y}(s, y) d s$. (We have used differentiation under the sign of integration which is true under our assumptions.) Using the hypothesis that $\frac{\partial M}{\partial y}(x, y)=\frac{\partial N}{\partial x}(x, y)$ we get $\frac{\partial F}{\partial y}=$ $N\left(y_{0}, y\right)+\int_{x_{0}}^{x} \frac{\partial N}{\partial x}(s, y) d s=N\left(y_{0}, y\right)+N(x, y)-N\left(x_{0}, y\right)=N(x, y)$.

Example: Let us look at the problem 36, page 72. In this case $M(x, y)=$ $\left(1+y e^{x y}\right)$ and $N(x, y)=2 y+x e^{x y}$. We need to check if the equation $M(x, y)+$ $N(x, y) y^{\prime}=0$ is exact. Using the Theorem 1.6.2 we see that we have to check that $\frac{\partial M}{\partial y}(x, y)=\frac{\partial N}{\partial x}(x, y)$. So, $\frac{\partial M}{\partial y}(x, y)=(1+x y) e^{x y}$ and $\frac{\partial N}{\partial x}(x, y)=(1+x y) e^{x y}$. In order to solve it we use the same method as in the proof of Theorem 3.1. First we integrate $M(x, y)$ with respect to $x$ and obtain $F(x, y)=x+e^{x y}+h(y)$. Then we differentiate this with respect to $y$ and obtain $N(x, y)=2 y+x e^{x y}=x e^{x y}+h^{\prime}(y)$. Hence $h(y)=2 y$ which implies $h(y)=y^{2}+C$. Then the general solution of this given equation is $x+y^{2}+e^{x y}+C=0$.

### 1.6.2 Reducible second order DE

The DE we are dealing with in these cases is of the form

$$
\begin{equation*}
F\left(x, y, y^{\prime}, y^{\prime \prime}\right)=0 \tag{1.34}
\end{equation*}
$$

In some situations just by making a substitution we can reduce this equation to a first order one. Case I. (Dependent variable y missing) In this case we substitute $v=y^{\prime}$ Then the equation becomes a first-order equation.

Case II. (Independent variable missing ) If the equation is written as $F\left(y, y^{\prime}, y^{\prime \prime}\right)=$ 0 then the substitution $v=y^{\prime}$ will give $y^{\prime \prime}=\frac{d v}{d y} \frac{d y}{d x}=v \frac{d v}{d y}$ and the equation becomes $F\left(y, v, v \frac{d v}{d y}\right)=0$ which is first-order DE.

Let us see an example like this. Problem 54, page 72 asks to solve the DE $y y^{\prime \prime}=3\left(y^{\prime}\right)^{2}$. If we substitute $v=y^{\prime}, y^{\prime \prime}=v \frac{d v}{d y}$ we get $y v \frac{d v}{d y}=3 v^{2}$. One particular solution of this equation is $v(x)=0$ for all $x$. Assuming that $v(x) \neq 0$ for $x$ in some interval $I$ we get $\frac{v^{\prime}}{v}=\frac{3}{y}$. Integrating with respect to $y$ we obtain $\ln |v|=\ln |y|^{3}+C$. From here $v(y)=k y^{3}$ with $k$ and arbitrary real constant. Since $y^{\prime}=k y^{3}$ we can integrate again since this is a separable equation to obtain $-\frac{1}{2 y^{2}}=k x+m$ where $m$ is another constant. This gives the general solution $y(x)=\frac{C_{1}}{\sqrt{1+C_{2} x}}$ with $C_{1}, C_{2} \in \mathbb{R}$.

## Homework:

Section 1.6 pages 71-73: 1-23, 34-36, 43-55, 63 and 66;

## Chapter 2

# Analytical Methods, Second and n-order Linear Differential Equations 

### 2.1 Lecture V

Quotation: "I could never resist a definite integral." G.H. Hardy

### 2.1.1 Stability, Euler's Method, Numerical Methods, Applications

Equilibrium solutions and stability for first-order autonomous DE, critical points, stable and instable critical points, bifurcation point, bifurcation diagram, vertical motion of a body with resistance proportional to velocity, Euler's approximation method, the error theorem in Euler's method

Another way of studying differential equations is to use qualitative methods in which one can say various things about a particular solution of the DE in question without necessarily solving for the solution in closed form or even in implicit form. Even for very simple differential equations which are autonomous first-order:

$$
\begin{equation*}
y^{\prime}=f(y) \tag{2.1}
\end{equation*}
$$

for some continuous function $f$, which leads to a separable DE , the integration $\int \frac{d y}{f(y)}$ may turn out to be very difficult. Not only that but these integrals may be impossible to be expressed in terms of elementary functions that we have reviewed earlier (polynomials, power functions, exponential and logarithmic ones, trigonometric functions and inverses of them).

One important concept for such an analysis in the case of DE of type (2.1) is the following notion:

Definition 2.1.1. A number $c$ for which $f(c)=0$ is called a critical point of the DE (2.1).

If $c$ is a critical point for (2.1), we have a particular solution of (2.1): $y(x)=c$ for all $x \in I$. Such a solution is called an equilibrium solution of (2.1). For the following concept let us assume that $f$ is also continuously differentiable on its domain of definition so that the existence and uniqueness theorem of Cauchy applies.

Definition 2.1.2. A critical point $c$ of (2.1) is said to be stable if for every $\epsilon>0$ there exist a positive number $\delta$ such that if $\left|y_{0}-c\right| \leq \delta$ then the solution of $y(x)$ of the initial value problem associated to (2.1)

$$
\left\{\begin{array}{l}
y^{\prime}=f(y)  \tag{2.2}\\
y(0)=y_{0}
\end{array}\right.
$$

satisfies $|y(x)-c| \leq \epsilon$ for all $x \geq 0$ and in the domain of the solution.
Notice that this is a very technical mathematical definition which is saying that if the initial point where a solution starts is close enough of the critical point $c$ then the whole solution is going to stay close to the corresponding equilibrium solution at any other point in time (black whole behavior). If this definition is not satisfied we say that $c$ is unstable.

Let us work the Problem 22, page 97 as we introduce all these related concepts and techniques.

Problem 2.1.3. Consider the $D E y^{\prime}=y+k y^{3}$ where $k$ is a parameter. Determine the critical points and classify them as stable or unstable.

Solution: The equation $y+k y^{3}=0$ has in general three solutions $y_{1}=0, y_{2,3}=$ $\pm \sqrt{-1 / k}$. There is only one solution if $k=0$, only one real solution if $k>0$ and three real ones if $k<0$.

If $k=0$ there is only one solution $y=0$ and if $y_{0}$ is given the initial value problem (2.2) has unique solution $y(x)=y_{0} e^{x}, x \in \mathbb{R}$, which has the property $\lim _{x \rightarrow \infty}|y(x)|=$ $\infty$, if $y_{0} \neq 0$, which shows that $y=0$ is unstable.

If $k>0$ then $y^{\prime}(x) \geq y(x)$ if $y(x)>0$ at least. Then a similar method to that of solving linear equations shows that $y^{\prime}(x) \geq y_{0} e^{x}$ and this implies the $y=0$ is unstable too. In fact one can integrate (2) and check that this is true. The general solution of

$$
\left\{\begin{array}{l}
y^{\prime}=y+k y^{3}  \tag{2.3}\\
y(0)=y_{0}
\end{array}\right.
$$

is given by $y(x)=\frac{y_{0} e^{x}}{\sqrt{1-k y_{0}^{2}\left(e^{2 x}-1\right)}}$ (please check!). This solution is defined only for $x$ satisfying $\left(e^{2 x}-1\right) k y_{0}^{2}<1$. So if $y_{0} \neq 0$ then $x \in(-\infty, T)$ where $T=\frac{1}{2} \ln \left(1+\frac{1}{k y_{0}^{2}}\right)$. Let us observe that $\lim _{x \rightarrow T}|y(x)|=\infty$ so the point 0 is indeed unstable.

If $k<0$ then the solution is well defined for all values of $x \in[0, \infty)$. We can write the expression of $y(x)$ in the form

$$
y(x)=\frac{y_{0}}{\sqrt{e^{-2 x}\left(1+k y_{0}^{2}\right)-k y_{0}^{2}}}
$$

which at the limit, as $x \rightarrow \infty$, is $\pm \sqrt{\frac{1}{-k}}$ depending upon $y_{0}$ is positive or negative. This shows that $y=0$ is still unstable. On the other hand, one can check that, for instance, $\left|y(t)-\sqrt{\frac{1}{-k}}\right| \leq \epsilon$ if $\left|y_{0}-\sqrt{\frac{1}{-k}}\right| \leq \delta$ where $\delta$ is chosen to be smaller than $\frac{1}{2 a}$ and $\frac{a \epsilon}{14}(a=\sqrt{-k})$. Similarly for the solution $-\sqrt{\frac{1}{-k}}$ which shows that both these critical points are stable. One could come to the same conclusion without solving the system (2.3). For $k<0$ the graph of $f(y)=y+k y^{3}$ as a function of $y$ looks like:


$$
f(y)=y+k y^{3}, k<0
$$

From this graph one can see that if the solution starts close to $y_{2}=\sqrt{\frac{1}{-k}}$ but below $y_{2}$ then the solution is going to have positive derivative. As a result it is going to increase as long as it is less than $y_{2}$. In fact, $y(x)$, is not going to reach the value $y_{2}$ because that is going to contradict the uniqueness theorem. Hence $y(x)$ is going to have a limit, say $L$. One can show that in this case the solution is defined for all $x \in[0, \infty)$, So, we can let $x$ go to infinity in the original DE and obtain that $\lim _{x \rightarrow \infty} y^{\prime}=L+k L^{3}$. One can show that $\lim _{x \rightarrow \infty} y^{\prime}=0$. This implies that $L=y_{2}$.

Remark: The point $k=0$ is called a bifurcation point. By definition, a value of a parameter $k$ is a bifurcation point if the behavior of the critical points (solutions of $f(y, k)=0)$ changes as $k$ increases. The graph of the points $(k, c)$ with $f(c, k)=0$ is called bifurcation diagram.

In some other texts, the a critical point which is stable is also called a sink. If the derivative of $f$ exists at such a point then one checks if $f^{\prime}(c)<0$ and concludes that the critical point is a sink or stable. If $f^{\prime}(c)>0$ then one sees that such a point is not stable or sometime called source. If $f^{\prime}(c)=0$ or $f^{\prime}(c)$ doesn't exist, the critical point is said to be a node.

Some useful ingredients here are:
Problem 2.1.4. Let $g$ be differentiable on $[0, \infty)$ such that $\lim _{x \rightarrow \infty} g(x)$ and $\lim _{x \rightarrow \infty} g^{\prime}(x)$ exist. Show that $\lim _{x \rightarrow \infty} g^{\prime}(x)=0$.

Problem 2.1.5. Suppose that $f$ is some differentiable function on $(a, b)$ with $c \in$ $(a, b)$ a critical point $(f(c)=0)$ such that $f(y)>0$ if $y<c$ and $f(y)<0$ for $y>c$. Show that the initial value problem

$$
\left\{\begin{array}{l}
y^{\prime}=f(y)  \tag{2.4}\\
y(0)=y_{0} \in(a, b)
\end{array}\right.
$$

has a unique solution $y(x)$ defined for all $x \geq 0$ and $\lim _{x \rightarrow \infty} y(x)=c$.
Problem 2.1.6. Let $k>0$ and $f$ be a differentiable function defined on $[0, \infty)$ such that $\lim _{x \rightarrow \infty}\left[f^{\prime}(x)+k f(x)\right]=L$. Show that $\lim _{x \rightarrow \infty} f^{\prime}(x)=0$.

Notice that Problem 0.6 generalizes Problem 0.4.

### 2.1.2 Vertical motion under gravitational force and air resistance proportional to the velocity

A simple application of this analysis can be done for the case of movement of a body with mass $m$ near the surface of the earth subject to gravitation and friction to the air. If one assumes that the friction force is $F=k v$ and opposed to the direction of the movement all the time we get the DE:

$$
m \frac{d v}{d t}=-k v+m g
$$

$$
\begin{equation*}
\frac{d v}{d t}=-\rho v+g . \tag{2.5}
\end{equation*}
$$

One can easily see that $v_{l}=\frac{\rho}{g}=\frac{m g}{k}$ is a stable solution of this differential equation. This speed is called the terminal speed. Please read the analysis done in the book for the case the friction is proportional to the square of the velocity. In this case the terminal speed is $\sqrt{\frac{\rho}{g}}$.

### 2.1.3 Euler's method of approximating the solution of a first-order DE

Algorithm: Given the initial value problem

$$
\left\{\begin{array}{l}
y^{\prime}=f(x, y)  \tag{2.6}\\
y\left(x_{0}\right)=y_{0}
\end{array}\right.
$$

Euler's method with step size $h$ consists in using the recurrent formula $y_{k+1}=y_{k}+$ $h f\left(x_{0}+k h, y_{k}\right)$ for $k=0,1,2, \ldots, n$ in order to compute the approximation $y_{n}$ of the solution of (2.6) at $x=x_{n}$. The difference $\left|y\left(x_{n}\right)-y_{n}\right|$ is called the cumulative error.

The figure below has been obtained with Maple using Euler's method with step size $h=0.001, n=6000$, for the initial value problem

$$
\left\{\begin{array}{l}
y^{\prime}=x^{2}-y^{2}  \tag{2.7}\\
y(-3)=1
\end{array}\right.
$$



The exact solution of this DE is difficult to calculate but there is an expression of it in terms of Bessel functions. Maple 9 doesn't handle it properly so we cannot compute the cumulative error. There is a theorem that gives some information about the cumulative error.

Theorem 2.1.7. Assume that the function $f$ in (2.6) is continuous and differentiable on some rectangle $\mathcal{R}=[a, b] \times[c, d]$. Then there exist a constant $C>0$ (independent of $h$ and as a result, independent of $n$ ) such that $\left|y\left(x_{n}\right)-y_{n}\right|<C h$ as long as $x_{n} \in(a, b)$, where $y_{n}$ is computed with the Euler's method with step size $h$.

This constant $C$ depends only on the function $f$ and on the rectangle $\mathcal{R}$. Theoretically this implies that by taking $h$ small enough we can get any accuracy we want for the solution.

One can obtain better approximations if one uses the improved Euler's approximation method or Runge-Kutta method (please see the book).

Homework: For the first test work problems at the first chapter review on page 76.

Section 2.2 pages 96-97: 1-12, 21, 22;

### 2.2 Lecture VI

Quotation: "If there is a problem you can't solve, then there is an easier problem you can solve: find it." George Pólya
Second-order linear DE, principle of superposition for linear homogeneous equations, existence and uniqueness for linear DE, initial value problem for second-order DE, linear independence of two functions, Wronskian, general solution of linear second-order homogeneous DE, constant coefficients, characteristic equation, the case of real roots, the case of repeated roots and the case of pure complex roots.

The type of equations we are going to be concerned with are DE that could be reduced to

$$
\begin{equation*}
y^{\prime \prime}+p(x) y^{\prime}+q(x) y=f(x) \tag{2.8}
\end{equation*}
$$

for some continuous functions $p, q, f$ on an open interval $I$. Recall that if $f=0$ then we called the DE homogeneous.

An important property for homogenous linear equations is the following:
Theorem 2.2.1. (Superposition principle) If $y_{1}$ and $y_{2}$ are two solutions of $y^{\prime \prime}+p(x) y^{\prime}+q(x) y=0$, then $z(x)=c_{1} y_{1}(x)+c_{2} y_{2}(x)$ is also a solution for every constants $c_{1}$ and $c_{2}$.

Proof. Because $y_{1}^{\prime \prime}+p(x) y_{1}^{\prime}+q(x) y_{1}=0$ and $y_{2}^{\prime \prime}+p(x) y_{2}^{\prime}+q(x) y_{2}=0$, we can multiply the first equation by $c_{1}$ and the second by $c_{2}$ and then add the two new equations together. Then we obtain $z^{\prime \prime}+p(x) z^{\prime}+q(x) z=0$.

The existence and uniqueness theorem takes a special form in this case.
Theorem 2.2.2. (Existence and Uniqueness) For the initial value problem

$$
\left\{\begin{array}{l}
y^{\prime \prime}+p(x) y+q(x) y=f(x)  \tag{2.9}\\
y(a)=b_{1} \\
y^{\prime}(a)=b_{2}
\end{array}\right.
$$

assume that $p, q$ and $f$ are continuous on an interval $I$ containing $a$. Then (2.9) has a unique solution on $I$.

The problem (2.9) is called an initial value problem associated to a secondorder DE.

Example: Suppose we take the differential equation in Problem 16, page 156: $y^{\prime \prime}+$ $\frac{1}{x} y^{\prime}+\frac{1}{x^{2}} y=0$ and the initial condition $y(1)=3$ and $y^{\prime}(1)=2$. Then by applying Theorem 0.2 we know that this initial value problem should have a solution defined on $(0, \infty)$. If we take the the two solutions given in the problem $y_{1}=\cos (\ln x)$ and $y_{2}=\sin (\ln x)$ we can use the Superposition Principle to find our solution by determining the constants $c_{1}$ and $c_{2}$ from the system:

$$
\left\{\begin{array}{l}
c_{1} y(1)+c_{2} y_{2}(1)=3  \tag{2.10}\\
c_{1} y^{\prime}{ }_{1}(1)+c_{2} y^{\prime}{ }_{2}(1)=2
\end{array}\right.
$$

or $c_{1}=3, c_{2}=2$. This gives $y(x)=3 \cos (\ln x)+2 \sin (\ln x)$ which exists on $(0, \infty)$, the largest interval on which $p(x)=\frac{1}{x}$ and $q(x)=\frac{1}{x^{2}}$ are defined and continuous.

Definition 2.2.3. Two functions $f, g$ defined on an interval I are said to be linearly independent on $I$, if $c_{1} f(x)+c_{2} g(x)=0$ for all $x \in I$ implies $c_{1}=c_{2}=0$.

If two functions are not linearly independent on $I$, they are called linearly dependent on $I$.

Example I: Suppose $f(x)=\cos 2 x$ and $g(x)=\cos ^{2} x-\frac{1}{2}$ and $I=\mathbb{R}$. These two functions are linearly dependent on $\mathbb{R}$ since $f(x)+(-2) g(x)=0$ for all $x \in \mathbb{R}$.
Example II: Let us take $f(x)=x$ and $g(x)=|x|$ and $x \in I=\mathbb{R}$. These two are linearly independent since $C_{1} f(x)+C_{2} g(x)=0$ for all $x \in \mathbb{R}$. This implies $C_{1}+C_{2}=0$ if $x=1$ for instance $x=1$ and $-C_{1}+C_{2}=0$ if $x=-1$. This attracts $C_{1}=C_{2}=0$ which means that $f$ and $g$ are linearly independent.

Definition 2.2.4. For two differentiable functions $f$ and $g$ on $I$, the Wronskian of $f$ and $g$ is the determinant

$$
W(f, g)(x)=\left|\begin{array}{cc}
f(x) & g(x) \\
f^{\prime}(x) & g^{\prime}(x)
\end{array}\right|=f(x) g^{\prime}(x)-f^{\prime}(x) g(x), \quad x \in I .
$$

The next theorem characterizes solutions of second-order DE which are linearly independent.

Theorem 2.2.5. Let $y_{1}$ and $y_{2}$ be two solutions of $y^{\prime \prime}+p(x) y^{\prime}+q(x) y=0$ defined on open interval $I$, where $p$ and $q$ are continuous. Then $y_{1}$ and $y_{2}$ are linearly independent if and only if $W\left(y_{1}, y_{2}\right)(x) \neq 0$ for all $x \in I$.

Proof. ( $\Longleftarrow)$ Let us assume that $W\left(y_{1}, y_{2}\right)(x) \neq 0$ for all $x \in I$. By way of contradiction if the solutions $y_{1}$ and $y_{2}$, are linearly dependent then $y_{1}=c y_{2}$ for some constant $c$. Then $W\left(y_{1}, y_{2}\right)=y_{1} y_{2}^{\prime}-y_{1}^{\prime} y_{2}=c y_{2} y_{2}^{\prime}-\left(c y_{2}\right)^{\prime} y_{2}=0$. So, if $W\left(y_{1}, y_{2}\right)(x) \neq 0$ for all $x \in I$. This contradiction shows that the two solutions must be linearly independent.
$(\Longrightarrow)$ Assume that $y_{1}$ and $y_{2}$ are linearly independent. This means that $c_{1} y_{1}+$ $c_{2} y_{2}=0$ on $I$ implies $c_{1}=c_{2}=0$. We will follow the idea from the Problem 32, page 156. Since $y_{1}^{\prime \prime}+p(x) y_{1}^{\prime}+q(x) y_{1}=0$ and $y_{2}^{\prime \prime}+p(x) y_{2}^{\prime}+q(x) y_{2}=0$ we can multiply the first equation by $y_{2}$ and the second by $y_{1}$ and subtract them. We get $y_{1}^{\prime \prime} y_{2}-y_{1} y_{2}^{\prime \prime}+p(x)\left(y_{1} y_{2}^{\prime}-y_{2} y_{1}^{\prime}\right)=0$. In a different notation $W^{\prime}(x)=p(x) W(x)$. This equation in $W$ is linear with the solution $W(x)=W_{0} e^{\int p(x) d x}$. This implies that $W(x) \neq 0$ if $W_{0} \neq 0$. So, we are done if $W_{0} \neq 0$. Again by way of contradiction let us assume that $W_{0}=0$. Then $W(x)=0$ for all $x \in I$. Hence $\left(y_{1} / y_{2}\right)^{\prime}=0$ which means $y_{1} / y_{2}=c$ or $y_{1}(x)-c y_{2}(x)=0$ for some nonzero constant $c$ and for all $x \in I$ where $y_{2}(x) \neq 0$. If $y_{2}(t)=0$, for some $t$, then as a corollary of Theorem 2.9 we cannot have $y_{2}^{\prime}(t)=0$ because that will attract $y_{2} \equiv 0$. Therefore $W(t)=0$ implies $y_{1}(t)=0$ which means $y_{1}(x)-c y_{2}(x)=0$ for al $x \in I$ and this contradicts the assumption on $y_{1}$ and $y_{2}$ as being linear independent. It remains that $W(x) \neq 0$ for all $x \in I$.

The next theorem tells us how the general solution of a homogeneous secondorder linear differential equation looks like.

Theorem 2.2.6. If $y_{1}$ and $y_{2}$ are two linearly independent solutions of $y^{\prime \prime}+p(x) y^{\prime}+$ $q(x) y=0$ defined on open interval $I$, where $p$ and $q$ are continuous then any solution $y$ can be written as $y=c_{1} y_{2}+c_{2} y_{2}$.

Proof. Let us start with an arbitrary solution $y$. Consider an arbitrary point $x_{0} \in I$. From the previous theorem we see that $W\left(x_{0}\right) \neq 0$. Hence by Crammer's rule the system

$$
\left\{\begin{array}{l}
c_{1} y_{1}\left(x_{0}\right)+c_{2} y_{2}\left(x_{0}\right)=y\left(x_{0}\right) \\
c_{1} y_{1}^{\prime}\left(x_{0}\right)+c_{2} y_{2}^{\prime}\left(x_{0}\right)=y^{\prime}\left(x_{0}\right)
\end{array}\right.
$$

has a unique solution in $c_{1}$ and $c_{2}$. Hence $y$ and $z=c_{1} y_{2}+c_{2} y_{2}$ both satisfy the initial value problem

$$
\left\{\begin{array}{l}
w^{\prime \prime}+p(x) w+q(x) w=0  \tag{2.11}\\
w\left(x_{0}\right)=y_{0} \\
w^{\prime}\left(x_{0}\right)=y^{\prime}\left(x_{0}\right)
\end{array}\right.
$$

Using the uniqueness property of the solution (Theorem 2.9) we see that the two solutions must coincide: $y=c_{1} y_{1}+c_{2} y_{2}$.

Two linearly independent solutions of a second-order linear homogeneous DE are called a fundamental set of solutions for this DE.

### 2.2.1 Linear second-order DE with constant coefficients

If the DE is of the form $a y^{\prime \prime}+b y^{\prime}+c y=0$ we can find two solutions which are linearly independent by going first to the characteristic equation:

$$
\begin{equation*}
a r^{2}+b r+c=0 \tag{2.12}
\end{equation*}
$$

Theorem 2.2.7. (a) If the roots of the equation (2.12) are real, say $r_{1}$ and $r_{2}$, and distinct then two linearly independent solutions of $a y^{\prime \prime}+b y^{\prime}+c y=0$ are $e^{r_{1} x}$ and $e^{r_{2} x}$. The general solution of the $D E$ is then given by $y(x)=c_{1} e^{r_{1} x}+c_{2} e^{r_{2} x}$.
(b) If the roots are real but $r_{1}=r_{2}=r$ then two linearly independent solutions of $a y^{\prime \prime}+b y^{\prime}+c y=0$ are $e^{r x}$ and $x e^{r x}$. The general solution of the $D E$ is then given by $y(x)=\left(c_{1}+c_{2} x\right) e^{r_{1} x}$.
(c) If the two roots are pure imaginary ones, say $r_{1,2}=\alpha+i \beta$ then two linearly independent solutions of $a y^{\prime \prime}+b y^{\prime}+c y=0$ are $e^{\alpha x} \sin \beta x$ and $e^{\alpha x} \cos \beta x$. The general solution of the $D E$ is then given by $y(x)=\left(c_{1} \sin \beta x+c_{2} \cos \beta x\right) e^{\alpha x}$.

Proof. We need to check in each case that the given pair of functions form a fundamental set of solutions.

Case (a) The functions $y_{1}(x)=e^{r_{1} x}$ and $y_{2}(x)=e^{r_{2} x}, x \in \mathbb{R}$, satisfy the DE, $a y^{\prime \prime}+b y^{\prime}+c y=0$, because $r_{1}$ and $r_{2}$ satisfy the characteristic equation (2.12). The Wronskian of these two functions is $W\left(y_{1}, y_{2}\right)(x)=\left(r_{2}-r_{1}\right) e^{\left(r_{1}+r_{2}\right) x} \neq 0$ since we assume in this case $r_{1} \neq r_{2}$.
Case (b) The two functions this time are $y_{1}(x)=e^{r_{1} x}$ and $y_{2}(x)=x e^{r_{1} x}$. The only novelty here is why $y_{2}$ must be a solution: $y_{2}^{\prime}(x)=\left(r_{1} x+1\right) e^{r_{1} x}, y_{2}^{\prime \prime}(x)=$ $\left(r_{1}^{2} x+2 r_{1}\right) e^{r_{1} x}$ and $\left.a y_{2}^{\prime \prime}+b y_{2}^{\prime}+c=\left[\left(a r_{1}^{2}+b r_{1}+c\right) x+2 a r_{1}+b\right)\right] e^{r_{1} x} \equiv 0$ since $r_{1}=r_{2}=\frac{-b}{2 a}$. We have $W\left(y_{1}, y_{2}\right)(x)=e^{r_{1} x} \neq 0$ for all $x \in \mathbb{R}$.
Case (c) Here $y_{1}(x)=e^{\alpha x} \sin \beta x$ and $y_{2}(x)=e^{\alpha x} \cos \beta x$. If we calculate $y_{1}^{\prime}(x)=$ $(\alpha \sin \beta x+\beta \cos \beta x) e^{\alpha x}$ and $y_{1}^{\prime \prime}(x)=\left[\left(\alpha^{2}-\beta^{2}\right) \sin \beta x+2 \alpha \beta \cos \beta x\right] e^{\alpha x}$, and $a y_{1}^{\prime \prime}+$ $b y_{1}+c y_{1}=\left[\left(a\left(\alpha^{2}-\beta^{2}\right)+b \alpha+c\right) \sin \beta x+(2 a \alpha \beta+b \beta) \cos \beta x\right] e^{\alpha x}$. But we know from the quadratic formula that $\alpha=\frac{-b}{2 a}$ and $\beta=\frac{\sqrt{4 a c-b^{2}}}{2 a}$. Hence $a\left(\alpha^{2}-\beta^{2}\right)+b \alpha+c=$ 0 and $2 a \alpha+b=0$ which in turn implies $a y_{1}^{\prime \prime}+b y_{1}+c \equiv 0$. Similarly one can check that $a y_{2}^{\prime \prime}+b y_{2}+c y_{2} \equiv 0$. The Wronskian is $W\left(y_{1}, y_{2}\right)(x)=-2 \beta e^{\alpha x} \neq 0$ for all $x \in \mathbb{R}$ (in this case $\beta \neq 0$ ).

Examples: Problem 34, page 156. The DE is $y$ " $+2 y^{\prime}-15 y=0$ whose characteristic equation is $r^{2}+2 r-15=0$. This has two real solutions $r_{1}=-5$ and $r_{2}=3$. Hence the general solution of this equation is $y(x)=c_{1} e^{3 x}+c_{2} e^{-5 x}, x \in \mathbb{R}$.

In Problem 40, page 156 the DE is $9 y^{\prime \prime}-12 y^{\prime}+4 y=0$. The characteristic equation is $9 r^{2}-12 r+4=0$ whose solutions are $r_{1}=r_{2}=\frac{2}{3}$. Hence the general solution of this DE is $y(x)=\left(c_{1} x+c_{2}\right) e^{2 x / 3}, x \in \mathbb{R}$.

If the DE is $y^{\prime \prime}+4 y^{\prime}+13 y=0$ then the characteristic equation $r^{2}+4 r+13=0$ has pure complex roots $r_{1,2}=-2 \pm 3 i$. Therefore the general solution of the given differential equation is

$$
y(x)=\left(c_{1} \sin 3 x+c_{2} \cos 3 x\right) e^{-2 x}
$$

## Homework:

Section 3.1 pages 156-157: 13-16, 24-26, 31-42, 51;

### 2.3 Lecture VII

Quotation: "If a nonnegative quantity was so small that it is smaller
than any given one, then it certainly could not be anything but zero. To those who ask what the infinitely small quantity in mathematics is, we answer that it is actually zero. Hence there are not so many mysteries hidden in this concept as they are usually believed to be. These supposed mysteries have rendered the calculus of the infinitely small quite suspect to many people. Those doubts that remain we shall thoroughly remove in the following pages, where we shall explain this calculus. " Leonhard Euler

Superposition Principle for n-order linear homogeneous DE; Existence and uniqueness for n-order linear DE; Linearly independent and linearly dependent set of functions; Wronskian of a set of $n$, $(n-1)$-times differentiable functions; Characterization theorem of independent solutions; General solutions of an n-order homogeneous linear DE; Complementary solution $y_{c}$ and particular solution $y_{p}$ of an n-order linear DE, Fundamental set of solutions of an n-order homogeneous linear DE; General solutions of an n-order linear DE, n-order linear homogeneous DE with constant coefficients

This lecture is basically a generalization of the previous one. Let us fix $n$ a natural number greater or equal to 2 . We are assuming that the n-order linear DE has been reduced to the form:

$$
\begin{equation*}
y^{(n)}+p_{1}(x) y^{(n-1)}+\ldots+p_{n}(x) y=f(x), \tag{2.13}
\end{equation*}
$$

where $p_{1}, p_{2}, \ldots, p_{n}, f$ are continuous on an open interval $I$. The homogeneous DE associated to (4.5) is

$$
\begin{equation*}
y^{(n)}+p_{1}(x) y^{(n-1)}+\ldots+p_{n}(x) y=0 \tag{2.14}
\end{equation*}
$$

As before an important property for the homogeneous case is the principle of superposition:

Theorem 2.3.1. (Superposition Principle) If $y_{k}, k=1 \ldots n$ are solutions of then (2.14) then the function $z(x)=\sum_{k=1}^{n} c_{k} y_{k}(x), x \in I$ is also a solution of (2.14) for every value of the constants $c_{k}, k=1 \ldots n$.

The proof of this is following exactly the same steps as in the case $n=2$.
The existence and uniqueness theorem needs to formulated in the following way.

Theorem 2.3.2. (Existence and Uniqueness) For the initial value problem

$$
\left\{\begin{array}{l}
y^{(n)}+p_{1}(x) y^{(n-1)}+\ldots+p_{n}(x) y=f(x)  \tag{2.15}\\
y(a)=b_{1}, y^{\prime}(a)=b_{2}, \ldots, y^{(n-1)}(a)=b_{n}
\end{array}\right.
$$

assume that $p_{1}, p_{2}, \ldots, p_{n}, f$ are continuous on an interval $I$ containing $a$. Then (2.15) has a unique solution on I.

The problem (2.15) is called the initial value problem associated to (4.5).
Example: Suppose we take the differential equation in Problem 20, page 168: $x^{3} y^{\prime \prime \prime}+6 x^{2} y^{\prime \prime}+4 x y^{\prime}-4 y=0$ and the initial condition $y(1)=1, y^{\prime}(1)=5$, $y^{\prime \prime}(1)=-11$. We can rewrite the equation as $y^{\prime \prime \prime}+\frac{6}{x} y^{\prime \prime}+\frac{4}{x^{2}} y^{\prime}-\frac{4}{x^{3}} y=0$.

Then by applying Theorem 2.3.2 we know that this initial value problem should have a solution defined on $(0, \infty)$. If we take the three solutions given in the problem's statement: $y_{1}=x, y_{2}=\frac{1}{x^{2}}$ and $y_{3}=\frac{\ln x}{x^{2}}$. We can use the Superposition Principle to find our solution by determining the constants $c_{1}, c_{2}$ and $c_{3}$ from the system:

$$
\left\{\begin{array}{l}
\left.c_{1} y_{( } 1\right)+c_{2} y_{2}(1)+c_{3} y_{3}(1)=1 \\
c_{1} y^{\prime}{ }_{1}(1)+c_{2} y^{\prime}{ }_{2}(1)+c_{3} y^{\prime}{ }_{3}(1)=5 \\
c_{1} y^{\prime \prime}{ }_{1}(1)+c_{2} y^{\prime \prime}{ }_{2}(1)+c_{3} y^{\prime \prime}{ }_{3}(1)=-11
\end{array}\right.
$$

or

$$
\left\{\begin{array}{l}
c_{1}+c_{2} \quad=1 \\
c_{1}-2 c_{2}+c_{3}=5 \\
\quad 6 c_{2}-5 c_{3}=-11
\end{array}\right.
$$

Solving this system of $3 \times 3$ linear equations we get $c_{1}=2, c_{2}=-1$ and $c_{3}=1$. This gives the unique solution of our initial value problem $y(x)=2 x-\frac{1}{x^{2}}+\frac{\ln x}{x^{2}}$ which exists on $I=(0, \infty)$, the largest interval on which $p_{1}, p_{2}, p_{3}$ are defined and continuous and of course containing the initial value for $x(1 \in I)$.

Definition 2.3.3. A set of functions $f_{k}, k=1 \ldots n$, defined on an interval $I$, is said to be linearly independent on $I$, if $\sum_{k=1}^{n} c_{k} f_{k}(x)=0$ for all $x \in I$ implies $c_{1}=c_{2}=$ $\ldots=c_{n}=0$.

If a set of functions is not linearly independent on $I$, the set is called linearly dependent on $I$. By negation of the above definition we see that a set of $n$ functions
is linearly dependent on $I$ if there exist $c_{k}$ not all zero such that $\sum_{k=1}^{n} c_{k} f_{k}(x)=0$ for all $x \in I$.

Example: Suppose $f_{1}(x)=\sin x, f_{2}(x)=\sin 3 x, \ldots, f_{n}(x)=\sin (2 n-1) x$ and $f_{n+1}(x)=\frac{(\sin n x)^{2}}{\sin x}$. These ( $\mathrm{n}+1$ ) functions are linearly dependent on $(0, \pi / 2)$ since $f_{1}+f_{2}+f_{3}+\ldots+f_{n}-f_{n+1} \equiv 0$ (please check!).

Definition 2.3.4. For $n$ functions, $f_{1}, \ldots, f_{n}$, which are $(n-1)$-times differentiable on $I$, the Wronskian of $f_{1}, \ldots, f_{n}$ is the function calculated by the following determinant

$$
W\left(f_{1}, \ldots, f_{n}\right)(x)=\left|\begin{array}{cccc}
f_{1}(x) & f_{2}(x) & \ldots & f_{n}(x) \\
f_{1}^{\prime}(x) & f_{2}^{\prime}(x) & \ldots & f_{n}^{\prime}(x) \\
\ldots & \ldots & \ldots & \ldots \\
f_{1}^{(n-1)}(x) & f_{2}^{(n-1)}(x) & \ldots & f_{n}^{(n-1)}(x)
\end{array}\right|, \quad x \in I .
$$

The next theorem characterizes a set of $n$ solutions of an $n$-order homogeneous linear DE to form a linearly independent set of functions.
Theorem 2.3.5. Let $y_{1}, y_{2}, \ldots, y_{n}$ be $n$-solutions of (2.14). Then $y_{1}, y_{2}, \ldots, y_{n}$ forms a set of linearly independent functions if and only if

$$
W\left(y_{1}, y_{2}, \ldots, y_{n}\right)(x) \neq 0
$$

for all $x \in I$.
Proof. ( $\Longleftarrow)$ For sufficiency let us proceed as before (in the case $n=2$ ) using an argument by contradiction. If the solutions $y_{1}, y_{2}, \ldots, y_{n}$ are linearly dependent then $\sum_{k=1}^{n} c_{k} y_{k} \equiv 0$ for some constants $c_{k}$ not all zero. Then $W\left(y_{1}, y_{2}, \ldots, y_{n}\right) \equiv 0$ because the determinant has one column is a linear combination of the others. So, if $W\left(y_{1}, y_{2}, \ldots, y_{n}\right)(x) \neq 0$ for all $x \in I$ then the set of solutions must be linearly independent.
$(\Longrightarrow)$ For necessity, let us assume the solutions $y_{1}, y_{2}, \ldots, y_{n}$ are linearly independent and again by way of contradiction suppose that their Wronskian is zero for some point $a \in I$. This means that the following homogeneous linear system of equations in $c_{1}, c_{2}, \ldots, c_{n}$

$$
\left\{\begin{array}{l}
c_{1} y_{1}(a)+c_{2} y_{2}(a)+\ldots+c_{n} y_{n}(a)=0  \tag{2.16}\\
c_{1} y_{1}^{\prime}(x)+c_{2} y_{2}^{\prime}(x)+\ldots+c_{n} y_{n}^{\prime}(x)=0 \\
\ldots \\
c_{1} y_{1}^{(n-1)}(a)+c_{2} y_{2}^{(n-1)}(a)+\ldots+c_{n} y_{n}^{(n-1)}(a)=0
\end{array}\right.
$$

has a non-trivial solution. We take such a non-trivial solution, $c_{1}, c_{2}, \ldots, c_{n}$, and consider the function

$$
z(x)=\sum_{k=1}^{n} c_{k} y_{k}(x)
$$

which by the Superposition Principle is a solution of (2.14). Since (2.16) can be written as $z(a)=z^{\prime}(a)=\ldots=z^{(n-1)}(a)=0$ we can apply the uniqueness of solution (Theorem 2.3.2) and conclude that $z \equiv 0$. But that contradicts the fact that $y_{1}, y_{2}, \ldots, y_{n}$ are linearly independent.

Example: The equation $y^{(n)}=0$ has $n$ linearly independent solutions: $y_{1}=1$, $y_{2}(x)=x, y_{3}(x)=x^{2}, \ldots, y_{n}=x^{n-1}$ on any given interval $I$ since the Wronskian of these functions is equal to $1!2!\ldots(n-1)$ ! for all $x \in I$ (please check !).

The next theorem tells us how the general solution of a homogeneous $n$-order linear differential equation looks like.

Theorem 2.3.6. If $y_{1}, y_{2}, \ldots, y_{n}$ are $n$ linearly independent solutions of (2.14) defined on the open interval $I$, then any solution $z$ of (2.14) can be written as $z(x)=$ $\sum_{k=1}^{n} c_{k} y_{k}(x), x \in I$, for some constants $c_{k}, k=1 \ldots n$.

Proof. The proof is the same as in the case $n=2$. Let us start with an arbitrary solution $z$. Consider an arbitrary point $a \in I$. From the previous theorem we see that $W\left(y_{1}, \ldots, y_{n}\right)(a) \neq 0$. Hence by Crammer's rule the system

$$
\left\{\begin{array}{l}
c_{1} y_{1}(a)+c_{2} y_{2}(a)+\ldots+c_{n} y_{n}(a)=z(a)  \tag{2.17}\\
c_{1} y_{1}^{\prime}(x)+c_{2} y_{2}^{\prime}(x)+\ldots+c_{n} y_{n}^{\prime}(x)=z^{\prime}(a) \\
\cdots \\
c_{1} y_{1}^{(n-1)}(a)+c_{2} y_{2}^{(n-1)}(a)+\ldots+c_{n} y_{n}^{(n-1)}(a)=z^{(n-1)}(a),
\end{array}\right.
$$

has a unique solution for $c_{1}, c_{2}, \ldots, c_{n}$. Again if we denote $w(x)=\sum_{k=1}^{n} c_{k} y_{k}(x)$, $x \in I$, then $w$ is a solution of (2.14) and satisfies the initial conditions $w(a)=z(a)$, $w^{\prime}(a)=z^{\prime}(a), \ldots, w^{(n-1)}(a)=z^{(n-1)}(a)$. Again by Theorem 2.3.2 there exist only one such solution. Therefore $z \equiv w$.

A set of $n$ linearly independent solutions of a $n$-order linear homogeneous DE is called a fundamental set of solutions for this DE . So, in solving such a DE we are looking for a fundamental set of solutions. If the differential equation is not homogeneous we have the following characterization of the general solution.

Theorem 2.3.7. If $y_{1}, y_{2}, \ldots, y_{n}$ are $n$ linearly independent solutions of (2.14) defined on the open interval $I$, and $y_{p}$ is a particular solution of (4.5), then any solution $z$
of (4.5) can be written as

$$
z(x)=y_{p}(x)+\sum_{k=1}^{n} c_{k} y_{k}(x), \quad x \in I
$$

for some constants $c_{k}, k=1 \ldots n$.

Proof. If $y_{p}$ is a particular solution of (4.5) then $z-y_{p}$ is a solution of (2.14). Hence by Theorem 2.3.6, $z(x)-y_{p}(x)=\sum_{k=1}^{n} c_{k} y_{k}(x), x \in I$, for some constants $c_{k}, k=1 \ldots n$.

Definition The function $\sum_{k=1}^{n} c_{k} y_{k}(x)$ is called a complementary function associated to (4.5).

### 2.3.1 Linear $n$-order linear DE with constant coefficients

We are going to study the particular situation of (4.5) or (2.14) in which the equation is of the form

$$
\begin{equation*}
a_{0} y^{(n)}+a_{1} y^{(n-1)}+\ldots+a_{n} y=0 . \tag{2.18}
\end{equation*}
$$

where $a_{k}$ are just constant real numbers.
As in the case $n=2$ the discussion here is going to be in terms of the solutions of the characteristic equation:

$$
\begin{equation*}
a_{0} r^{n}+a_{1} r^{n-1}+\ldots+a_{n}=0 \tag{2.19}
\end{equation*}
$$

Theorem 2.3.8. A fundamental set of solutions, $\mathcal{S}$, for (2.18) can be obtained using the following rules
(a) If a root $r$ of the equation (2.19) is real and has multiplicity $k$ then the contribution of this root to $\mathcal{S}$ is with the functions

$$
e^{r x}, x e^{r x}, \ldots, x^{k-1} e^{r x}
$$

(b) If $a$ root $r=a+i b$ of the equation (2.19) is pure complex (i.e. $b \neq 0$ ) and has multiplicity $k$ then the contribution of this root to $\mathcal{S}$ is with the functions

$$
e^{a x} \cos b x, e^{a x} \sin b x, x e^{a x} \cos b x, x e^{a x} \sin b x, \ldots, x^{k-1} e^{a x} \cos b x, x^{k-1} e^{a x} \sin b x
$$

A not very difficult proof of this theorem can be given if one uses an unified approach of the cases (a) and (b) and employing complex-valued functions instead of real-valued ones.

Examples: Problem 12, page 180. The differential equation is $y^{(4)}-3 y^{(3)}+3 y^{\prime \prime}-y^{\prime}=$ 0 . The characteristic equation is $r^{4}-3 r^{3}+3 r^{2}-r=0$. The roots of this equation are $r_{1}=1$ with multiplicity 3 and $r_{2}=0$. Hence the general solution of this equation is $y(x)=\left(c_{1}+c_{2} x+c_{3} x^{2}\right) e^{x}+c_{4}$

Problem 18, page 180. The differential equation is $y^{(4)}=16 y$. The associated characteristic equation is $r^{4}-16=0$. The roots of this equation are $r_{1,2}= \pm 2$ and $r_{3,4}= \pm 2 i$. Therefore the general solution of this DE is $y(x)=c_{1} e^{2 x}+c_{2} e^{-2 x}+$ $c_{3} \cos 2 x+c_{4} \sin 2 x, x \in \mathbb{R}$.

## Homework:

Section 3.2 pages 168-169, 14-20, 27, 28-30, 32-36, 43, 44;
Section 3.3 pages 180-181, 1-20, 24-26, 30-32, 34-36, 45, 46, 50;

### 2.4 Lecture VIII

Quotation: "To divide a cube into two other cubes, a fourth power or in general any power whatever into two powers of the same denomination above the second is impossible, and I have assuredly found an admirable proof of this, but the margin is too narrow to contain it." Pierre Fermat
Topics: Mechanical vibrations (damped, undamped, free, forced, amplitude, circular frequency, phase angle, period, frequency, time lag, critical damping, overdamped, underdamped), nonhomogeneous equations, undetermined coefficients, variation of parameters

Suppose we have a body of mass $m$ attached at one end to an ordinary spring. Hooke's law says that the spring acts on the body with a force proportional to the displacement from the equilibrium position.

Denote by $x$ this displacement. Then this force is $F_{s}=-k x$ where $k$ is called the spring constant. Also let us assume that at the other end the body is attached to a shock absorber that provides a force that is proportional to the speed of the body: $F_{r}=-c \frac{d x}{d t}$. The number $c$ is called the damping constant. If there is also an external force $F_{e}=F(t)$ then according to the Newton's law:

$$
\begin{gather*}
F=F_{s}+F_{r}+F_{e}=m \frac{d^{2} x}{d t^{2}} \\
m \frac{d^{2} x}{d t^{2}}+c \frac{d x}{d t}+k x(t)=F(t) \tag{2.20}
\end{gather*}
$$

Some terminology here has become classic: if we ignore all friction forces, i.e. $c=0$ we say we have an undamped system and it is called damped if $c>0$. If the exterior force is zero we say the system is free and if the exterior forces are present the movement is called forced motion.

### 2.4.1 Free undamped motion

We have basically the equation

$$
\begin{equation*}
m \frac{d^{2} x}{d t^{2}}+k x(t)=0 \tag{2.21}
\end{equation*}
$$

which has the general solution

$$
x(t)=A \cos \omega_{0} t+B \sin \omega_{0} t
$$

where $\omega_{0}=\sqrt{k / m}$ called the circular frequency. This can be written as

$$
\begin{equation*}
x(t)=C \cos \left(\omega_{0} t-\alpha\right), \tag{2.22}
\end{equation*}
$$

where $C=\sqrt{A^{2}+B^{2}}$ is called the amplitude and

$$
\alpha=\left\{\begin{array}{l}
\arctan \left(\frac{B}{A}\right) \quad \text { if } A, B>0,  \tag{2.23}\\
\pi+\arctan \left(\frac{B}{A}\right) \quad \text { if } A<0, \\
2 \pi+\arctan \left(\frac{B}{A}\right) \quad \text { if } A>0 \text { and } B<0, \\
\pi / 2 \quad \text { if } A=0 \text { and } B \geq 0 \\
3 \pi / 2 \quad \text { if } A=0 \text { and } B<0,
\end{array}\right.
$$

is called the phase angle. The period of this simple harmonic motion is simply $T=\frac{2 \pi}{\omega_{0}}$. The physical interpretation is the time necessary to complete one full oscillation. The frequency is defined as the inverse of $T$, i.e. $\nu=\frac{1}{T}$, is usually measured in hetzs (Hz) and measures the number of complete cycles per second. The time lag is the quantity $\delta=\frac{\alpha}{\omega_{0}}$ represents how long it takes to reach the first time the amplitude.

### 2.4.2 Free damped motion

The equation (4.5) becomes:

$$
\begin{equation*}
x^{\prime \prime}+2 p x^{\prime}+\omega_{0}^{2} x=0 \tag{2.24}
\end{equation*}
$$

where $\omega_{0}$ is as before and $p=\frac{c}{2 m}$. The characteristic equation has roots $r_{1,2}=-p \pm$ $\sqrt{p^{2}-\omega_{0}^{2}}$. As we have seen this leads to a discursion in terms of the discriminant of the equation $p^{2}-\omega_{0}^{2}=\frac{c^{2}-4 m k}{4 m^{2}}$. We have a critical damping coefficient for $c_{c r}=$ $2 \sqrt{\mathrm{~km}}$. If $c>c_{c r}$ we say the system is over-damped in which case $x(t) \rightarrow 0$ as $t \rightarrow \infty$ since the general solution is

$$
x(t)=A e^{r_{1} t}+B e^{r_{2} t}, t \in \mathbb{R} .
$$

There are no oscillations around the equilibrium position and the body passes through the equilibrium position at most once.

If $c=c_{c r}$, the system is critically-damped and the general solution is of the form

$$
x(t)=(A+B t) e^{r t}, t \in \mathbb{R}
$$

and as before $x(t) \rightarrow 0$ as $t \rightarrow \infty$ and again the body passes through the equilibrium position at most once.

If $c<c_{c r}$ we say the system is under-damped. The general solution in this case is

$$
x(t)=e^{-p t}\left(A \cos \omega_{1} t+B \sin \omega_{1} t\right)=e^{-p t} \cos \left(\omega_{1} t-\alpha\right),
$$

using the same notations as before. In this case, $\omega_{1}=\frac{\sqrt{4 m k-c^{2}}}{2 m}$ is called circular pseudo-frequency, and $T_{1}=\frac{2 \pi}{\omega_{1}}$ is its pseudo-period.

### 2.4.3 Nonhomogeneous linear equations, undetermined coefficients method

In order to determine a particular solution of a nonhomogeneous linear equation of the form

$$
\begin{equation*}
a_{0} y^{(n)}+a_{1} y^{(n-1)}+\ldots+a_{n} y=f(x), \tag{2.25}
\end{equation*}
$$

in special situations one can try a solution of a certain form. This method applies whenever the function $f$ is a finite linear combination of products of polynomials, exponentials, cosines or sines. One needs to apply two rules.

Rule 1: If none of the terms of the function $f$ contains solutions of the homogeneous DE associated to (2.25), then the reccomended function to try is $y_{p}$ a combination of the terms in $f$ and all their derivatives that form a finite set of linearly independent functions.

Example 1: Let us suppose the DE is $y^{(3)}+y=\sin x+e^{x}$. Since the complementary solution of this equation is $y_{c}=a_{1} e^{-x}+\left[a_{2} \cos (x \sqrt{3} / 2)+a_{3} \sin (x \sqrt{3} / 2)\right] e^{x / 2}$ we can try a particular solution to be $y=c_{1} \sin x+c_{2} \cos x+c_{3} e^{x}$. After substituting in the equation we get $\left(c_{2}-c_{1}\right) \cos x+\left(c_{2}+c_{1}\right) \sin x+2 c_{3} e^{x}=\sin x+e^{x}$. So, it follows that $c_{1}=c_{2}=c_{3}=1 / 2$. So, the general solution of the given equation is $y(x)=a_{1} e^{-x}+\left[a_{2} \cos (x \sqrt{3} / 2)+a_{3} \sin (x \sqrt{3} / 2)\right] e^{x / 2}+\frac{1}{2}\left(\sin x+\cos x+e^{x}\right)$

Example 2: Suppose the DE is $y^{\prime \prime}+2 y^{\prime}-3 y=x^{2} e^{2 x}$. Because the complementary solution of this equation is $y_{c}=a_{1} e^{x}+a_{2} e^{-3 x}$ we can take as a particular solution $y_{p}=\left(c_{1} x^{2}+c_{2} x+c_{3}\right) e^{2 x}$. Since $y_{p}^{\prime}=\left[2 c_{1} x^{2}+\left(2 c_{1}+2 c_{2}\right) x+c_{2}+2 c_{3}\right] e^{2 x}$ and $y_{p}^{\prime \prime}=\left[4 c_{1}+\left(8 c-1+4 c_{2}\right) x+2 c_{1}+4 c_{2}+4 c_{3}\right) e^{2 x}$, we see that $y^{\prime \prime}+2 y^{\prime}-3 y=$ $\left.\left[5 c_{1} x^{2}+\left(12 c_{1}+5 c_{2}\right) x+2 c_{1}+6 c_{2}+5 c_{3}\right)\right] e^{2 x}=x^{2} e^{2 x}$. Therefore $c_{1}=\frac{1}{5}, c_{2}=\frac{-12}{25}$ and $c_{3}=\frac{62}{125}$. Thus the general solution of the given equation is $y(x)=a_{1} e^{x}+a_{2} e^{-3 x}+$ $\frac{25 x^{2}-60 x+62}{125} e^{2 x}$.
Rule 2: If the function $f$ contains terms which are solutions of the homogeneous linear DE associated, then one should try as a particular solution, $y_{p}$, a linear combination of these terms and their derivatives which are linearly independent multiplied by a power of $x$, say $x^{s}$, where $s$ is the smallest nonnegative integer which makes all the new terms not to be solutions of the homogeneous problem.
Example 3: Let us assume the differential equation we want to solve is $y^{\prime \prime}+2 y^{\prime}+y=$ $x^{2} e^{-x}$. So we need to determine the coefficients of the particular solution $y_{p}=$ $\left(c_{1} x^{4}+c_{2} x^{3}+c_{3} x^{2}\right) e^{-x}$. After a simple calculation we get $c_{1}=\frac{1}{12}$ and $c_{2}=c_{3}=0$.
Example 4: Suppose we are given the DE $y^{(4)}-4 y^{(3)}+6 y^{\prime \prime}-4 y^{\prime}+y=e^{x} \sin x$. The complementary solution is $y_{c}=\left(a_{1}+a_{2} x+a_{3} x^{2}+a_{4} x^{3}\right) e^{x}$ we need to look for a particular solution of the form $y_{p}=\left(c_{1} \sin x+c_{2} \cos x\right) e^{x}$. If we introduce the differential operator $D=\frac{d}{d x}$ then the equation given is equivalent to $(D-1)^{4} y=$ $e^{x} \sin x$ and we are looking for a particular solution $y_{p}=u(x) e^{x}$. Since $(D-1) y_{p}=$ $(D u) e^{x}$ (please check!) we see that $(D-1)^{4} y_{p}=\left(D^{4} u\right) e^{x}=\left(c_{1} \sin x+c_{2} \cos x\right) e^{x}$ so $c_{1}=1$ and $c_{2}=0$.

### 2.4.4 Nonhomogeneous linear equations, Variation of parameters method

We are going to describe the method in the case $n=2$ but this works in fact for the $n$-order linear nonhomogeneous DE.

Theorem 2.4.1. A particular solution of the differential equation $y^{\prime \prime}+p(x) y^{\prime}+$ $q(x) y=f(x)$ is given by

$$
\begin{equation*}
y_{p}=-y_{1}(x) \int \frac{y_{2}(x) f(x)}{W\left(y_{1}, y_{2}\right)(x)} d x+y_{2}(x) \int \frac{y_{1}(x) f(x)}{W\left(y_{1}, y_{2}\right)(x)} d x \tag{2.26}
\end{equation*}
$$

where $y_{1}$ and $y_{2}$ is a fundamental set of solutions of $y^{\prime \prime}+p(x) y^{\prime}+q(x) y=0$.
Proof. We are looking for a solution of the form $y_{p}=u_{1}(x) y_{1}(x)+u_{2}(x) y_{2}(x)$. Differentiating with respect to $x$ we obtain $y_{p}=u_{1}^{\prime} y_{1}+u_{2}^{\prime} y_{2}+u_{1} y_{1}^{\prime}+u_{2} y_{2}^{\prime}$. We are going to make an assumption that is going to simplify the next differentiation:

$$
\begin{equation*}
u_{1}^{\prime} y_{1}+u_{2}^{\prime} y_{2}=0 \tag{2.27}
\end{equation*}
$$

Hence, $y_{p}^{\prime \prime}=u_{1}^{\prime} y_{1}^{\prime}+u_{2}^{\prime} y_{2}^{\prime}+u_{1} y_{1}^{\prime \prime}+u_{2} y_{2}^{\prime \prime}$ and then $y_{p}^{\prime \prime}+p(x) y_{p}^{\prime}+q(x) y_{p}=u_{1}\left(y_{1}^{\prime \prime}+\right.$ $\left.p(x) y_{1}^{\prime}+q(x) y_{1}\right)+u_{2}\left(y_{2}^{\prime \prime}+p(x) y_{2}^{\prime}+q(x) y_{1}\right)+u_{1}^{\prime} y_{1}^{\prime}+u_{2}^{\prime} y_{2}^{\prime}=f(x)$. Thus this reduces to

$$
\begin{equation*}
u_{1}^{\prime} y_{1}^{\prime}+u_{2}^{\prime} y_{2}^{\prime}=f(x) . \tag{2.28}
\end{equation*}
$$

Using (2.27) and (2.28) to solve for $u_{1}^{\prime}$ and $u_{2}^{\prime}$ that gives $u_{1}^{\prime}(x)=-\frac{y_{2}(x) f(x)}{W\left(y_{1}, y_{2}\right)(x)}$ and $u_{2}^{\prime}(x)=\frac{y_{1}(x) f(x)}{W\left(y_{1}, y_{2}\right)(x)}$ which gives (2.26).
Example: Let us work Problem 58, page 208. The DE is $x^{2} y^{\prime \prime}-4 x y^{\prime}+6 y=x^{3}$. Dividing the equation by $x^{2}$ we get $y^{\prime \prime}-4 y^{\prime} / x+6 y / x^{2}=x$. So, we have $p(x)=-4 / x$ and $q(x)=6 / x^{2}$ and $f(x)=x$. Two linearly independent solutions are given: $y_{1}=$ $x^{2}$ and $y_{2}=x^{3}$. We have $W\left(y_{1}, y_{2}\right)(x)=3 x^{4}-2 x^{4}=x^{4} \neq 0$ for $x \in(0, \infty)$. Then $u_{1}^{\prime}(x)=-\frac{y_{2}(x) f(x)}{W\left(y_{1}, y_{2}\right)(x)}=-1$ which gives $u_{1}(x)=-x$ and $u_{2}^{\prime}(x)=\frac{y_{1}(x) f(x)}{W\left(y_{1}, y_{2}\right)(x)}=1 / x$ which implies $u_{2}(x)=\ln x$. Therefore a particular solution of this equation is $y_{p}=x^{3}(\ln x-1)$ with $x \in(0, \infty)$.

## Homework:

Section 3.4 pages 192-193, 1-4, 13, 15, 16, 32, 33;
Section 3.5 pages 207-208, 1-20, 31-40, 47-56, 58-63.

### 2.5 Lecture IX

Quotation: "Finally, two days ago, I succeeded - not on account of my hard efforts, but by the grace of the Lord. Like a sudden flash of lightning, the riddle was solved. I am unable to say what was the conducting thread that connected what I previously knew with what made my success possible." Carl Friedrich Gauss

Topics: Forced Oscillations, Beats, Resonance, Boundary Value Problems

### 2.5.1 Undamped Forced Oscillations

In the previous lecture we studied the mechanical vibrations of a body under the action of a spring, damped forces and exterior forces. The DE was:

$$
\begin{equation*}
m \frac{d^{2} x}{d t^{2}}+c \frac{d x}{d t}+k x(t)=F(t) \tag{2.29}
\end{equation*}
$$

Now, we assume the exterior force $F(t)$ is of the form $F(t)=F_{0} \cos \omega t$ and the damping coefficient $c=0$. The differential equation that we need to study is of the form

$$
\begin{equation*}
m x^{\prime \prime}+k x=F_{0} \cos \omega t \tag{2.30}
\end{equation*}
$$

which admits as a complementary solution $x_{c}(t)=c_{1} \cos \omega_{0} t+c_{2} \sin \omega_{0} t$, where $\omega_{0}=\sqrt{\frac{k}{m}}$. First let us assume that $\omega \neq \omega_{0}$. Then, to find a particular solution of (4.5) we try $x_{p}(t)=A \cos \omega t$ using the undetermined coefficient method (no term in $\sin \omega t$ is needed as we can see from the following computation):

$$
-A m \omega^{2} \cos \omega t+A k \cos \omega t=F_{0} \cos \omega t
$$

which implies $A=\frac{F_{0}}{k-m \omega^{2}}=\frac{F_{0} / m}{\omega_{0}^{2}-\omega^{2}}$. Therefore, the general solution of (4.5) is

$$
x(t)=c_{1} \cos \omega_{0} t+c_{2} \sin \omega_{0} t+\frac{F_{0} / m}{\omega_{0}^{2}-\omega^{2}} \cos \omega t
$$

This shows that the solution is a combination of two harmonic oscillations having different frequencies:

$$
\begin{equation*}
x(t)=C \cos \left(\omega_{0} t-\alpha\right)+\frac{F_{0} / m}{\omega_{0}^{2}-\omega^{2}} \cos \omega t . \tag{2.31}
\end{equation*}
$$

where $C=\sqrt{c_{1}^{2}+c_{2}^{2}}$ and $\alpha$ is defined by (2.23) but some notation has changed so we update it here:

$$
\alpha=\left\{\begin{array}{l}
\arctan \left(\frac{c_{2}}{c_{1}}\right) \quad \text { if } \quad c_{1}, c_{2} \geq 0, \\
\pi+\arctan \left(\frac{c_{2}}{c_{1}}\right) \quad \text { if } \quad c_{1}<0, \\
2 \pi+\arctan \left(\frac{c_{2}}{c_{1}}\right) \quad \text { if } \quad c_{1}>0 \\
3 \pi / 2 \quad \text { if } c_{1}=0 \quad \text { and } \quad c_{2}<0 .
\end{array}\right.
$$

### 2.5.2 Beats

If the amplitude $C=\sqrt{c_{1}^{2}+c_{2}^{2}}=\frac{F_{0} / m}{\left|\omega_{0}^{2}-\omega^{2}\right|}$ and the phase $\alpha$ is zero if $\omega>\omega_{0}$ or $\pi$ if $\omega<\omega_{0}$, which can be accomplished by imposing the initial condition $x(0)=x^{\prime}(0)=$ 0 , then the general solution can be written as

$$
x(t)=\frac{F_{0} / m}{\omega_{0}^{2}-\omega^{2}}\left(\cos \omega t-\cos \omega_{0} t\right)
$$

or

$$
x(t)=\frac{2 F_{0} / m}{\omega_{0}^{2}-\omega^{2}} \sin \frac{\left(\omega_{0}-\omega\right) t}{2} \sin \frac{\left(\omega_{0}+\omega\right) t}{2}
$$

If we assume that the two frequencies are close to one another (i.e. $\omega \approx \omega_{0}$ ) the expression above explains the behavior of the solution in some sense. We have a product of two harmonic functions, one with a big circular frequency, $\left(\omega_{0}+\omega\right) / 2$, and the other with a smaller one $\left|\omega_{0}-\omega\right| / 2$ which gives the phenomenon of beats. The graph below is the graph of a function of this type: $f(t)=\sin (t) \sin (30 t)$ on the interval $t \in[-2 \pi, 4 \pi]$.


### 2.5.3 Resonance

Suppose that we have $\omega$ getting closer and closer (within any given $\epsilon>0$ ) to $\omega_{0}$. Then $A(t)=\frac{F_{0} / m}{\omega_{0}^{2}-\omega^{2}}$ goes to infinity. This is the phenomenon of resonance. In fact a particular solution of the problem (4.5) in the case $\omega=\omega_{0}$ is $x_{p}(t)=t \sin \omega_{0} t$. The graph of the of a function of this type is included below:


This phenomenon is considered to be the explanation of a lot of disasters like the one that happened in 1940 with the Tacoma Narrows Bridge near Seattle. It seeamed like the exterior forces created by the wind created exactly this kind of explosion of the amplitude of the oscillations in the vertical suspension cables. Another classical example is the collapse of the Broughton Bridge near Manchester in England of 1831 when soldiers marched upon it.

### 2.5.4 Endpoint problems and eigenvalues

We are concerned with second order linear and homogenous DE which have a special type of initial conditions. One such endpoint problem is:

$$
\left\{\begin{array}{l}
y^{\prime \prime}+p(x) y^{\prime}+\lambda q(x) y=0  \tag{2.32}\\
a_{1} y(a)+a_{2} y^{\prime}(a)=0 \\
b_{1} y(b)+b_{2} y^{\prime}(b)=0
\end{array}\right.
$$

where $a \neq b$. In general only the trivial solution $y \equiv 0$ satisfies (2.32). But for some values of the parameter $\lambda$ the problem (2.32) may have non-zero solutions. These values are called eigenvalues and the corresponding functions are called eigenfunctions. A general method to solve (2.32) is to write the general solution of the DE as $y=A y_{1}(x, \lambda)+B y_{2}(x, \lambda)$, where $y_{1}$ and $y_{2}$ is a fundamental set of solutions which is also going to depend of $\lambda$. We impose the two initial boundary
conditions and rewrite these equations as a system in $A$ and $B$ :

$$
\left\{\begin{array}{l}
\alpha_{1}(\lambda) A+\beta_{1}(\lambda) B=0  \tag{2.33}\\
\alpha_{2}(\lambda) A+\beta_{2}(\lambda) B=0
\end{array}\right.
$$

This system in $A$ and $B$ has a non-trivial solution if and only if

$$
\begin{equation*}
\alpha_{1}(\lambda) \beta_{2}(\lambda)-\alpha_{2}(\lambda) \beta_{1}(\lambda)=0 . \tag{2.34}
\end{equation*}
$$

One solves this equation and obtains the eigenvalues of (2.32).
Example: Let us work Problem 6, page 240. The DE is $y^{\prime \prime}+\lambda y=0$ and the boundary conditions are $y^{\prime}(0)=0$ and $y(1)+y^{\prime}(1)=0$. We are given that all eigenvalues are nonnegative, so we write $\lambda=\alpha^{2}$.
(a) We have to show that $\lambda=0$ is not an eigenvalue. If by way of contradiction we assume it is, then some non-zero solution of our problem must exist: $y(x)=$ $A+B x, 0=y^{\prime}(0)=B$ and then $0=y(1)+y^{\prime}(1)=A$ which is a contradiction.
(b) We need to show that the eigenvalues of this problem are the solutions in $\lambda$ of the equation

$$
\tan \sqrt{\lambda}=\frac{1}{\sqrt{\lambda}}
$$

Let $y(x)=A \cos \alpha t+B \sin \alpha t$ be the general solution of our DE without the boundary conditions. Since $y^{\prime}(0)=0$ we have $0=(-\alpha A \sin \alpha t+B \alpha \cos \alpha t)_{\mid t=0}$ or $B=0$. Then $y(1)+y^{\prime}(1)=0$ implies $A \cos \alpha-A \alpha \sin \alpha=0$. Since we assume there exit a non-zero solution, we must have $A \neq 0$. Therefore $\alpha$ must satisfy $\cos \alpha-\alpha \sin \alpha=0$ or $\tan \alpha=\frac{1}{\alpha}$. Corresponding eigenfunctions are $y(x)=A \cos \alpha x$. Since a picture is worth a thousand words let us include the graph of $\lambda \rightarrow \tan \sqrt{\lambda}$ and $\lambda \rightarrow \frac{1}{\sqrt{\lambda}}$ for $\lambda>0$.


## Homework:

Section 3.6 page 219 Problems 21, 22;
Section 3.8 page 240 Problems 1-6, 13, 14.

## Chapter 3

## Systems of Differential Equations

### 3.1 Lecture X

Quotation: "We [he and Halmos] share a philosophy about linear algebra: we think basis-free, we write basis-free, but when the chips are down we close the office door and compute with matrices like fury. Paul Halmos: Celebrating 50 Years of Mathematics." Irving Kaplansky

### 3.1.1 A non-linear classical example: Kepler's laws of planetary motion

After analyzing observations of Tycho Brache, Johannes Kepler arrived to the following laws of planetary motion:

1. The orbits of planets are ellipses (with the sun in one of the foci).
2. The planets move in such a way on the orbit, that their corresponding ray wipes out an area that varies at a constant rate.
3. The square of the planet's period of revolution is proportional to the cube of the major semi-axis of the elliptical orbit.

We are going to make an assumption here which is not very far from what it happens in the reality (neglect the influence of the planet in question on the sun). The sun contains more than $99 \%$ of the mass in the solar system, so the influence of the planets on the sun could be, on a first analysis, neglected. Intuitively it is not hard to believe that the planet $X$ is moving in a fixed plane although this is also a consequence of the movement under the gravitational field. Let us take the origin of the coordinates in this plane centered at the sun.

The position vector corresponding to the planet $X$ is denoted here by $\vec{r}=$ $x(t) \vec{i}+y(t) \vec{j}$ where $\vec{i}=(1,0)$ and $\vec{j}=(0,1)$. The distance between the sun and the planet $X$ is $r=\sqrt{x(t)^{2}+y(t)^{2}}$. According to Newton's law the planet $X$ moves under the action of a force which is inverse proportional to the square of the distance $r$. The law can be written as a differential equation in the following way:

$$
\begin{equation*}
\vec{r}^{\prime \prime}=-k \frac{\vec{r}}{r^{3}} \tag{3.1}
\end{equation*}
$$

We are going to derive Kepler's laws from (3.1). First let us observe that (3.1) is just the vectorial form of the following second-order non-linear autonomous system of differential equations:

$$
\left\{\begin{array}{l}
x^{\prime \prime}=-k \frac{x}{\left(x^{2}+y^{2}\right)^{3 / 2}}  \tag{3.2}\\
y^{\prime \prime}=-k \frac{y^{\prime}}{\left(x^{2}+y^{2}\right)^{3 / 2}}
\end{array}\right.
$$

It is really a significant fact that this can be reduced to a differential equation that we know how to solve. To see this, let us first introduce polar coordinates, by assuming that the trajectory is written in polar coordinates, $r=r(\theta)$, and we consider two unit vectors that will help us simplify the calculations:

$$
\begin{aligned}
& \vec{u}=\cos \theta \vec{i}+\sin \theta \vec{j} \\
& \text { and } \\
& \vec{v}=-\sin \theta \vec{i}+\cos \theta \vec{j} .
\end{aligned}
$$

It is easy to check that $\vec{u} \cdot \vec{v}=0$, and these two vectors clearly depend of time because $\theta$ is. Differentiating these two vectors with respect to time we get

$$
\begin{gather*}
\frac{d \vec{u}}{d t}=(-\sin \theta \vec{i}+\cos \theta \vec{j}) \frac{d \theta}{d t}=\vec{v} \frac{d \theta}{d t} \\
\text { and }  \tag{3.3}\\
\frac{d \vec{v}}{d t}=(-\cos \theta \vec{i}-\sin \theta \vec{j}) \frac{d \theta}{d t}=-\vec{u} \frac{d \theta}{d t} .
\end{gather*}
$$

Since $\vec{r}=r \vec{u}$ after differentiating this equality, we obtain

$$
\frac{d \vec{r}}{d t}=r^{\prime} \vec{u}+r \frac{d \vec{u}}{d t}=r^{\prime} \vec{u}+r \vec{v} \frac{d \theta}{d t}
$$

Differentiating one more time and using (3.3) we get:

$$
\frac{d^{2} \vec{r}}{d t^{2}}=r^{\prime \prime} \vec{u}+2 r^{\prime} \vec{v} \frac{d \theta}{d t}-r \vec{u}\left(\frac{d \theta}{d t}\right)^{2}+r \vec{v} \frac{d^{2} \theta}{d t^{2}}
$$

or

$$
\frac{d^{2} \vec{r}}{d t^{2}}=\left(r^{\prime \prime}-r\left(\frac{d \theta}{d t}\right)^{2}\right) \vec{u}+\left(2 r^{\prime} \frac{d \theta}{d t}+r \frac{d^{2} \theta}{d t^{2}}\right) \vec{v}=-k \frac{\vec{u}}{r^{2}}
$$

Identifying the coefficients of $\vec{u}$ and $\vec{v}$ in the above relation we obtain

$$
\left\{\begin{array}{l}
r^{\prime \prime}-r\left(\frac{d \theta}{d t}\right)^{2}=-\frac{k}{r^{2}}  \tag{3.4}\\
2 r^{\prime} \frac{d \theta}{d t}+r \frac{d^{2} \theta}{d t^{2}}=0 .
\end{array}\right.
$$

The second relation in (3.4) is equivalent to $\frac{d}{d t}\left(r^{2} \frac{d \theta}{d t}\right)=0(r \neq 0)$. This means $r^{2} \frac{d \theta}{d t}=h$ for some constant $h$. This proves the second Kepler's law since

$$
\frac{d A}{d t}=\frac{d A}{d \theta} \frac{d \theta}{d t}=\left[\lim _{\Delta \theta \rightarrow 0} \frac{r(\theta+\Delta \theta) r(\theta) \sin (\Delta \theta)}{2 \Delta \theta}\right] \frac{d \theta}{d t}=\frac{1}{2} r^{2} \frac{d \theta}{d t}=\frac{h}{2} .
$$

The first equation in (3.4) can be transformed using the substitution $r=\frac{1}{z}$ and changing the independent variable to $\theta$ instead of $t$ : $\frac{d r}{d t}=-\frac{1}{z^{2}} \frac{d z}{d \theta} \frac{d \theta}{d t}=-h \frac{d z}{d \theta}$ and then $\frac{d^{2} r}{d t^{2}}=-h \frac{d^{2} z}{d \theta^{2}} \frac{d \theta}{d t}=-h^{2} z^{2} \frac{d^{2} z}{d \theta^{2}}$ which gives

$$
-h^{2} z^{2} \frac{d^{2} z}{d \theta^{2}}-\frac{1}{z} h^{2} z^{4}=-k z^{2}
$$

Equivalently, this can be written as

$$
\begin{equation*}
\frac{d^{2} z}{d \theta^{2}}+z=\frac{k}{h^{2}} . \tag{3.5}
\end{equation*}
$$

As we have seen the general solution of this is $z=A \cos \theta+B \sin \theta+\frac{k}{h^{2}}=$ $\frac{k}{h^{2}}(1+e \cos (\theta-\alpha))$ where $e=\frac{h^{2}}{k} \sqrt{A^{2}+B^{2}}, \cos \alpha=\frac{A}{\sqrt{A^{2}+B^{2}}}$ and $\sin \alpha=\frac{B}{\sqrt{A^{2}+B^{2}}}$. This gives

$$
\begin{equation*}
r=\frac{L}{1+e \cos (\theta-\alpha)} \tag{3.6}
\end{equation*}
$$

which represents an ellipse if the eccentricity $e$ satisfies $0 \leq e<1$, a parabola if $e=1$ or a hyperbola if $e>1$. Since the orbits of the planets are bounded it must be the case that $e<1$. Comets, by definition, are having parabolic or hyperbolic orbits. [So, according to this definition, Halley's comet is actually not a comet.] This proves the first Kepler's law.

To derive the third Kepler's law, let us integrate the area formula $\frac{d A}{d t}=h / 2$ over the interval $[0, T]$, where $T$ is the period of the orbit. Then we get $h T / 2=$ Area(Ellipse), but the area of an ellipse is equal to $\pi a b$, where $a$ and $b$ are the two semiaxes. The big axis is $a=\left(\frac{L}{1+e}+\frac{L}{1-e}\right) / 2=\frac{L}{1-e^{2}}$ and $b=\frac{L}{\sqrt{1-e^{2}}}$. This means that $\frac{h^{2} T^{2}}{4}=\frac{\pi^{2} L^{4}}{\left(1-e^{2}\right)^{3}}$. From here we see that

$$
T^{2}=\frac{4 \pi^{2} L}{h^{2}} a^{3}=\frac{4 \pi^{2}}{k} a^{3}
$$

or

$$
\frac{T^{2}}{a^{3}}=\frac{4 \pi}{k}=\text { constant }
$$

which is the third Kepler's law.

### 3.1.2 Linear systems of differential equations

Our general setting here is going to be

$$
\begin{equation*}
x^{\prime}(t)=P(t) x(t)+f(t) \tag{3.7}
\end{equation*}
$$

where $P(t)$ is a $n \times n$ matrix whose coefficients are continuous functions on an interval $I, x(t)=\left[x_{1}(t), x_{2}(t), \ldots, x_{n}(t)\right]^{t}$ is the column vector of the unknown functions and $f(t)=\left[f_{1}(t), f_{2}(t), \ldots, f_{n}(t)\right]^{t}$ is a vector-valued function assumed continuous on $I$ as well.

Theorem 3.1.1. (Existence and uniqueness) The initial value problem

$$
\left\{\begin{array}{l}
x^{\prime}(t)=P(t) x(t)+f(t)  \tag{3.8}\\
x_{1}(a)=b_{1}, x_{2}(a)=b_{2}, \ldots, x_{n}(a)=b_{n}
\end{array}\right.
$$

with $a \in I$ has a unique solution defined on $I$.

We notice here that the number of initial conditions in (4.2) is equal to the number of unknowns.

Theorem 3.1.2. (Principle of superposition) If $y_{1}, y_{2}, \ldots, y_{n}$ are solutions of the homogeneous problem associated to (3.7), i.e. $x^{\prime}(t)=P(t) x(t)$, then the vector function $y=\sum_{k=1}^{n} c_{k} y_{k}$ is also a solution for every values of the constants $c_{k}$.

Definition 3.1.3. As before, we say that a set of vector-valued functions, $\left\{f_{1}, f_{2}, \ldots, f_{k}\right\}$, is called linearly independent if $\sum_{i=1}^{k} c_{i} f_{k}(t)=0$ for all $t \in I$, implies $c_{i}=0$ for all $i=1 . . . k$.

For a set of $n$ vector valued functions, $y_{1}, \ldots, y_{n}$, the Wronskian in this case, $W\left(y_{1}, \ldots, y_{n}\right)$, is constructed as the determinant of the matrix $\left(y_{j k}\right)_{j, k=1 . . n}$ where $y_{j}=\left[y_{1 j}, y_{2 j}, \ldots, y_{n j}\right]^{t}$. We have a similar characterization of linear independence.

Theorem 3.1.4. If the system $x^{\prime}(t)=P(t) x(t)$ admits $y_{1}, \ldots, y_{n}$ as solutions then these are linearly independent if and only if the Wronskian asssociated to them, $W\left(y_{1}, \ldots, y_{n}\right)(t)$, is not zero for all $t \in I$.

The proof of this theorem goes the same way as the one for the similar theorem we studied in the case of n-order linear differential equations. Its proof is based on the existence and uniqueness theorem. Based on this, let us remark that every homogeneous system $x^{\prime}=P x$ admits a fundamental set of solutions, i.e. a set of $n$ linearly independent solutions.

Indeed, one has to use the existence and uniqueness theorem and denote by $y_{k}$, the solution of the initial value problem

$$
\left\{\begin{array}{l}
x^{\prime}(t)=P(t) x(t) \\
x_{1}(a)=0, \ldots, x_{k}(a)=1, \ldots x_{n}(a)=0 \quad(a \in I)
\end{array}\right.
$$

Then the Wronskian of the solutions $\left\{y_{1}, y_{2}, \ldots, y_{n}\right\}$ at $a$ is equal to the determinant of the identity matrix, which is, in particular, not zero. By Theorem 3.1.4 we then see that this set of solutions must be linearly independent.

Theorem 3.1.5. Suppose $\left\{y_{1}, y_{2}, \ldots, y_{n}\right\}$ is a fundamental system of solutions of the system $x^{\prime}(t)=P(t) x(t)$. Then every other solution of this system, $z$, can be written as $z(t)=\sum_{k=1}^{n} c_{k} y_{k}(t), t \in I$, for some parameters $c_{k}$.

For the nonlinear we have the following theorem:

Theorem 3.1.6. Suppose $\left\{y_{1}, y_{2}, \ldots, y_{n}\right\}$ is a fundamental system of solutions of the system $x^{\prime}(t)=P(t) x(t)$ and $y_{p}$ is a particular solution of $x^{\prime}(t)=P(t) x(t)+f(t)$ Then every other solution of $x^{\prime}(t)=P(t) x(t)+f(t)$, say $z$, can be written as $z(t)=$ $y_{p}+\sum_{k=1}^{n} c_{k} y_{k}(t), t \in I$, for some choice of the parameters $c_{k}$.

## Homework:

Section 3.6 page 219 Problems 21, 22;
Section 3.8 page 240 Problems 1-6, 13, 14.

### 3.2 Lecture XI

Quotation: "The reader will find no figures in this work. The methods which I set forth do not require either constructions or geometrical or mechanical reasonings: but only algebraic operations, subject to a regular and uniform rule of procedure." Joseph-Louis Lagrange, Preface to Mcanique Analytique.

### 3.2.1 The eigenvalue method for homogeneous with constant coefficients

In this subsection we assume that the matrix-valued $P(t)$ is a constant function: $A=\left(a_{j, k}\right)_{j, k=1 . . n}$. We are going to consider the homogeneous problem

$$
\begin{equation*}
x^{\prime}(t)=A x(t) . \tag{3.9}
\end{equation*}
$$

We remind the reader the definition of an eigenvalue and eigenvector for a matrix $A$.

Definition 3.2.1. A complex number $\lambda$ is called an eigenvalue for the matrix $A$ if there exist a nonzero vector $u$ such that $A u=\lambda u$.

It turns out that the eigenvalues of a matrix are the zeros of its characteristic polynomial $p(\lambda)=\operatorname{det}(A-\lambda I)$. We have the following simple theorem:

Theorem 3.2.2. (a) Suppose $\lambda$ is a real eigenvalue of $A$ with a corresponding eigenvector $u$. Then the vector-valued function $v(t)=e^{\lambda t} u$ is a solution of the $D E$ (3.9).
(b) If $\lambda=p+i q$ and the corresponding eigenvector is $u=a+i b$ the $v_{1}(t)=$ $e^{p t}(a \cos q t-b \sin q t), v_{2}=e^{p t}(b \cos q t+a \sin q t)$ are solutions of the the $D E$ (3.9).

Proof. (a) Since $\frac{d v}{d t}=\lambda e^{\lambda t} u$ and $A v(t)=e^{\lambda t} A u=\lambda e^{\lambda t} u$ we see that the function $v$ satisfies (3.9).
(b) Because $\lambda$ is a solution of the characteristic polynomial whose coefficients are real, the complex conjugate of $\lambda, \bar{\lambda}$, must also be an eigenvalue. In fact, the complex conjugate of $u, \bar{u}=a-i b$, is the corresponding eigenvector of $\bar{\lambda}$. Because $w_{1}=e^{\lambda t} u$ is a solution of (3.9), as we have seen above, and then $w_{2}=e^{\bar{\lambda} t} \bar{u}$ is also a solution. Therefore, by the superposition principle, any linear combination of these is a solution also. But then, we are done, since a simple calculation shows that $v_{1}=\left(w_{1}+w_{2}\right) / 2$ and $v_{2}=-i\left(w_{1}-w_{2}\right) / 2$.

Theorem 3.2.3. Suppose the matrix $A$ has $n$ different solutions $\lambda_{1}, \ldots, \lambda_{k}, \lambda_{k+1}, \overline{\lambda_{k+1}}, \ldots$ where $\lambda_{j}$ are real for $j=1 \ldots k$ and pure complex for the rest of them. We denote by $u_{j}$ the corresponding eigenvectors of the eigenvalue $\lambda_{j}$. Then a fundamental system of solutions of (3.9) can be given by $\left\{v_{1}, \ldots, v_{n}\right\}$ where $v_{j}=e^{\lambda t} u_{j}$ for $j=1 \ldots k$ and $v_{j}=e^{p_{j} t}\left(a_{j} \cos q t-b_{j} \sin q t\right), v_{j+1}=e^{p_{j} t}\left(b_{j} \cos q t+a_{j} \sin q t\right), \ldots$ for $j \geq k+1$, where $\lambda_{j}=p+i q$ and $u_{j}=a_{j}+i b_{j}$ for $j \geq k+1$.

Proof. Let us show that the system is linearly independent. First, let us assume that all of the vector value functions are of the form $v_{j}=e^{\lambda t} u_{j}$. We observe that $u_{1}, \ldots, u_{n}$ are linearly independent as vectors in $\mathbb{R}^{n}$. Indeed, this is happening because if $\sum_{j=1}^{n} c_{j} u_{j}=0$ then, applying $A$ several times to this equality, we get $\sum_{j=1}^{n} \lambda_{j}{ }^{s} c_{j} u_{j}=0$. This implies $c_{j} u_{j l}=0$ for every $l=1$..n because the main the determinant of the homogeneous linear system obtained is a Vandermonde determinant. The Vandermonde determinant is equal to $\prod_{j<l}\left(\lambda_{j}-\lambda_{l}\right) \neq 0$. Because each $u_{j}$ is not zero, there exist a component $u_{j l}$ which is not zero. Then we get $c_{j}=0$. Hence $u_{1}, \ldots, u_{n}$ are linearly independent.

Using this we get that $\operatorname{det}\left(\left[u_{1}, \ldots, u_{n}\right]\right) \neq 0$. But then the Wronskian of $v_{1}, \ldots, v_{n}$ is $e^{\left(\lambda_{1}+\ldots+\lambda_{n}\right) t} \operatorname{det}\left(\left[u_{1}, \ldots, u_{n}\right]\right) \neq 0$. Using Theorem 3.1.4 we see that $v_{1}, \ldots, v_{n}$ are linearly independent.

The case when we have some pure complex eigenvalues, using elementary properties of determinants, we obtain that the Wronskian value for the given functions $v_{1}, \ldots, v_{n}$ change from the value calculated above for functions of the type $e^{\lambda_{j} t} u_{j}$ just by a multiple of a power of 2 .

Example: Let's take the Problem 20, page 312:

$$
\left\{\begin{array}{l}
x_{1}^{\prime}=5 x_{2}+x_{2}+3 x_{3} \\
x_{2}^{\prime}=x_{1}+7 x_{2}+x_{3} \\
x_{3}^{\prime}=3 x_{1}+x_{2}+5 x_{3} .
\end{array}\right.
$$

Here the matrix is $A=\left(\begin{array}{lll}5 & 1 & 3 \\ 1 & 7 & 1 \\ 3 & 1 & 5\end{array}\right)$. The characteristic polynomial is $p(\lambda)=$ $\operatorname{det}(A-\lambda)$. Let us notice that

$$
\begin{aligned}
& \operatorname{det}\left(\begin{array}{ccc}
5-\lambda & 1 & 3 \\
1 & 7-\lambda & 1 \\
3 & 1 & 5-\lambda
\end{array}\right)=\operatorname{det}\left(\begin{array}{ccc}
9-\lambda & 9-\lambda & 9-\lambda \\
1 & 7-\lambda & 1 \\
3 & 1 & 5-\lambda
\end{array}\right)= \\
&(9-\lambda) \operatorname{det}\left(\begin{array}{ccc}
1 & 1 & 1 \\
1 & 7-\lambda & 1 \\
3 & 1 & 5-\lambda
\end{array}\right)=(9-\lambda) \operatorname{det}\left(\begin{array}{ccc}
1 & 1 & 1 \\
0 & 6-\lambda & 0 \\
0 & -2 & 2-\lambda
\end{array}\right)= \\
&(9-\lambda)(6-\lambda)(2-\lambda) .
\end{aligned}
$$

Hence $\lambda_{1}=9, \lambda_{2}=2$, and $\lambda_{3}=6$. Three eigenvectors corresponding to these are $u_{1}=[1,1,1]^{t}, u_{2}=[-1,0,1]^{t}, u_{3}=[1,-2,1]^{t}$. Then the general solution of this system is

$$
\left\{\begin{array}{l}
x_{1}=c_{1} e^{9 t}-c_{2} e^{2 t}+c_{3} e^{6 t} \\
x_{2}=c_{1} e^{9 t}-2 c_{3} e^{6 t} \\
x_{3}=c_{1} e^{9 t}+c_{2} e^{2 t}+c_{3} e^{6 t}
\end{array}\right.
$$

## Homework:

Section 5.1 pages 297-298 Problems 1-30;
Section 5.2 page 312 problem 1-26;

### 3.3 Lexture XII

Quotation: "Regarding $\sqrt{-1}$ or what we denote these days by $i$, the building block of imaginary complex number system]: ... we can repudiate completely and which we can abandon without regret because one does not know what this pretended sign signifies nor what sense one ought to attribute to it." Augustin-Louis Cauchy said in 1847.

### 3.3.1 Multiplicity vs Dimension of the Eigenspace

In this part we assume that the system

$$
\begin{equation*}
x^{\prime}(t)=A x(t) \tag{3.10}
\end{equation*}
$$

has constant coefficients: $A=\left(a_{j, k}\right)_{j, k=1 . . n}, a_{j, k} \in \mathbb{R}$.
Definition 3.3.1. We say that an eigenvalue $\lambda_{0}$ has multiplicity $m$ if the characteristic polynomial $p(\lambda)=\operatorname{det}(A-\lambda I)=\left(\lambda-\lambda_{0}\right)^{m} q(\lambda)$ with $q\left(\lambda_{0}\right) \neq 0$. The multiplicity of $\lambda_{0}$ will be denoted by $m\left(\lambda_{0}\right)$.

Definition 3.3.2. The set of vectors $E_{\lambda_{0}}:=\left\{v: A v=\lambda_{0} v\right\}$ is a linear subspace invariant to $A$ and the dimension of it is called the dimension of the eigenspace associated to $\lambda_{0}$ denoted de $\left(\lambda_{0}\right)$.

In general we have the following relationship between these numbers:
Theorem 3.3.3. If $\lambda$ is an eigenvalue of $A$ we have $\operatorname{de}(\lambda) \leq m(\lambda)$.

In the case $m(\lambda)=d e(\lambda)$ we call $\lambda$ complete. As we have seen before we have the first general simple solution of the system (4.5) in the situation that every eigenvalue is complete. In this case we also say that $A$ is diagonalizable.

Theorem 3.3.4. If the eigenvalues of $A$ are $\lambda_{1}, \lambda_{2}, \ldots, \lambda_{k}$ for which $\operatorname{de}\left(\lambda_{j}\right)=m\left(\lambda_{j}\right)$ for every $j=1 \ldots k$ then the general solution of (4.5) is

$$
x(t)=\sum_{s=1}^{n} c_{s} e^{\lambda_{s} t} v_{s}
$$

where $v_{1}, \ldots, v_{n}$ is a basis of eigenvectors.
If we have $d e(\lambda)<m(\lambda)$ then we call the eigenvalue $\lambda$ to be defective. Let us notice that if an eigenvalue is not complete it is defective. For a defective eigenvalue the number $m(\lambda)-d e(\lambda)$ is called the defect of $\lambda$.

### 3.4 Defect one and multiplicity two

Suppose that $A v=\lambda v$ and $(A-\lambda) w=v$. It turns out that in this case there is always a solution in $w$ of this last equation. Let us check that the vector function $u(t)=(w+t v) e^{\lambda t}$ is a solution of (4.5). We have $u^{\prime}(t)=[\lambda(w+t v)+v] e^{\lambda t}$ and $A u(t)=(\lambda w+v+\lambda t v) e^{\lambda t}$. So, we have $u^{\prime}(t)=A u(t)$. This solution is linearly independent of $e^{\lambda t} v$. Indeed, if $c_{1} e^{\lambda t} v+c_{2}(w+t v) e^{\lambda t}=0$ for all $t$, we can get rid first
of $e^{\lambda t}$ to obtain $c_{1} v+c_{2}(w+t v)=0$ for all $t$. After differentiation with respect to $t$ we obtain $c_{2} v=0$ which implies $c_{2}=0$ and then automatically $c_{1}=0$. Hence the two vector valued functions that one has to take corresponding to the eigenvalue $\lambda$ are: $e^{\lambda t} v$ and $e^{\lambda t}(w+t v)$.
Example: Let us take the following example $A:=\left[\begin{array}{lll}1 & 2 & 2 \\ 2 & 1 & 2 \\ 0 & 4 & 1\end{array}\right]$ and let us say we want to solve the following initial value problem

$$
\left\{\begin{array}{l}
x^{\prime}=A x \\
x_{1}(0)=1, x_{2}(0)=2, x_{3}(0)=3
\end{array}\right.
$$

The characteristic polynomial is

$$
\begin{aligned}
& p(\lambda)=\operatorname{det}\left[\begin{array}{ccc}
1-\lambda & 2 & 2 \\
2 & 1-\lambda & 2 \\
0 & 4 & 1-\lambda
\end{array}\right]=\operatorname{det}\left[\begin{array}{ccc}
0 & (1+\lambda)(3-\lambda) / 2 & (1+\lambda) \\
2 & 1-\lambda & 2 \\
0 & 4 & 1-\lambda
\end{array}\right]= \\
& -2(1+\lambda)\left[\begin{array}{cc}
(3-\lambda) / 2 & 1 \\
4 & 1-\lambda
\end{array}\right]=-(\lambda+1)^{2}(\lambda-5) .
\end{aligned}
$$

Hence $\lambda_{1}=\lambda_{2}=-1$ with multiplicity 2 and $\lambda_{3}=5$. Let us solve for the eigenvectors of $\lambda_{3}=5$ :
$\left[\begin{array}{ccc}-4 & 2 & 2 \\ 2 & -4 & 2 \\ 0 & 4 & -4\end{array}\right]\left[\begin{array}{l}x_{1} \\ x_{2} \\ x_{3}\end{array}\right]=\left[\begin{array}{l}0 \\ 0 \\ 0\end{array}\right]$ which gives simply a one-dimensional eigenspace
$\left[\begin{array}{l}x_{1} \\ x_{2} \\ x_{3}\end{array}\right]=t\left[\begin{array}{l}1 \\ 1 \\ 1\end{array}\right], t \in \mathbb{R}$.
We will just take $v_{1}=\left[\begin{array}{l}1 \\ 1 \\ 1\end{array}\right]$.
Now we solve $\left(A-\lambda_{1}\right) v=0$ : $\left[\begin{array}{lll}2 & 2 & 2 \\ 2 & 2 & 2 \\ 0 & 4 & 2\end{array}\right]\left[\begin{array}{l}x_{1} \\ x_{2} \\ x_{3}\end{array}\right]=0$ which gives also a onedimensional eigenspace generated by $v_{2}=\left[\begin{array}{c}1 \\ 1 \\ -2\end{array}\right]$.

Then we solve for $\left(A-\lambda_{2}\right) w=v_{2}$. One solution of this equation can be taken to be $w=\left[\begin{array}{c}3 / 2 \\ 0 \\ -1\end{array}\right]$. Therefore the general solution of our equation is $x(t)=$
$c_{1} v_{1} e^{5 t}+\left[c_{2} v_{2}+c_{3}\left(w+t v_{2}\right)\right] e^{-t}$. This gives

$$
\left\{\begin{array}{l}
x_{1}(t)=c_{1} e^{5 t}+\left[c_{2}+c_{3}(3 / 2+t)\right] e^{-t} \\
x_{2}(t)=c_{1} e^{5 t}+\left(c_{2}+c_{3} t\right) e^{-t} \\
x_{3}(t)=c_{1} e^{5 t}-\left[2 c_{2}+c_{3}(1+2 t)\right] e^{-t}
\end{array} .\right.
$$

Then we need to determine the constants $c_{1}, c_{2}, c_{3}$ in order to get the initial condition satisfied. This gives $c_{1}=\frac{19}{9}, c_{2}=-\frac{1}{9}$, and $c_{3}=-\frac{2}{3}$. Finally substituting we get

$$
\left\{\begin{array}{l}
x_{1}(t)=\frac{19}{9} e^{5 t}-\frac{10+6 t}{9} e^{-t} \\
x_{2}(t)=\frac{19}{9} e^{5 t}-\frac{1+6 t}{9} e^{-t} \\
x_{3}(t)=\frac{19}{9} e^{5 t}-\frac{8+12 t}{9} e^{-t}
\end{array} .\right.
$$

### 3.4.1 Generalized vectors

Let $\lambda$ be an eigenvalue of $A$.
Definition 3.4.1. $A$ vector $v$ is called a rank r generalized eigenvector associated to $\lambda$ if $(A-\lambda I)^{r} v=0$ and $(A-\lambda I)^{r-1} v \neq 0$.

Clearly, every eigenvector is a rank one generalized eigenvector. Notice that if $v$ is a rank $r$ generalized eigenvector associated to $\lambda$ then we can define $v_{1}=$ $(A-\lambda I)^{r-1} v$ which is not zero by definition and satisfies $A v_{1}=\lambda v_{1}$. This means $v_{1}$ is a regular eigenvector. We define in general $v_{2}=(A-\lambda I)^{r-2} v, \ldots, v_{r-1}=(A-\lambda) v$, $v_{r}=v$. These vectors are all not equal to zero because otherwise $v_{1}$ becomes zero. In practice we need to work our way backwards in order to determine a generalized eigenvector. First we find $v_{1}$ as usual since it is a eigenvector. Then we find $v_{2}$ from the equation $(A-\lambda I) v_{2}=v_{1}$. Next we solve for $v_{3}$ from the equation $(A-\lambda I) v_{3}=v_{2}$ and so on. One can show that $v_{1}, v_{2}, \ldots, v_{r}$ are linearly independent. We have the following important theorem from linear algebra which is sometime called the Jordan representation theorem because it allows one to represent the matrix, up to a similarity, i.e. $S A S^{-1}$, as a direct sum of Jordan blocks: $\lambda I_{k}+N$ where $N$ is a nilpotent matrix that has ones above the diagonal and zero for the rest of the entries.

Theorem 3.4.2. For every $n \times n$ matrix $A$ there exists a basis of generalized eigenvectors. For each eigenvalue $\lambda$ of multiplicity $m(\lambda)$ there exists $m(\lambda)$ generalized linearly independent vectors associated.

In general, if we have an eigenvector $v_{1}$ such that $v_{r}$ is is a rank $r$ generalized eigenvector corresponding to the eigenvalue $\lambda$, the contribution of these to a
fundamental set of solutions for (4.5) is with the following set of functions:

$$
\begin{aligned}
& u_{1}(t)=e^{\lambda t} v_{1}, \\
& u_{2}(t)=\left(v_{2}+t v_{1}\right) e^{\lambda t} \\
& u_{3}(t)=\left(v_{3}+t v_{2}+\frac{t^{2}}{2!} v_{1}\right) e^{\lambda t} \\
& \ldots, \\
& u_{r}(t)=\left(v_{r}+t v_{r-1}+\frac{t^{2}}{2!} v_{r-2}+\ldots+\frac{t^{r-1}}{(r-1)!} v_{1}\right) e^{\lambda t} .
\end{aligned}
$$

One can show that these are linearly independent vector-valued functions which are solutions of (4.5). We may have for a certain eigenvalue different sets of this type. Putting all together will give a fundamental set of solutions of (4.5). This fact is insured by the Theorem 3.4.2. In the case the eigenvalue is pure complex we just take the real and imaginary parts of these vector valued functions.

## Example: Find the general solution of the differential system:

$$
\left\{\begin{array}{l}
x_{1}^{\prime}=x_{1}+x_{2} \\
x_{2}^{\prime}=x_{2}+x_{3} \\
x_{3}^{\prime}=x_{3}
\end{array}\right.
$$

The matrix of this system is $A=\left[\begin{array}{lll}1 & 1 & 0 \\ 0 & 1 & 1 \\ 0 & 0 & 1\end{array}\right]$. Clearly we have $\lambda=1$ as eigenvalue of multiplicity 3 and a corresponding eigenvector is $v_{1}=\left[\begin{array}{l}1 \\ 0 \\ 0\end{array}\right]$. Then the equation $(A-\lambda I) v_{2}=v_{1}$ has a nonzero solution $v_{2}=\left[\begin{array}{l}0 \\ 1 \\ 0\end{array}\right]$. This makes $v_{2}$ a rank 2 generalized vector. If we continue, the equation $(A-\lambda I) v_{3}=v_{2}$ has a nontrivial solution $v_{3}=\left[\begin{array}{l}0 \\ 0 \\ 1\end{array}\right]$. This means that $v_{3}$ is a rank 3 generalized vector. Then a fundamental set of solutions of our system is $u_{1}(t)=e^{t} v_{1}, u_{2}=\left(t v_{1}+v_{2}\right) e^{t}$ and $u_{3}=\left(t^{2} v_{1} / 2+t v_{2}+v_{3}\right) e^{t}$. Therefore, the general solution of the differential system is $\left\{\begin{array}{l}x_{1}=\left(c_{1} e^{t}+c_{2} t+c_{3} t^{2} / 2\right) e^{t} \\ x_{2}=\left(c_{2}+t c_{3}\right) e^{t} \\ x_{3}=c_{3} e^{t}\end{array}\right.$.

### 3.4.2 Fundamental Matrix of Solutions, Exponential Matrices

Suppose we are still solving a system $x^{\prime}=A x$ as before. A fundamental matrix of solutions is a $n \times n$ matrix formed with $n$ linearly independent solutions of the system (4.5). Such a matrix can be simply calculated by using the Taylor formula, $e^{x}=1+x+x^{2} / 2!+\ldots$, for matrices instead of numbers.

Definition 3.4.3. The matrix $e^{t A}$ is the result of the infinite series $\sum_{k=0}^{\infty} A^{k} \frac{t^{k}}{k!}$, $t \in \mathbb{R}$.

Theorem 3.4.4. The matrix $e^{t A}$ is a fundamental matrix of solutions of $x^{\prime}=A x$. The solution of the initial value problem $\left\{\begin{array}{l}x^{\prime}=A x \\ x(0)=x_{0}\end{array} \quad\right.$ is given by $x(t)=e^{t A} x_{0}$.

Example: Let us take Problem 12, page 356 as an exemplification of this. The system given is $\left\{\begin{array}{l}x_{1}^{\prime}=5 x_{1}-4 x_{2} \\ x_{2}^{\prime}=3 x_{1}-2 x_{2}\end{array} \quad\right.$ with the matrix $A=\left[\begin{array}{ll}5 & -4 \\ 3 & -2\end{array}\right]$. We need to calculate $e^{t A}$.

The characteristic polynomial of $A$ is $p(\lambda)=\lambda^{2}-3 \lambda+2$. (In general the characteristic polynomial for $2 \times 2$ matrix $A=\left[\begin{array}{ll}a & b \\ c & d\end{array}\right]$ is $\lambda^{2}-\operatorname{tr}(A) \lambda+\operatorname{det}(A)$, where $\operatorname{tr}(A)=a+d)$. The eigenvalues in this case are $\lambda_{1}=1$ and $\lambda_{2}=2$. Two eigenvectors are in this case $v_{1}=\left[\begin{array}{l}1 \\ 1\end{array}\right]$ and $v_{2}=\left[\begin{array}{l}4 \\ 3\end{array}\right]$ corresponding to $\lambda_{1}$ and $\lambda_{2}$ respectively. If we denote the matrix $\left[v_{1} \mid v_{2}\right]$ by $S:=\left[\begin{array}{ll}1 & 4 \\ 1 & 3\end{array}\right]$ then let us observe that $A S=\left[v_{1} \mid 2 v_{2}\right]=S D$ where $D:=\left[\begin{array}{ll}1 & 0 \\ 0 & 2\end{array}\right]$ is a diagonal matrix with entries exactly the two eigenvalues. Therefore $A=S D S^{-1}$. This allows us to calculate $A^{n}=S D^{n} S^{-1}$. Notice that $e^{t D}=\left[\begin{array}{cc}e^{t} & 0 \\ 0 & e^{2 t}\end{array}\right]$. In general the inverse of a $2 \times 2$ matrix matrix $A=\left[\begin{array}{ll}a & b \\ c & d\end{array}\right]$ is given by $A^{-1}=\frac{1}{\operatorname{det}(A)}\left[\begin{array}{cc}d & -b \\ -c & a\end{array}\right]$. Hence

$$
e^{t A}=S e^{t D} S^{-1}=\left[\begin{array}{ll}
1 & 4 \\
1 & 3
\end{array}\right]\left[\begin{array}{cc}
e^{t} & 0 \\
0 & e^{2 t}
\end{array}\right]\left[\begin{array}{cc}
-3 & 4 \\
1 & -1
\end{array}\right]
$$

or

$$
e^{t A}=\left[\begin{array}{ll}
-3 e^{t}+4 e^{2 t} & 4 e^{t}-4 e^{2 t} \\
-3 e^{t}+3 e^{2 t} & 4 e^{t}-3 e^{2 t}
\end{array}\right]
$$

## Homework:

Section 5.4 pages 342-343 Problems 23-33;
Section 5.5 page 356 problems 1-20, 25-30;

## Chapter 4

## Nonlinear Systems and Qualitative Methods

### 4.1 Lecture XIII

Quotation: "I entered an omnibus to go to some place or other. At that moment when I put my foot on the step the idea came to me, without anything in my former thoughts seeming to have paved the way for it, that the transformations I had used to define the Fuchsian functions were identical with non-Euclidean geometry." Henri Poincaré

### 4.1.1 Nonlinear systems and phenomena, linear and almost linear systems

We are going to discuss the behavior for solutions of autonomous systems DE of the form

$$
\left\{\begin{array}{l}
\frac{d x}{d t}=F(x, y)  \tag{4.1}\\
\frac{d y}{d t}=G(x, y)
\end{array}\right.
$$

Let us assume that the two functions $F$ and $G$ are continuous on a region $\mathcal{R}=\left\{(x, y) \in \mathbb{R}^{2} \mid a<x<b, c<y<d\right\}$. This region is going to be called phase plane. By a similar theorem of existence and uniqueness we have a unique solution to the IVP:

$$
\left\{\begin{array}{l}
\frac{d x}{d t}=F(x, y)  \tag{4.2}\\
\frac{d y}{d t}=G(x, y) \\
x(0)=x_{0}, y(0)=y_{0}
\end{array}\right.
$$

where $\left(x_{0}, y_{0}\right) \in \mathcal{R}$. The curve $(x(t), y(t)), t \in(-\epsilon, \beta)$ represented in $\mathcal{R}$ is called a trajectory. For each point of $\mathcal{R}$ there exist one and only one trajectory containing it.

Definition 4.1.1. A critical point for the system (4.5) is a point $(a, b) \in \mathcal{R}$ such that $F(a, b)=G(a, b)=0$.

Clearly, if $(a, b)$ is a critical point the constant function $(x(t), y(t))=(a, b)$ for every $t \in \mathbb{R}$ is a solution of (4.5) which is called an equilibrium solution. We observe that the trajectory of a critical point is just a point.

The phase portrait is a sketch of the phase plane and a few typical trajectories together with all critical points.

Definition 4.1.2. A critical point $(a, b)$ is called a node if either every trajectory approaches $(a, b)$ or every trajectory recedes form $(a, b)$ tangent to a line at $(a, b)$. A node can be $a \operatorname{sink}$ if all trajectories approach the critical point or a source if all trajectories emanate from it.

A node can be proper or improper depending upon the number of tangent lines that the trajectories have: infinitely many or only two. The following notion of stability is the same as for ODE case:

Definition 4.1.3. A critical point $(a, b)$ is called stable if for every $\epsilon>0$ there exists $a \delta>0$ such that if $\left|x_{0}-a\right|<\delta$ and $\left|y_{0}-b\right|<\delta$ the solution of the IVP (4.2) statistics $|x(t)-a|<\epsilon$ and $|y(t)-b|<\epsilon$.

In general nodal sinks are stable critical points. A critical point which is not stable is called unstable.

A critical point can be stable without having the trajectories approach the critical point. If a stable critical point is surrounded by simple closed trajectories representing periodic solutions then such a critical point is called (stable) center.

Definition 4.1.4. A critical point is called asymptotically stable if it is stable and for some $\delta>0$ if $\left|x_{0}-a\right|<\delta$ and $\left|y_{0}-b\right|<\delta$ then $\lim _{t \rightarrow \infty} x(t)=a$ and $\lim _{t \rightarrow \infty} y(t)=b$.

An asymptotically stable critical point with the property that every trajectory spirals around it is called spiral sink. A spiral source is a critical point as before but with time $t$ going to $-\infty$ instead of $\infty$. If for a critical point there are two trajectories that approach the critical point but all the other ones are unbounded as $t \rightarrow \infty$ then we say the critical point is a saddle point. Under certain conditions one can show that there are only four possibilities for trajectories:

1. $(x(t), y(t))$ approaches a critical point $(a, b)$ as $t \rightarrow \infty$;
2. $(x(t), y(t))$ is unbounded;
3. $(x(t), y(t))$ is a periodic solution;
4. $(x(t), y(t))$ spirals toward a periodic solution as $t \rightarrow \infty$.

### 4.2 Linear and almost linear systems

In the linear case, as usual, we can always have more to say. A critical point $(a, b)$ is called isolated if there are no other critical points in a neighborhood of $(a, b)$. We assume that $F$ and $G$ are differentiable around $(a, b)$. The linearized sytem associated to (4.5) is

$$
\left\{\begin{array}{l}
\frac{d x}{d t}=\frac{\partial F}{\partial x}(a, b)(x-a)+\frac{\partial F}{\partial y}(a, b)(y-b)  \tag{4.3}\\
\frac{d y}{d t}=\frac{\partial G}{\partial x}(a, b)(x-a)+\frac{\partial G}{\partial y}(a, b)(y-b) .
\end{array}\right.
$$

A system is said to be almost linear at the isolated critical point $(a, b)$ if

$$
\lim _{x \rightarrow a, y \rightarrow b} \frac{F(x, y)-\frac{\partial F}{\partial x}(a, b)(x-a)-\frac{\partial F}{\partial y}(a, b)(y-b)}{\sqrt{(x-a)^{2}+(y-b)^{2}}}=0
$$

and

$$
\lim _{x \rightarrow a, y \rightarrow b} \frac{G(x, y)-\frac{\partial G}{\partial x}(a, b)(x-a)-\frac{\partial G}{\partial y}(a, b)(y-b)}{\sqrt{(x-a)^{2}+(y-b)^{2}}}=0 .
$$

In this case its linearization is (4.8). This is a linear system with constant coefficients. In practice, the functions $F$ and $G$ are continuously differentiable. This insures that the system (4.5) is almost linear around a critical point.

| Eigenvalues $\lambda_{1}, \lambda_{2}$ |  |
| :---: | :---: |
| Type of critical point |  |
| Real, different, same sign | Improper node |
| Real, unequal, opposite signs | Saddle |
| Real and equal, | Proper or improper node |
| Pure complex, | Spiral point |
| Pure imaginary | Center |

Table 4.1: Type of critical points for linear systems

### 4.3 Critical points classification of linear systems

We are going to consider only two dimensional linear systems of the type

$$
\left\{\begin{array}{l}
x^{\prime}=a x+b y  \tag{4.4}\\
y^{\prime}=c x+d y
\end{array}\right.
$$

which we already know how to solve. Let us assume that the two eigenvalues are $\lambda_{1}$ and $\lambda_{2}$. Then, what type of critical point $(0,0)$ of $(4.4)$ is could be determined by the following chart:

For the stability of the critical point of (4.4) we have:
Theorem 4.3.1. Let $\lambda_{1}$ and $\lambda_{2}$ be the eigenvalues of the system (4.4) which has $(0,0)$ as an isolated critical point. The critical point $(0,0)$ is
(a) Asymptotically stable if the real parts of $\lambda_{1}$ and $\lambda_{2}$ are both negative;
(b) Stable but not asymptotically stable if the real parts of $\lambda_{1}$ and $\lambda_{2}$ are both zero;
(c) Unstable if either $\lambda_{1}$ or $\lambda_{2}$ has a positive real part.

The next theorem says that for an almost liner system the effect of small perturbations around an isolated critical point is almost determined by its linearization at that point.

Theorem 4.3.2. Let $\lambda_{1}$ and $\lambda_{2}$ be the eigenvalues of the linearization system at an isolated critical point of an almost linear system (4.5). Then
(a) If $\lambda_{1}=\lambda_{2}$, we have a node or spiral point; in this case the critical point is asymptotically stable if $\lambda_{1}=\lambda_{2}<0$ or unstable if $\lambda_{1}=\lambda_{2}>0$.
(b) If $\lambda_{1}$ and $\lambda_{2}$ are pure imaginary, then we have either a center or a spiral point (and undetermined stability)
(c) Otherwise the type and stability of $(a, b)$ is the same as the one for the linearization system.

## Homework:

Section 6.1 pages 375-376 Problems 1-8;
Section 6.2 page 389 Problems 1-32;

### 4.4 Lecture XIV

Quotation: "Everything is vague to a degree you do not realize till you have tried to make it precise." Bertrand Russell British author, mathematician and philosopher (1872-1970)

### 4.4.1 Nonlinear spring

Let us assume that Hooks's law is "corrected" a little to: $F=-k x+\beta x^{3}$. In a sense, it is natural to think that there are some other terms in the Taylor expansion of the force acting on a body of mass $m$ in terms of the displacement $x$. In this case the differential equation that we obtain from Newton's law is $m x^{\prime \prime}=-k x+\beta x^{3}$. We can turn this into a system if we introduce $y=\frac{d x}{d t}$ :

$$
\left\{\begin{array}{l}
\frac{d x}{d t}=y  \tag{4.5}\\
\frac{d y}{d t}=-\frac{k}{m} x+\frac{\beta}{m} x^{3} .
\end{array}\right.
$$

The discussion of what happens with the mechanical system depends clearly on $\beta$.

Case $\beta<0$ "hard spring": In this situation we have only one critical point: $(0,0)$. The Jacobian of the almost linear system (4.5) is

$$
J(x, y)=\left[\begin{array}{cc}
0 & 1 \\
-\frac{k}{m}+3 \frac{\beta}{m} x^{2} & 0
\end{array}\right]
$$

so at the critical point $J(0,0)=\left[\begin{array}{cc}0 & 1 \\ -\frac{k}{m} & 0\end{array}\right]$
The eigenvalues of $J(0,0)$ are $\pm i \sqrt{\frac{k}{m}}$ and according to the theorem about transfer of stability and type from the previous lecture we see that all we can say is that $(0,0)$ is either a center or a spiral point. From the point stability the theorem does not say what happens.

But we can determine what is going on by integrating (4.5). By multiplying the second equation in (4.5) by $y$ and integrating we obtain

$$
\frac{y^{2}}{2}+\frac{k x^{2}}{2 m}+\frac{|\beta| x^{4}}{4 m}=\text { constant }
$$

These curves are almost like circles around $(0,0)$ (see figure below). The stability but not asymptotic stability of $(0,0)$ follows.


Figure 1
This means that we have almost regular oscillations around the equilibrium position.

Case $\beta>0$ "soft spring": The system has two more critical points: $\left( \pm \sqrt{\frac{k}{\beta}}, 0\right)$. The Jacobian at the critical points is at the critical point $J\left( \pm \sqrt{\frac{k}{\beta}}, 0\right)=\left[\begin{array}{cc}0 & 1 \\ \frac{2 k}{m} & 0\end{array}\right]$ which means the eigenvalues are real one positive and on negative: $\pm \sqrt{\frac{2 k}{m}}$. so by the same theorem we have that both the new critical points are saddle points and unstable.

In fact if we use the method of integration as before we get

$$
\frac{y^{2}}{2}+\frac{k x^{2}}{2 m}-\frac{\beta x^{4}}{4 m}=c
$$

The phase portrait is as below:


Figure 2
This suggests that for a certain constant $c$ (say $c_{0}$, the Energy) we get solutions which go directly to an static position without any oscillations. For values of $c$ smaller than that quantity we have oscillations around the equilibrium position. For values of the energy bigger than $c_{0}$ the displacement goes to infinity. This is not corresponding to anything real so it seems like this model does not work.

### 4.4.2 Chaos

Let us start with the logistic differential equation

$$
\begin{equation*}
\frac{d P}{d t}=a P-b P^{2}, \quad(a, b>0) \tag{4.6}
\end{equation*}
$$

which models in general a bounded population.
We are going to look into the discretization of this equation according to Euler's method: compute $P_{n+1}-P_{n}=\left(a P_{n}-b P_{n}^{2}\right) h$ which are the values of the approximation of $P(t)$ at the times $t_{k}=t_{0}+k h$, where $h$ is the step size. This can written as $P_{n+1}=r P_{n}-s P_{n}^{2}$ where $r=1+a h$ and $s=b h$. Then if we substitute $P_{n}=\frac{r}{s} x_{n}$ then the recurrence becomes

$$
\begin{equation*}
x_{n+1}=r x_{n}\left(1-x_{n}\right) . \tag{4.7}
\end{equation*}
$$

We notice that the function $f(x)=r x(1-x)$ has a maximum at $x=1 / 2$ equal to $r / 4$. So, for $r<4$ this function maps the interval $[0,1]$ into itself. So one can iterate
and compute the sequence defined in (4.7) for large values of $n$. It turns out that for $r<r_{c}$ where $r_{c} \approx 3.57$.. the behavior of the sequence can be predicted in the sense that approaches a certain number of limit points. The sequence becomes chaotic for $r=r_{c}$, i.e. there are infinitely many limit points that fill out the interval $[0,1]$. In other words the sequence is unpredictable. For instance we cannot tell, unless we compute it precisely where $x_{1000000}$ is in the interval $[0,1]$.

The number of periodic orbits for $r<r_{c}$ changes at some specific values $r_{k}$. The mathematician Mitchell Feigenbaum discovered that

$$
\lim _{n \rightarrow \infty} \frac{r_{k}-r_{k-1}}{r_{k+1}-r_{k}}=\text { constant } \approx 4.669 \ldots
$$

This constant appears in other places in mathematics, so it gained a status similar to those constants such as $\pi, e$ or $\gamma$. Some differential equations have the same type of behavior for certain values of the parameters. Such examples are $m x^{\prime \prime}+c x^{\prime}+$ $k x+\beta x^{3}=F_{0} \cos \omega t$ (forced Duffing equation) or the famous Lorenz system

$$
\left\{\begin{array}{l}
\frac{d x}{d t}=-s x+s y  \tag{4.8}\\
\frac{d y}{d t}=-x z+r x-y \\
\frac{d z}{d t}=x y-b z
\end{array}\right.
$$

## Homework:

Section 6.4 page 418, Problems 2, 5-8;
Section 7.1 pages 444-445, Problems 1-42.

## Chapter 5

## Laplace Transform

### 5.1 Lecture XV

Quotation: "Given for one instant an intelligence which could comprehend all the forces by which nature is animated and the respective positions of the beings which compose it, if moreover this intelligence were vast enough to submit these data to analysis, it would embrace in the same formula both the movements of the largest bodies in the universe and those of the lightest atom; to it nothing would be uncertain, and the future as the past would be present to its eyes." Pierre Simon De Laplace (17491827), French mathematician, philosopher. Theorie Analytique de Probabilites: Introduction, v. VII, Oeuvres (1812-1820).

### 5.1.1 Definition and a few examples

The Laplace transform is a transformation on functions as the operator $D$ of differentiation that we have encountered earlier. The study of it in this course is motivated by the fact that some differential equations can be converted via the Laplace transform into an algebraic equation. This is in general thought as being easier to solve and then one obtains the solution of the given differential equation by taking the inverse Laplace transform for the solution of the the corresponding algebraic equation.

In order to define this transform we need a few definitions beforehand.
Definition 5.1.1. A function $f$ is called piecewise continuous on the interval $[a, b]$ if there is a partition of the interval $x_{0}=a<x_{1}<x_{2}<\ldots<x_{n}=b$ such that $f$ is continuous on each interval $\left(x_{k}, x_{k+1}\right)$ and it has sided limits at each point $x_{k}$.

A function $f$ defined on an unbounded interval is said to be piecewise continuous if it is so on each bounded subinterval.

Definition 5.1.2. A function $f:[0, \infty) \rightarrow \mathbb{R}$ is called of exponential type at $\infty$ if there exist nonnegative constants $M, c$ and $T$ such that $|f(t)| \leq M e^{c t}$ for all $t \geq T$.

Definition 5.1.3. For every function $f:[0, \infty) \rightarrow \mathbb{R}$ which is piecewise continuous on some interval $[T, \infty)$, integrable on $[0, T]$ and of exponential type at infinity the Laplace transform $\mathcal{L}(f)$ is the new function of the variable $s$ defined by

$$
\mathcal{L}(f)(s)=\int_{0}^{\infty} e^{-s t} f(t) d t
$$

The domain of $\mathcal{L}(f)$ is taken to be the set of all $s$ for which the improper integral exists, i.e. $\lim _{n \rightarrow \infty} \int_{0}^{n} e^{-s t} f(t) d t$ exists.

The next theorem tells us that the above definition is meaningful. We are going to denote the class of these functions by $\mathcal{D}(\mathcal{L})$.

Theorem 5.1.4. Under the assumption in the definition (5.1.3)the Laplace transform $\mathcal{L}(f)(s)$ exists for every $s>c$.

Before we prove this theorem let us compute the Laplace transform for some simple functions.
Example 1: Suppose we take $f(t)=1$ for all $t \in[0, \infty)$. Then $\mathcal{L}(f)(s)=$ $\int_{0}^{\infty} e^{-s t} d t=-\left.\frac{e^{-s t}}{s}\right|_{0} ^{\infty}=\frac{1}{s}$, for all $s>0$. Therefore we write

$$
\mathcal{L}(1)(s)=\frac{1}{s}, s>0
$$

Example 2: Let us take $f(t)=e^{c t}$ for all $t \geq 0$. Then if $s>c, \mathcal{L}(f)(s)=$ $\int_{0}^{\infty} e^{-s t} e^{c t} d t=\int_{0}^{\infty} e^{-(s-c) t} d t=-\left.\frac{e^{-(s-c) t}}{s-c}\right|_{0} ^{\infty}=\frac{1}{s-c}$. Hence

$$
\mathcal{L}\left(e^{c \cdot}\right)(s)=\frac{1}{s-c}, s>c
$$

Proof of Theorem 5.1.4. We need to show that the limit $\lim _{n \rightarrow \infty} \int_{0}^{n} e^{-s t} f(t) d t$ exists. Using Cauchy's characterization of the existence of a limit it suffices to show that $\lim _{n, m \rightarrow \infty} \int_{m}^{n} e^{-s t} f(t) d t=0$. We have the estimate

$$
\begin{aligned}
& \left|\int_{m}^{n} e^{-s t} f(t) d t\right| \leq \int_{m}^{n} e^{-s t}|f(t)| d t \leq M \int_{m}^{n} e^{-(s-c) t} d t= \\
& \frac{M}{(s-c)}\left(e^{-(s-c) m}-e^{-(s-c) n}\right) \rightarrow 0
\end{aligned}
$$

as $m, n \rightarrow \infty(T \leq m<n)$ provided that $s>c$.
Note: Cauchy's characterization is one of the most common tools in analysis. Augustin Louis Cauchy was born on 21st of August, 1789 in Paris, France, and died May 23rd, 1857 in Sceaux near Paris.

### 5.1.2 General Properties

Corollary 5.1.5. Let us assume that $f$ is as in the Theorem 5.1.4. Then $\lim _{s \rightarrow \infty} \mathcal{L}(f)(s)=$ 0 .

Proof. Since we know the limit in the definition of $\mathcal{L}(f)(s)$ exists we let $n$ go to infinity in the sequence of inequalities (5.1) but fix $m=T$. That gives

$$
|\mathcal{L}(f)(s)| \leq\left|\int_{0}^{m} e^{-s t} f(t) d t\right|+\frac{M}{s-c}
$$

for every $s>c$ and the conclusion of our corollary follows from this and a theorem of convergence under the integral sign.

The Laplace transform may exist even for functions that are unbounded on a finite interval. One such example is $f(t)=t^{a}, t>0$ with $a>-1$. Notice that for $a \in(-1,0)$ the integral $\int_{0}^{\infty} e^{-s t} t^{a} d t$ is also improper at 0 . To compute $\mathcal{L}(f)(s)$ we change the variable $s t=u(s>0)$ and obtain

$$
\mathcal{L}(f)(s)=\frac{1}{s^{a+1}} \int_{0}^{\infty} e^{-u} u^{a} d u=\frac{\Gamma(a+1)}{s^{a+1}},
$$

where $\Gamma$ is defined by $\Gamma(x)=\int_{0}^{\infty} e^{-t} t^{x-1} d x$ and exists for all $x>0$. An integration by parts shows that

$$
\Gamma(x+1)=\int_{0}^{\infty} e^{-t} t^{x} d x=-\left.e^{-t} t^{x}\right|_{t=0} ^{t=\infty}+x \int_{0}^{\infty} e^{-t} t^{x-1} d x=x \Gamma(x) .
$$

because $\Gamma(1)=1$ we get by induction $\Gamma(n+1)=n$ ! for $n \in \mathbb{N}$.
In particular we get for instance $\mathcal{L}\left(t^{5}\right)(s)=\frac{\Gamma(6)}{s^{6}}=\frac{5!}{s^{6}}, s>0$. For fractional values of $a$ one needs to know $\Gamma(a)$. One interesting fact here is that $\Gamma(1 / 2)=\sqrt{\pi}$. To see this let us change the variable $t=u^{2}$ in $\Gamma(1 / 2)=\int_{0}^{\infty} e^{-t} t^{-1 / 2} d t$.

We obtain $\Gamma(1 / 2)=\int_{0}^{\infty} e^{-u^{2}} u^{-1} 2 u d u=2 \int_{0}^{\infty} e^{-u^{2}} d u$. Now we calculate $\Gamma(1 / 2)^{2}=4 \int_{0}^{\infty} e^{-u^{2}} d u \int_{0}^{\infty} e^{-v^{2}} d v=4 \int_{0}^{\infty} \int_{0}^{\infty} e^{-\left(u^{2}+v^{2}\right)} d u d v$. We then use polar coordinates $u=r \cos s, v=r \sin s$. So we get $\Gamma(1 / 2)^{2}=4 \int_{0}^{\pi / 2}\left(\int_{0}^{\infty} e^{-r^{2}} r d r\right) d s=\pi$ which implies $\Gamma(1 / 2)=\sqrt{\pi}$. This allows one to compute for instance $\Gamma(3 / 2)=$ $\frac{3}{2} \Gamma(1 / 2)=\frac{3 \sqrt{\pi}}{2}$.

Proposition 5.1.6. The Laplace transform is linear.
Proof. The integral and the limit are linear transformations on functions. One needs to check also that $\mathcal{D}(\mathcal{L})$ is a linear space of functions.
Example: $\mathcal{L}\left(3 t^{2}+2 \sqrt{t}\right)(s)=3 \mathcal{L}\left(t^{2}\right)(s)+2 \mathcal{L}(\sqrt{t})(s)=3 \frac{2!}{s^{3}}+2 \frac{\Gamma(3 / 2)}{s^{3 / 2}}=\frac{6}{s^{3}}+\frac{3 \sqrt{\pi}}{s \sqrt{s}}$.
Another example we would like to do involves the Laplace transform of a complex valued function which is a natural extension of the Laplace transform of real valued functions.

Example: We take $f(t)=e^{z t}$ where $z=a+i b$. Notice that $|f(t)|=e^{a t}$ so this function is of exponential type at infinity. We have $\mathcal{L}(f)(s)=\int_{0}^{\infty} e^{-s t} e^{z t} d t=$ $\int_{0}^{\infty} e^{-(s-z) t} d t=\lim _{t \rightarrow \infty}\left(\frac{1}{s-z}-\frac{e^{-(s-z) t}}{s-z}\right)=\frac{1}{s-z}$, provided that $s>a$.

This is happening because $\left|\frac{e^{-(s-z) t}}{s-z}\right|=\frac{e^{-(s-a) t}}{|s-z|} \rightarrow 0$ as $t \rightarrow \infty$. Since $\mathcal{L}$ is linear $\operatorname{Re} \mathcal{L}(f)(s)=\mathcal{L}\left(e^{a t} \cos b t\right)(s)=\operatorname{Re} \frac{1}{s-z}=\frac{s-a}{(s-a)^{2}+b^{2}}$ and $\operatorname{Im} \mathcal{L}(f)(s)=$ $\mathcal{L}\left(e^{a t} \sin b t\right)(s)=\operatorname{Im} \frac{1}{s-z}=\frac{b}{(s-a)^{2}+b^{2}}$.
Example: This example involves the Laplace transform of a function denoted by $u$ and defined by

$$
u(t)=\left\{\begin{array}{lll}
1 & \text { if } & x \geq 0 \\
0 & \text { if } & x<0,
\end{array} \quad \text { or a translation of } u \text { which is denoted by } u_{a}\right. \text { and defined }
$$ as $u_{a}(t)=u(t-a), t \in \mathbb{R}$.

We get $\mathcal{L}\left(u_{a}\right)(s)=\int_{0}^{\infty} e^{-s t} u_{a}(t) d t=\int_{a}^{\infty} e^{-s t} d t=\frac{e^{-s a}}{s}$. Let us record the main formulas that we have discovered so far in the table 5.1.2:

Proposition 5.1.7. The Laplace transform is a one-to-one map in the following sense: $\mathcal{L}(f)(s)=\mathcal{L}(g)(s)$ for all $s>s_{0}$ implies that the functions $f$ and $g$ coincide at all their continuity points.

Since this is true we are allowed to take the inverse of the Laplace transform denoted by $\mathcal{L}^{-1}$ by simply inverting the table above (in other words it is not ambiguous to talk about the inverse of the Laplace transform).

| $f(t)$ | $\mathcal{L}(f)(s)$ |
| ---: | ---: |
| $t^{n}$ | $\frac{n!}{s^{n+1}}$ |
| $e^{a t} \cos b t$ | $\frac{s-a}{(s-a)^{2}+b^{2}}$ |
| $e^{a t} \sin b t$ | $\frac{b}{(s-a)^{2}+b^{2}}$ |
| $u_{a}(t)$ | $\frac{e^{-s a}}{s}$ |
| $t^{a}(a>-1)$ | $\frac{\Gamma(a+1)}{s^{a+1}}$ |
| $e^{a t}$ | $\frac{1}{s-a}$ |

Table 5.1: Laplace Transform Formulae

Theorem 5.1.8. Given the function $f:[0, \infty) \rightarrow \mathbb{R}$ of exponential type at infinity which is continuous and whose derivative is piecewise continuous, then $\mathcal{L}\left(f^{\prime}\right)$ exists and $\mathcal{L}\left(f^{\prime}\right)(s)=s \mathcal{L}(f)(s)-f(0)$.

Proof. First we assume that the derivative is continuous at all points. Then an integration by parts will give

$$
\mathcal{L}(f)(s)=\int_{0}^{\infty} e^{-s t} f^{\prime}(t) d t=\left.e^{-s t} f(t)\right|_{0} ^{\infty}+s \int_{0}^{\infty} e^{-s t} f(t) d t=s \mathcal{L}(f)(s)-f(0)
$$ where $\lim _{t \rightarrow \infty} e^{-s t} f(t)=0$ because of the hypothesis on $f$ to be of exponential type.

The proof in the general case goes the same way with the only change that the fundamental formula of calculus holds true for $f$ under the given hypothesis.

Corollary 5.1.9. If the function $f$ is of exponential type and it has derivatives of order $k,(k \leq n)$, exist with $f^{(n)}$ piecewise continuous then

$$
\mathcal{L}\left(f^{n}\right)(s)=s^{n} \mathcal{L}(f)(s)-s^{n-1} f(0)-\ldots-f^{(n-1)}(0)
$$

Let us solve a differential equation using the Laplace transform now. Problem 6 , page 455 asks for the following initial value problem of a second order linear DE with constant coefficients but non-homogeneous: $x^{\prime \prime}+4 x=\cos t, x(0)=x^{\prime}(0)=0$. We first apply the Laplace transform to both sides of the equation and use the formula for the Laplace transform of the derivative of a function: $s^{2} \mathcal{L}(x)(s)-s x(0)-$ $x^{\prime}(0)+4 \mathcal{L}(x)=\frac{s}{s^{2}+1}$. Hence we get $\mathcal{L}(x)(s)\left(s^{2}+4\right)=\frac{s}{s^{2}+1}$. Solving for $\mathcal{L}(x)(s)$ we obtain

$$
\begin{equation*}
\mathcal{L}(x)(s)=\frac{s}{\left(s^{2}+1\right)\left(s^{2}+4\right)} . \tag{5.2}
\end{equation*}
$$

In order to take the inverse Laplace transform we need to write the right hand side of (5.2) in its partial fraction decomposition. There are some shortcuts that
one can use in order to obtain the partial fraction decomposition. These techniques will be discussed in class. In this simple case it is easy to see that we have

$$
\mathcal{L}(x)(s)=\frac{1}{3} \frac{s}{\left(s^{2}+1\right)}-\frac{1}{3} \frac{s}{\left(s^{2}+4\right)}
$$

Equivalently, if we remember the table 5.1.2 of Laplace transforms we can rewrite this equality as

$$
\mathcal{L}(x)(s)=\frac{1}{3} \mathcal{L}(\cos t)(s)-\frac{1}{3} \mathcal{L}(\cos 2 t) .
$$

Because the Laplace transform is linear and injective we conclude that $x(t)=$ $\frac{1}{3} \cos t-\frac{1}{3} \cos 2 t$ for all $t$.

Theorem 5.1.10. If $f$ is piecewise continuous and of exponential type at infinity then

$$
\mathcal{L}\left(\int_{0}^{t} f(x) d x\right)(s)=\frac{1}{s} \mathcal{L}(f)(s) .
$$

Proof. One can see that $g(t)=\int_{0}^{t} f(x) d x$ is continuous and whose derivative is piecewise continuous. It is easy to see that it is also of exponential type. Hence one can apply the Theorem 5.1 .8 to $g: \mathcal{L}\left(g^{\prime}\right)(s)=s \mathcal{L}(g)(s)-g(0)$. This is exactly the identity that we want to establish.

## Homework:

Section 7.2 page 455, Problems 1-37.

### 5.2 Lecture XVI

Quotation: "Mathematics is the search of structure out there in the most incredible places of the human intellect and at the same time apparently unrelated, in all the corners of what presents itself as reality." Anonymous

### 5.3 More properties of the Laplace transform

We have shown how to obtain the Laplace transform for the functions in the table below:

| $f(t)$ | $\mathcal{L}(f)(s)$ |
| :---: | :---: |
| $t^{n}, n \in \mathbb{N}$ | $\frac{n!}{s^{n+1},}, s>0$ |
| $e^{a t} \cos b t$ | $\frac{s-a}{(s-a)^{2}+b^{2}}, s>a$ |
| $e^{a t} \sin b t$ | $\frac{b}{\left(s-a^{2}+b^{2}\right.}, s>a$ |
| $u_{a}(t)$ | $\frac{e^{-s a}}{s}, s>0$ |
| $t^{a}(a>-1)$ | $\frac{\Gamma(a+1)}{s^{a+1}}, s>0$ |
| $e^{z t}$ | $\frac{s}{s-z}, s>\operatorname{Re} z$ |

Using the theorem about the Laplace transform of the derivative of a function we may obtain additional transforms using the technique exemplified in the next example:
Example: Problem 27, page 456. We consider the function $f_{n}(t)=t^{n} e^{z t}$ with $z=a+i b$ and $n \in \mathbb{N}$. Then $f$ is continuous of exponential type $(c=R e z)$ and its derivative exists everywhere, $f_{n}^{\prime}(t)=n t^{n-1} e^{z t}+z t^{n} e^{z t}$, and $f^{\prime}$ is piecewise continuous (in fact continuous on $[0, \infty)$ ). Hence $\mathcal{L}\left(f^{\prime}\right)(s)=s \mathcal{L}(f)(s)-f(0)$ or $n \mathcal{L}\left(f_{n-1}\right)(s)+z \mathcal{L}\left(f_{n}\right)(s)=s \mathcal{L}\left(f_{n}\right)(s)$. Solving for $\mathcal{L}\left(f_{n}\right)(s)$ we get

$$
\mathcal{L}\left(f_{n}\right)(s)=\frac{n}{s-z} \mathcal{L}\left(f_{n-1}\right)(s), s>R e z
$$

This recurrence can be used inductively to prove then that

$$
\begin{equation*}
\mathcal{L}\left(t^{n} e^{z t}\right)(s)=\frac{n!}{(s-z)^{n+1}}, s>\operatorname{Re} z . \tag{5.3}
\end{equation*}
$$

Let us observe that this formula generalizes several of the formulas that we have seen so far but will give two new ones if we take the real part and the imaginary part of both sides $(z=a+i b)$ :

$$
\begin{equation*}
\mathcal{L}\left(t^{n} e^{a t} \cos b t\right)(s)=\frac{n!\sum_{0 \leq j \leq(n+1) / 2}\binom{n+1}{2 j}(-1)^{j}(s-a)^{n+1-2 j} b^{2 j}}{\left[(s-a)^{2}+b^{2}\right]^{n+1}}, \tag{5.4}
\end{equation*}
$$

and

$$
\begin{equation*}
\mathcal{L}\left(t^{n} e^{a t} \sin b t\right)(s)=\frac{n!\sum_{0 \leq j \leq n / 2}\binom{n+1}{2 j+1}(-1)^{j}(s-a)^{n-2 j} b^{2 j+1}}{\left[(s-a)^{2}+b^{2}\right]^{n+1}} . \tag{5.5}
\end{equation*}
$$

Using the theorem about the Laplace transform of the integral of a function we may also obtain some inverse Laplace transforms of functions that contain a power of $s$ at the denominator. We again use an example to exemplify this.

Example: Problem 24, page 456. In this problem we need to find the Laplace transform of $F(s)=\frac{1}{s(s+1)(s+2)}$. Because $\mathcal{L}\left(\int_{0}^{t} f(x) d x\right)(s)=\frac{\mathcal{L}(f)(s)}{s}$ we see that if the right hand side is $F(s)$ we need to find what the inverse Laplace transform just for

$$
G(s)=\frac{1}{(s+1)(s+2)}=\frac{1}{s+1}-\frac{1}{s+2}=\mathcal{L}\left(e^{-t}\right)(s)-\mathcal{L}\left(e^{-2 t}\right)(s)
$$

Therefore $\mathcal{L}^{-1}(G)(t)=e^{-t}-e^{-2 t}$ and our function is $f(t)=\int_{0}^{t} e^{-x}-e^{-2 x} d x=$ $1-e^{-x}-\left(\frac{1}{2}-\frac{1}{2} e^{-2 t}\right)=\frac{1}{2}+\frac{e^{-2 t}}{2}-e^{-t}$.

Next we are going to generalize the theorem about the Laplace transform of the derivative of a function.

Theorem 5.3.1. Suppose $f:[0, \infty) \rightarrow \mathbb{C}$ is piecewise continuous of exponential type (of constant c), which has a derivative $f^{\prime}$ at the points of continuity with the exception of maybe an isolated set of points. Then

$$
\begin{align*}
\mathcal{L}\left(f^{\prime}\right)(s)=s \mathcal{L}(f)(s)-f(0)- & \left.\sum^{t \text { discontinuity }} \begin{array}{l} 
\\
\\
\\
\\
\text { point of } f
\end{array} \quad . \begin{array}{l}
-s t
\end{array} f(t+0)-f(t-0)\right], s>c . \tag{5.6}
\end{align*}
$$

Proof. Let us assume that the origin $t=0$ and the discontinuity points of $f$ are $\left\{t_{n}\right\}_{n \in \mathcal{D}}$; so $t_{1}=0<t_{2}<\ldots, \mathcal{D} \subset \mathbb{N}$. Next we assume first that $f^{\prime}$ exists on each interval $\left(t_{k}, t_{k+1}\right)$. Then $\mathcal{L}\left(f^{\prime}\right)(s)=\lim _{a \rightarrow \infty} \int_{0}^{a} e^{-s t} f^{\prime}(t) d t=\lim _{a \rightarrow \infty}\left[\sum_{n=1}^{n(a)+1} \int_{t_{n}^{\prime}}^{t_{n+1}^{\prime}} e^{-s t} f^{\prime}(t) d t\right]$ where $n(a)$ is the greatest index for which $t_{n(a)}<a$ and $t_{k}^{\prime}=t_{k}$ if $k \leq n(a)$ and $t_{n(a)+1}^{\prime}=a$.

On each interval $\left[t_{n}^{\prime}, t_{n+1}^{\prime}\right]$ we apply the integration by parts to the function which becomes continuous at the endpoints when $f$ is extended with the sided limits:

$$
\begin{aligned}
& \int_{t_{n}^{\prime}}^{t_{n+1}^{\prime}} e^{-s t} f^{\prime}(t) d t=\left.e^{-s t} f(t)\right|_{t=t_{n}^{\prime}+0} ^{t=t_{n+1}^{\prime}-0}+s \int_{t_{n}^{\prime}}^{t_{n+1}^{\prime}} e^{-s t} f(t) d t \text {. Then } \\
& \mathcal{L}\left(f^{\prime}\right)(s)=\lim _{a \rightarrow \infty} s \int_{0}^{a} e^{-s t} f(t) d t+\lim _{a \rightarrow \infty}\left[\sum_{n=1}^{n(a)+1} e^{-s t_{n+1}} f\left(t_{n+1}^{\prime}-0\right)-e^{-s t_{n}} f\left(t_{n}^{\prime}+0\right)\right]
\end{aligned}
$$

. Rearranging the summation and letting $a \rightarrow \infty$ we obtain (5.6). The general case is handled the same way with the observation we made before that the fundamental formula of calculus works under our more relaxed assumptions.

For an application let us work Problem 34, page 456. We apply formula (5.6) for $f(x)=(-1)^{\lfloor x\rfloor}$ where $\lfloor x\rfloor$ is the greatest integer function. Figure 1 below gives an idea of what the graph of $f$ looks like.


Figure 1
This function has a discontinuity for every $n \in \mathbb{N}$ and a jump of $f(2 n+0)$ $f(2 n-0)=1-(-1)=2$ for even discontinuity points and $f(2 n+1+0)-f(2 n+1-$ $0)=-1-1=-2$ for every odd one. In other words $f(n+0)-f(n-0)=2(-1)^{n}$ for every $n \in \mathbb{N}$. Since the derivative is basically zero where it exists applying (5.6) we obtain $0=s \mathcal{L}(f)(s)-f(0)-\sum_{n=1}^{\infty} 2(-1)^{n} e^{-n s}$. Using the formula for summing a sequence in geometric progression $\left(1+r+r^{2}+\ldots=\frac{1}{1-r}\right.$, whenever $\left.r<1\right)$ this turns into

$$
s \mathcal{L}(f)(s)=1+2\left(-e^{-s}\right) \frac{1}{1+e^{-s}}=\frac{1-e^{-s}}{1+e^{-s}} .
$$

Another way of writing this using the hyperbolic functions $\sinh x=\frac{e^{x}-e^{-x}}{2}$, $\cosh x=\frac{e^{x}+e^{-x}}{2}$ is $\mathcal{L}(f)(s)=\frac{1}{s} \frac{e^{s / 2}-e^{-s / 2}}{e^{s / 2}+e^{-s / 2}}=\frac{1}{s} \tanh (s / 2)$.
Theorem 5.3.2. (Translation along the s-axis) If $f$ is such that $\mathcal{L}(f)(s)$ exists for all $s>c$ then $\mathcal{L}\left(e^{a t} f(t)\right)(s)$ exists for all $s>a+c$ and

$$
\mathcal{L}\left(e^{a t} f(t)\right)(s)=\mathcal{L}(f)(s-a)
$$

The proof of this theorem is straightforward. Let us work out a few other examples from this next homework:

Problem 10, page 465. Find the inverse Laplace transform of the function $F(s)=$ $\frac{2 s-3}{9 s^{2}-12 s+20}$.
Observe that

$$
\begin{aligned}
F(s)= & \frac{2 s-3}{(3 s-2)^{2}+16}=\frac{2 s-3}{9\left[(s-2 / 3)^{2}+16 / 9\right]}, \text { or } \\
& F(s)=\frac{2}{9} \frac{s-2 / 3}{(s-2 / 3)^{2}+(4 / 3)^{2}}-\frac{5}{36} \frac{4 / 3}{(s-2 / 3)^{2}+(4 / 3)^{2}} .
\end{aligned}
$$

Hence $\mathcal{L}^{-1}(F)(t)=\frac{2}{9} e^{2 t / 3} \cos 4 t / 3-\frac{5}{36} e^{2 t / 3} \sin 4 t / 3$.
Problem 34, page 465. The DE is $x^{(4)}+13 x^{\prime \prime}+36 x=0$ with initial conditions $x(0)=x^{\prime \prime}(0)=0, x^{\prime}(0)=2$ and $x^{\prime \prime \prime}(0)=-13$. Applying the Laplace transform we get $s^{4} \mathcal{L}(x)-s^{3} x(0)-s^{2} x^{\prime}(0)-s x^{\prime \prime}(0)-x^{\prime \prime \prime}(0)+13\left[s^{2} \mathcal{L}(x)-s x(0)-x^{\prime}(0)\right]+36 \mathcal{L}(x)=0$.

Substituting the initial conditions we get

$$
\mathcal{L}(x)(s)\left(s^{4}+13 s^{2}+36\right)-2 s^{2}+13-26=0 .
$$

Solving for $\mathcal{L}(x)$ gives $\mathcal{L}(x)(s)=\frac{2 s^{2}+13}{s^{4}+13 s^{2}+36}=\frac{s^{2}+4+s^{2}+9}{\left(s^{2}+4\right)\left(s^{2}+9\right)}$, or $\mathcal{L}(x)(s)=$ $\frac{1}{2} \frac{2}{s^{2}+4}+\frac{1}{3} \frac{3}{s^{2}+9}=\frac{1}{2} \mathcal{L}(\sin 2 t)(s)+\frac{1}{3} \mathcal{L}(\sin 3 t)(s)$.

Taking the inverse Laplace transform we obtain $x(t)=\frac{1}{2} \sin 2 t+\frac{1}{3} \sin 3 t$.
Problem 24, page 465 Find the inverse Laplace transform for the function $F(s)=\frac{s}{s^{4}+4 a^{4}}$. Using the idea given in the textbook we factor the denominator of fraction in $F$ :

$$
\begin{aligned}
& F(s)=\frac{s}{\left(s^{2}-2 a s+2 a^{2}\right)\left(s^{2}+2 a s+2 a^{2}\right)}=\frac{s}{\left[(s-a)^{2}+a^{2}\right]\left[(s+a)^{2}+a^{2}\right]} \\
& =\frac{1}{4 a}\left(\frac{1}{(s-a)^{2}+a^{2}}-\frac{1}{(s+a)^{2}+a^{2}}\right)=\frac{1}{4 a^{2}}\left(\frac{a}{(s-a)^{2}+a^{2}}-\frac{a}{(s+a)^{2}+a^{2}}\right) .
\end{aligned}
$$

Hence $\mathcal{L}^{-1}(F)(t)=\frac{1}{4 a^{2}}\left(e^{a t} \sin a t-e^{-a t} \sin a t\right)$ or

$$
\mathcal{L}^{-1}(F)(t)=\frac{1}{2 a^{2}} \sinh a t \sin a t .
$$

## Homework:

Section 7.3 page 455, Problems 1-38.

### 5.4 Lecture XVII

Quotation: "The key to make progress in the process of learning mathematics is to ask the right questions." Anonymous

### 5.4.1 Convolution of two functions

Let us assume we have two functions, $f$ and $g$, which are piecewise continuous and of exponential type (with the same constant $c$ )

Definition 5.4.1. The convolution of $f$ and $g$ is the new function $(f \star g)(t)=$ $\int_{0}^{t} f(x) g(t-x) d x, t \geq 0$.

The convolution defined this way is commutative: $f \star g=g \star f$. This can be easily seen by a change of variables: $y=t-x$,

$$
\begin{aligned}
& (f \star g)(t)=\int_{0}^{t} f(x) g(t-x) d x=\int_{t}^{0} f(t-y) g(y)(-d y) \\
& =\int_{0}^{t} g(y) f(t-y) d y=(g \star f)(t) .
\end{aligned}
$$

Theorem 5.4.2. The convolution of the two functions of exponential type (with constant c) is also of exponential type (with constant $c+\epsilon$ ). The Laplace transform of the convolution of two functions is the product of the individual Laplace transforms:

$$
\mathcal{L}(f \star g)(s)=\mathcal{L}(f)(s) \mathcal{L}(g)(s), s>c
$$

Proof. Let us denote by $F(s)$ the Laplace transform of $f$ and by $G(s)$ the Laplace transfrom of $g$. If $s>c$ then $F(s) G(s)=\int_{0}^{\infty} e^{-s t} f(t) d t \int_{0}^{\infty} e^{-s x} g(x) d x$. The function of two variables $(t, x)->e^{-s t} f(t) e^{-s x} g(x)$ is absolutely integrable over the domain $[0, \infty) \times[0, \infty)$ with the same proof as we did when we showed the existence of the Laplace transform. Then we can rewrite $F(s) G(s)=\int_{0}^{\infty} \int_{0}^{\infty} e^{-s(t+x)} f(t) g(x) d t d x$. We can make a substitution now $t=u$ and $t+x=v$. The domain $[0, \infty) \times[0, \infty)$ can now be described as $\{(v, u): v \in[0, \infty)$ and $u \in[0, v]\}$. The Jacobian of the transformation is $J(x, t)=\operatorname{det}\left[\begin{array}{cc}1 & 0 \\ -1 & 1\end{array}\right]=1$, so, the double integral becomes

$$
F(s) G(s)=\int_{0}^{\infty} e^{-s v}\left[\int_{0}^{v} f(u) g(v-u) d u\right] d v=\mathcal{L}(f * g)(s) .
$$

The double integral can be understood in a limit sense (over rectangles) which makes the above computation possible.

Let us use this theorem to solve Problem 14, page 474. We need to find the inverse Laplace transform of the function $F(s)=\frac{s}{s^{4}+5 s^{2}+4}$. We can rewrite $F(s)=\frac{s}{\left(s^{2}+1\right)\left(s^{2}+4\right)}=\mathcal{L}(\sin t)(s) \mathcal{L}(\cos 2 t)(s)$

So, $\mathcal{L}^{-1}(F)(t)=\int_{0}^{t} \sin x \cos 2(t-x)=\int_{0}^{t} \frac{1}{2}[\sin (2 t-x)+\sin (3 x-2 t)] d x$. Then

$$
\begin{aligned}
& \mathcal{L}^{-1}(F)(t)=\left.\frac{1}{2} \cos (2 t-x)\right|_{x=0} ^{x=t}-\left.\frac{1}{6} \cos (3 x-2 t)\right|_{x=0} ^{x=t}= \\
& \frac{1}{2}(\cos t-\cos 2 t)-\frac{1}{6}(\cos t-\cos 2 t)=\frac{1}{3}(\cos t-\cos 2 t), t \geq 0 .
\end{aligned}
$$

Theorem 5.4.3. [Integration of Transform formula] Suppose that $f$ is piecewise continuous for $t \geq 0$, has exponential type at infinity (with constant $c$ ) and that $\lim _{t \rightarrow 0^{+}} \frac{f(t)}{t}$ exists. If the Laplace transform of $f$ is $F$, then the improper integral $\int_{s}^{\infty} F(u) d u$ exists for every $s>c$ and

$$
\mathcal{L}\left(\frac{f(t)}{t}\right)(s)=\int_{s}^{\infty} F(u) d u, s>c .
$$

Proof. Since $F(s)=\int_{0}^{\infty} e^{-s t} f(t) d t$ the function $F$ is continuous. The improper integral $\int_{s}^{\infty} F(u) d u$ exists because of the estimate we got when we proved the existence of the Laplace transform. Then

$$
\int_{s}^{\infty} F(u) d u=\int_{s}^{\infty} \int_{0}^{\infty} e^{-u t} f(t) d t d u
$$

It turns out that the function of two variables $e^{-u t} f(t)$ is integrable on the domain $[s, \infty) \times[0, \infty)$ in the sense of limits on arbitrary rectangles and so the interchange of the integrals is possible. Thus

$$
\int_{s}^{\infty} F(u) d u=\int_{0}^{\infty}\left[\int_{s}^{\infty} e^{-u t} d u\right] f(t) d t=\int_{0}^{\infty} e^{-s t} \frac{f(t)}{t} d t
$$

As we can see the hypothesis that $\lim _{t \rightarrow 0^{+}} \frac{f(t)}{t}$ exists can be relaxed to the existence of the integral $\int_{0}^{1} \frac{|f(t)|}{t} d t$.

We are going to work out Problem 20, page 474. We need to find the Laplace transform of $g(t)=\frac{1-\cos 2 t}{t}$. Consider the map $f(t)=1-\cos 2 t$. We have $\mathcal{L}(f)(s)=$
$\frac{1}{s}-\frac{s}{s^{2}+4}$. Since $\lim _{t \rightarrow 0^{+}} \frac{f(t)}{t}=0$ we can apply Theorem 5.4.3 and obtain $\mathcal{L}(g)(s)=$ $\int_{s}^{\infty} \frac{1}{u}-\frac{u}{u^{2}+4} d u=\left.\ln \frac{u}{\sqrt{u^{2}+4}}\right|_{s} ^{\infty}=\ln \frac{\sqrt{s^{2}+4}}{s}$ for all $s>0$.

Theorem 5.4.4. [Differentiation of the transform] If $f$ is piecewise continuous and of exponential type (with constant c), then if $F$ is the Laplace transform of $f$ we have

$$
\mathcal{L}(-t f(t))(s)=F^{\prime}(s), \quad s>c
$$

Proof. Since $F(s)=\int_{0}^{\infty} e^{-s t} f(t) d t$ and $\mathcal{L}(-t f(t))(s)=\int_{0}^{\infty} e^{-s t}(-t f(t)) d t$ exist, we can calculate

$$
\frac{F(s)-F\left(s_{0}\right)}{s-s_{0}}-\mathcal{L}(-t f(t))\left(s_{0}\right)=\int_{0}^{\infty} \frac{e^{-s t}-e^{-s_{0} t}-\left(s-s_{0}\right)(-t) e^{-s_{0} t}}{s-s_{0}} f(t) d t
$$

Using the generalized mean value theorem: $h(b)=h(a)+(b-a) h^{\prime}(a)+$ $\frac{(b-a)^{2}}{2} h^{\prime \prime}(\xi)$ for some $\xi \in(a, b)$, we obtain $\left(h(u)=e^{-u t}, a=s_{0}, b=s\right)$

$$
\frac{e^{-s t}-e^{-s_{0} t}-\left(s-s_{0}\right)(-t) e^{-s_{0} t}}{s-s_{0}}=t^{2} \frac{s-s_{0}}{2} e^{-\xi(s) t}, \xi(s) \in\left(s_{0}, s\right) .
$$

Hence

$$
\begin{equation*}
\left|\frac{F(s)-F\left(s_{0}\right)}{s-s_{0}}-\mathcal{L}(-t f(t))\left(s_{0}\right)\right| \leq \frac{\left|s-s_{0}\right|}{2} \int_{0}^{\infty} e^{-s_{0} t} t^{2}|f(t)| d t \rightarrow 0 \tag{5.7}
\end{equation*}
$$

as $s$ tends to $s_{0}$. The integral $\int_{0}^{\infty} e^{-s_{0} t} t^{2}|f(t)| d t$ is finite if $s_{0}>c$, fact that goes the same way as the existence of the Laplace transform. Then passing to the limit in (5.7) $\left(s \rightarrow s_{0}\right)$ we get $\mathcal{L}(-t f(t))(s)=F^{\prime}(s), \quad s>c$.

Applying this theorem several time we get:
Corollary 5.4.5. Under the same assumptions of Theorem 5.4.4, for every $n \in \mathbb{N}$, $\mathcal{L}\left(t^{n} f(t)\right)(s)=(-1)^{n} F^{(n)}(s), s>c$.

Example: Problem 26, page 474. We need to calculate the inverse Laplace transform of $F(s)=\arctan \frac{3}{s+2}$. Since $F^{\prime}(s)=\frac{-\frac{3}{(s+2)^{2}}}{\frac{9}{(s+2)^{2}}+1}=-\frac{3}{(s+2)^{2}+9}$. If $f$ is the inverse Laplace transform of $F$ then by Theorem 5.4.4 we get $-t f(t)=$ $\mathcal{L}^{-1}\left(-\frac{3}{(s+2)^{2}+9}\right)=-e^{-2 t} \sin 3 t$. This gives $f(t)=e^{-2 t} \frac{\sin 3 t}{t}$.

Another application of this formula is finding a nontrivial solution of the Bessel's type equation in Problem 34: $t x^{\prime \prime}+(4 t-2) x^{\prime}+(13 t-4) x=0, x(0)=0$. We denote by $X(s)=\mathcal{L}(x(t))(s)$. We have $\mathcal{L}\left(x^{\prime \prime}\right)(s)=s^{2} X(s)-a$ where $x^{\prime}(0)=a$, and $\mathcal{L}\left(x^{\prime}\right)(s)=s X(s)$. Hence $\mathcal{L}\left(t x^{\prime \prime}\right)=-\frac{d}{d s}\left(s^{2} X(s)-a\right)=-2 s X(s)-s^{2} X^{\prime}(s)$, $\mathcal{L}\left(t x^{\prime}\right)(s)=-X(s)-s X^{\prime}(s)$ and $\mathcal{L}(t x)=-X^{\prime}(s)$.

Then the equation becomes $-2 s X(s)-s^{2} X^{\prime}(s)-4 X(s)-4 s X^{\prime}(s)-2 s X(s)-$ $13 X^{\prime}(s)-4 X(s)=0$. This reduces to a simple differential equation in $X(s): \frac{X^{\prime}(s)}{X(s)}=$ $-\frac{8+4 s}{(s+2)^{2}+9}$.

Integrating we get $\ln |X(s)|=-2 \ln \left[(s+2)^{2}+9\right]+C$ and from here $X(s)=$ $\frac{k}{\left[(s+2)^{2}+9\right]^{2}}$. Because $\mathcal{L}^{-1}\left(\frac{k}{\left[(s+2)^{2}+9\right]}\right)(t)=\frac{k}{3} e^{-2 t} \sin 3 t$, then we can use the convolution formula to get $x(t)=\frac{k}{9} \int_{0}^{t}\left(e^{-2 u} \sin 3 u\right)\left(e^{-2(t-u)} \sin 3(t-u)\right) d u=\frac{k e^{-2 t}}{18} \int_{0}^{t} \cos (6 u-$ $3 t)-\cos 3 t d u$. Therefore $x(t)=\frac{k e^{-2 t}}{18}\left(\left.\frac{\sin (6 u-3 t)}{6}\right|_{0} ^{t}-t \cos 3 t\right)=\frac{k e^{-2 t}}{54}(\sin 3 t-3 t \cos 3 t)$ or $x(t)=A e^{-2 t}(\sin 3 t-3 t \cos 3 t), t \geq 0$.

Now we are going to review all the important formulae that we have introduced so far:

| $f(t)$ on $[0, \infty)$ | $F(s)=\mathcal{L}(f)(s)$ |
| :---: | :---: |
| $\|f(t)\| \leq M e^{c t}$ | $\int_{0}^{\infty} e^{-s t} f(t) d t, s>c$ |
| $t^{n}, n \in \mathbb{N}$ | $\frac{n!}{s^{n+1}}, s>0$ |
| $e^{a t} \cos b t$ | $\frac{s-a}{(s-a)^{2}+b^{2}}, s>a$ |
| $e^{a t} \sin b t$ | $\frac{b}{(s-a)^{2}+b^{2}}, s>a$ |
| $t e^{a t} \cos b t$ | $\frac{(s-a)^{2}-b^{2}}{\left.\\|(s-a)^{2}+b^{2}\right]^{2}}, s>a$ |
| $t e^{a t} \sin b t$ | $\frac{2 b(s-a)}{(s-a)^{2}+\left.b^{2}\right\|^{2}}, s>a$ |
| $u_{a}(t)$ | $\frac{e^{-s a}}{s}, s>0$ |
| $t^{a}(a>-1)$ | $\frac{\Gamma(a+1)}{s^{a+1}}, s>0$ |
| $e^{z t}$ | $\frac{1}{s-z}, s>\operatorname{Re} z$ |
| $t^{n} e^{z t}$ | $\frac{n!}{(s-z)^{n+1}}, s>\operatorname{Re} z$ |
| $e^{z t} f(t)$ | $F(s-z), s>c+\operatorname{Rez}$ |
| $f^{\prime}(t)$ | $s F(s)-f(0), s>c$ |
| $(f * g)(t)$ | $F(s) G(s), s>c$ |
| $t f(t)$ | $-F^{\prime}(s)$, |
| $\int_{0}^{t} f(x) d x$ | $F(s) / s, s>c$ |
| $f(t) / t$ | $\int_{s}^{\infty} F(u) d u, s>c$ |
| some less important |  |
| $(-1)^{\lfloor x\rfloor}$ | $\frac{1}{s} \tanh \frac{s}{2}, s>0$ |
| $\frac{1}{2 b^{3}} e^{a t}(\sin b t-b t \cos b t)$ | $\frac{1}{\left[(s-a)^{2}+b^{2}\right]^{2}}$ |
| $\frac{1}{2 b} e^{a t}(\sin b t+b t \cos b t)$ | $\frac{(s-a)^{2}}{\left[(s-a)^{2}+b^{2}\right]^{2}}$ |
| $e^{a t} \cosh b t$ | $\frac{s-a}{(s-a)^{2}-b^{2}}, s>a$ |
| $e^{a t} \sinh b t$ | $\frac{b}{(s-a)^{2}-b^{2}}, s>a$ |

## Homework:

Section 7.3 page 474, Problems 1-38.

### 5.5 Lecture XVIII

Quotation: "When a truth is necessary, the reason for it can be found by analysis, that is, by resolving it into simpler ideas and truths until the primary ones are reached. It is this way that in mathematics speculative theorems and practical canons are reduced by analysis to definitions, axioms and postulates. " (Leibniz, 1670)

### 5.5.1 Periodic and piecewise continuous input functions

We have already showed that $\mathcal{L}\left(u_{a}\right)(s)=\frac{e^{-a s}}{s}, s>0$, where

$$
u_{a}(t)=\left\{\begin{array}{l}
0 \text { for } t<a \\
1 \text { for } t \geq a
\end{array}\right.
$$

Theorem 5.5.1. Let us consider $a \geq 0$. If $\mathcal{L}(f)(s)$ exists for $s>c$ then

$$
\mathcal{L}\left(u_{a}(t) f(t-a)\right)(s)=e^{-s a} \mathcal{L}(f)(s), \text { for } s>a \text {. }
$$

Proof. This is just a simple calculation:

$$
\begin{aligned}
& \mathcal{L}\left(u_{a}(t) f(t)\right)(s)=\int_{0}^{\infty} e^{-s t} u_{a}(t) f(t-a) d t=\int_{a}^{\infty} e^{-s t} f(t-a) d t= \\
& \int_{0}^{\infty} e^{-s(u+a)} f(u) d u=e^{-s a} \mathcal{L}(f)(s)
\end{aligned}
$$

for all $s>c$.
Let us observe that if $0<a<b$ then $u_{a}-u_{b}$ is the function:

$$
u_{a, b}(t)=\left\{\begin{array}{l}
0 \text { for } x<a \\
1 \text { for } x \in[a, b) \\
0 \text { for } x \geq b
\end{array}\right.
$$

whose graph is


Figure 12
and it is called the characteristic function of the interval $[a, b)$. Let us solve Problem 18, page 484: we need to compute the Laplace transform of

$$
f(t)=\left\{\begin{array}{l}
\cos \frac{1}{2} \pi t \text { if } 3 \leq t \leq 5 \\
0 \text { if } t<3 \text { or } t>5
\end{array}\right.
$$

This function is essentially the same as $t \rightarrow u_{3,5}(t) \cos \frac{1}{2} \pi t$ for $t \in[0, \infty)$. Hence the Laplace transform of it is $\mathcal{L}\left(u_{3}(t) \cos \frac{t \pi}{2}\right)-\mathcal{L}\left(u_{5}(t) \cos \frac{t \pi}{2}\right)$. Because $\cos \frac{t \pi}{2}=$ $\cos \left(\frac{(t-3) \pi}{2}+\frac{3 \pi}{2}\right)=\sin \frac{(t-3) \pi}{2}$ and similarly for $\cos \frac{t \pi}{2}=\cos \left(\frac{(t-5) \pi}{2}+\frac{5 \pi}{2}\right)=-\sin \frac{(t-5) \pi}{2}$, we obtain that $\mathcal{L}(f)(s)=\left(e^{-5 t}+e^{-3 t}\right) \frac{2 \pi}{4 s^{2}+\pi^{2}}$.

The last theorem that we are going to do is about the transform of a periodic function:

Theorem 5.5.2. [Laplace transform of periodic function] If $f$ is periodic piecewise continuous with period $p$ on $[0, \infty)$ the Laplace transfrom of $f$ exists and

$$
\mathcal{L}(f)(s)=\frac{1}{1-e^{-p s}} \int_{0}^{p} e^{-s t} f(t) d t, s>0 .
$$

Proof. This is also a calculation:

$$
\begin{aligned}
& \mathcal{L}(f)(s)=\int_{0}^{\infty} e^{-s t} f(t) d t=\sum_{k=0}^{\infty} \int_{k p}^{k(p+1)} e^{-s t} f(t) d t \\
& \sum_{k=0}^{\infty} \int_{0}^{p} e^{-s t-k p s} f(t+k p) d t=\sum_{k=0}^{\infty} e^{-k p s} \int_{0}^{p} e^{-s t} f(t) d t=\frac{1}{1-e^{-p s}} \int_{0}^{p} e^{-s t} f(t) d t
\end{aligned}
$$

using the sum of the geometric progression $\sum_{k=0}^{\infty} e^{-k p s}=\frac{1}{1-e^{-s p}}$.
We are going to use this Theorem to compute the Laplace transform of the function in Problem 28, page 485. The graph of $f$ (for $a=1$ is shown below):


Figure 2
Basically we need to compute

$$
\int_{0}^{2 a} e^{-s t} f(t) d t=\int_{0}^{a} e^{-s t} t d t=\left.e^{-s t}\left(\frac{t}{-s}-\frac{1}{s^{2}}\right)\right|_{t=0} ^{t=a}=\frac{1-(1+s a) e^{-s a}}{s^{2}}
$$

So, the Laplace transform of $f$ is $\mathcal{L}(f)(s)=\frac{1-(1+s a) e^{-s a}}{s^{2}\left(1-e^{-2 s a}\right)}$.

### 5.5.2 Impulses and delta function

Definition 5.5.3. The Dirac delta function at a, denoted by $\delta_{a}$ is a transformation on continuous functions defined by $\delta_{a}(g)=g(a)$ for every continuous function.

In general most of the maps $\varphi$ having properties of linearity and bounded on continuous functions defined for $t \in[0, \infty)$ is of the form $\varphi(g)=\int_{0}^{\infty} g(t) h(t) d t$. The $\operatorname{map} \delta_{a}$ is an example not of this form. When we have a differential equation of the type, let's say, as in Problem 2, page 495, $x^{\prime \prime}+4 x=\delta_{0}+\delta_{\pi}$, with initial conditions $x(0)=x^{\prime}(0)=0$, we interpret this as the model of movement of a mass $(m=1)$ attached to a spring with no dashpot with two instantaneous blows of unit intensity at moments $t=0$ and $t=\pi$.

So if we apply both functions to the function $t \rightarrow e^{-s t}$ we get $s^{2} X(s)+4 X(s)=$ $1+e^{-s \pi}$. Then we solve for $X(s)=\frac{1}{s^{2}+4}+\frac{e^{-s \pi}}{s^{2}+4}$ and then take the inverse Laplace transform $x(t)=\frac{\sin 2 t}{2}+u_{\pi}(t) \frac{\sin 2(t-\pi)}{2}$ or $x(t)=\left(1+u_{\pi}(t) \frac{\sin 2 t}{2}\right.$

The graph of this solution is shown below. This solution is still a continuous function but it is not differentiable at every point.


Figure 3

Now we summarize all the important Laplace transform formulae that we have studied so far:

| $f(t)$ on $[0, \infty)$ | $F(s)=\mathcal{L}(f)(s)$ |
| :---: | :---: |
| $\|f(t)\| \leq M e^{c t}$ | $\int_{0}^{\infty} e^{-s t} f(t) d t, s>c$ |
| $t^{n}, n \in \mathbb{N}$ | $\frac{n!}{s^{n+1}}, s>0$ |
| $e^{a t} \cos b t$ | $\frac{s-a}{(s-a)^{2}+b^{2}}, s>a$ |
| $e^{a t} \sin b t$ | $\frac{b}{(s-a)^{2}+b^{2}}, s>a$ |
| $t e^{a t} \cos b t$ | $\frac{(s-a)^{2}-b^{2}}{(s-a)^{2}+b^{2}}, s>a$ |
| $t e^{a t} \sin b t$ | $\frac{2 b(s-a)}{(s-a)^{2}+b^{2}}, s>a$ |
| $u_{a}(t)$ | $\frac{e^{-s a}}{s}, s>0$ |
| $t^{a}(a>-1)$ | $\frac{\Gamma(a+1)}{s^{a+1}}, s>0$ |
| $e^{z t}$ | $\frac{1}{s-z}, s>R e z$ |
| $t^{n} e^{z t}$ | $\frac{n!}{(s-z)^{n+1}}, s>\operatorname{Re} z$ |
| $e^{z t} f(t)$ | $F(s-z), s>c+\operatorname{Rez}$ |
| $f^{\prime}(t)$ | $s F(s)-f(0), s>c$ |
| $(f * g)(t)$ | $F(s) G(s), s>c$ |
| $t f(t)$ | $-F^{\prime}(s)$, |
| $\int_{0}^{t} f(x) d x$ | $F(s) / s, s>c$ |
| $f(t) / t$ | $\int_{s}^{\infty} F(u) d u, s>c$ |
| some less important |  |
| $(-1)^{\lfloor x\rfloor}$ | $\frac{\tanh \frac{s}{2}}{s}, s>c$ |
| $\frac{1}{2 b^{3}} e^{a t}(\sin b t-k t \cos b t)$ | $\frac{s}{\left[(s-a)^{2}+b^{2}\right]^{2}}$ |
| $\frac{1}{2 b} e^{a t}(\sin b t+k t \cos b t)$ | $\frac{(s-a)^{2}}{\left[(s-a)^{2}+b^{2}\right]^{2}}$ |
| $e^{a t} \cosh b t$ | $\frac{s-a}{(s-a)^{2}-b^{2}}, s>a$ |
| $e^{a t} \sinh b t$ | $\frac{b}{(s-a)^{2}-b^{2}}, s>a$ |
| $u_{a}(t) f(t-a)$ | $e^{-s a} \mathcal{L}(f)(s)$ |
| $f(t)$ periodic f of period p | $\frac{1}{1-e^{-s p}} \int_{0}^{p} f(t) d t$ |

## Homework:

Section 7.5 page 484, Problems 1-35.
Section 7.6, pages 495, Problems 1-8.

## Chapter 6

## Power Series Methods

### 6.1 Lecture XIX

Quotation: "The heart of mathematics is its problems." Paul Halmos

### 6.1.1 Power series review

The method that we are going to study in this Chapter applies to a variety of DE such as the Bessel's equation (of order $n$ ),

$$
x^{2} y^{\prime \prime}+x y^{\prime}+\left(x^{2}-n^{2}\right) y=0,
$$

or Legendre's equation

$$
\left(1-x^{2}\right) y^{\prime \prime}-2 x y^{\prime}+n(n+1) y=0
$$

which appear in many applications.
A power series around the point $x=a$ is an infinite sum of the form

$$
\begin{equation*}
\sum_{n=0}^{\infty} a_{n}(z-a)^{n} \tag{6.1}
\end{equation*}
$$

where as usual the convergence is understood in the usual sense, i.e.

$$
\lim _{n \rightarrow \infty} \sum_{k=0}^{n} a_{k}(z-a)^{k}
$$

exists.
The series (2.27) defines a function $f(z)$ on a disc of radius $R$ (called radius of convergence ) centered at $a: D_{a}(R):=\{z| | z-a \mid<R\}$.

The radius of convergence is given by the formula:

$$
\begin{equation*}
\frac{1}{R}=\limsup _{n \rightarrow \infty}\left|a_{n}\right|^{\frac{1}{n}} \tag{6.2}
\end{equation*}
$$

The series (6.1) converges at least for $z=a$ but we are going to be interested in series for which the radius of convergence is a positive real number or infinity. Most of the elementary functions have a power series expansion around any point which is not a singularity (all the derivative are defined there). The function defined by a power series is continuous and differentiable on $D_{a}(R)$. Moreover the derivative can be computed differentiating term by term. The derivative has the same radius of convergence and hence the function is infinitely many times differentiable. The coefficients $a_{n}$ are given then by the formula:

$$
a_{n}=\frac{f^{(n)}(a)}{n!}, n \geq 0
$$

This allows one to compute various power series for most of the elementary functions. Two power series can be added or subtracted term by term. This corresponds to adding or subtracting the corresponding functions. The product has to be done in the Cauchy sense. The following theorem is important:

Theorem 6.1.1. [Identity Principle] If $\sum_{n=0}^{\infty} a_{n} x^{n}=\sum_{n=0}^{\infty} b_{n} x^{n}$ for every $x$ in some open interval then $a_{n}=b_{n}$ for all $n \geq 0$.

### 6.1.2 Series solutions around ordinary points

We are going to consider solving the DE

$$
\begin{equation*}
y^{\prime \prime}+P(x) y^{\prime}+Q(x) y=0 \tag{6.3}
\end{equation*}
$$

where $P$ and $Q$ are functions defined around point $a$. If these functions have a power series expansion around $a$ then the point $x=a$ is called an ordinary point for (6.3). A point $a$ will be called singular for (6.3) if at least one of the functions $P$ or $Q$ is not analytic around $a$ (which means there is no power series centered at $a$ that sums up to the given function). The next theorem shows what happens in the situation of ordinary points.

Theorem 6.1.2. There are two linearly independent solutions of (6.3) around every ordinary point whose radius of convergence is at least as large as the distance from a to the nearest (real or complex) singular point of (6.3).

Let us solve the Legendre Equation:

$$
\begin{equation*}
\left(1-x^{2}\right) y^{\prime \prime}-2 x y^{\prime}+\alpha(\alpha+1) y=0 \tag{6.4}
\end{equation*}
$$

or if we put it in the form (6.3) we get

$$
y^{\prime \prime}+\frac{-2 x}{1-x^{2}} y^{\prime}+\frac{\alpha(\alpha+1)}{1-x^{2}} y=0,
$$

which makes it clear that $a=0$ is an ordinary point for (6.4). According to the theorem above there are two linearly independent solutions that can be written as power series whose radius of convergence is at least 1 . Let us look for a solution of the form $y(x)=\sum_{k=0}^{\infty} a_{k} x^{k}$.

Then the equation (6.4) becomes

$$
\left(1-x^{2}\right) \sum_{k=0}^{\infty} k(k-1) a_{k} x^{k-2}-2 x \sum_{k=0}^{\infty} a_{k} k x^{k-1}+\alpha(\alpha+1) \sum_{k=0}^{\infty} a_{k} x^{k}=0
$$

or

$$
\sum_{k=0}^{\infty}(k+2)(k+1) a_{k+2} x^{k}-\sum_{k=0}^{\infty} k(k-1) a_{k} x^{k}-2 \sum_{k=0}^{\infty} a_{k} k x^{k}+\alpha(\alpha+1) \sum_{k=0}^{\infty} a_{k} x^{k}=0 .
$$

Using the identity principle we get the following equations in terms of the coefficients $a_{k}$ :

$$
(k+2)(k+1) a_{k+2}-k(k-1) a_{k}-2 k a_{k}+\alpha(\alpha+1) a_{k}=0, k \geq 0 .
$$

This gives

$$
\begin{equation*}
a_{k+2}=\frac{(k-\alpha)(k+\alpha+1)}{(k+1)(k+2)} a_{k}, k \geq 0 . \tag{6.5}
\end{equation*}
$$

Instead of continuing this in the most general case we are going to make an assumption here that may help to see how something interesting could happen here.

Let's say $\alpha=3$. Then (6.5) gives $a_{2}=-\frac{12}{2} a_{0}=-6 a_{0}, a_{3}=-\frac{10}{6} a_{1}=-\frac{5}{3} a_{1}$, $a_{4}=-\frac{6}{12} a_{2}=3 a_{0}, a_{5}=0$. From here we see that the next coefficients $a_{2 k+1}=0$ for $k \geq 2$. So, one of the solutions is $y_{1}(x)=a_{1}\left(x-\frac{5}{3} x^{3}\right)=\frac{a_{1}}{3}\left(3 x-5 x^{3}\right)=-\frac{2 a_{1}}{3} P_{3}(x)$ where $P_{3}(x)=\frac{1}{2}\left(5 x^{3}-3 x\right)$ is the called Legendre polynomial of degree 3. Similarly for every $\alpha$ a non-negative integer $n$ one of the solutions is just going to be a polynomial which turns out to be the Legendre polynomial of degree $n$.

Next we rewrite (6.3) as

$$
\begin{equation*}
y^{\prime \prime}+\frac{p(x)}{x} y^{\prime}+\frac{q(x)}{x^{2}} y=0 \tag{6.6}
\end{equation*}
$$

Definition 6.1.3. The singular point $x=0$ of (6.3) is a regular singular point if the functions $p$ and $q$ are both analytic around 0 . Otherwise 0 is an irregular singular point.

### 6.1.3 The Method of Frobenius

We are going to use a slightly modified version of power series method to solve differential equations of second order for which $x=0$ is regular singular point. As before consider the equation written in the form

$$
\begin{equation*}
x^{2} y^{\prime \prime}+x p(x) y^{\prime}+q(x) y=0 . \tag{6.7}
\end{equation*}
$$

The idea is to look simply for a solution of the form

$$
\begin{equation*}
x^{r} \sum_{k=0}^{\infty} a_{k} x^{k}, \quad x>0 . \tag{6.8}
\end{equation*}
$$

We have the following theorem:
Theorem 6.1.4. Suppose that $x=0$ is a regular singular point for (6.7) and let $p(x)=\sum_{k=0}^{\infty} p_{k} x^{k}$ and $q(x)=\sum_{k=0}^{\infty} p_{k} x^{k}$ be the power series representations of $p$ and $q$. If we denote the solutions of the quadratic equation $r(r-1)+p_{0} r+q_{0}=0$ by $r_{1}$ and $r_{2}$ then:
(a) if $r_{1}$ and $r_{2}$ are real, say $r_{1} \geq r_{2}$, there exist a solution of the form (6.8) with $r=r_{1}$;
(b) if $r_{1}$ and $r_{2}$ are real, say $r_{1} \geq r_{2}$, and $r_{1}-r_{2}$ is not an integer (i.e. $\left(p_{0}-1\right)^{2}-4 q_{0}$ is not the square of an integer) the there exists a second linearly independent solution of (6.7) of the form (6.8) with $r=r_{2}$.

Let us solve Problem 18, page 535. We need to solve the DE: $2 x y^{\prime \prime}+3 y^{\prime}-y=0$. In this case if we write this equation in the form (6.7) we get $x^{2} y^{\prime \prime}+\frac{3}{2} x y^{\prime}-\frac{x}{2} y=0$. This gives $x=0$ as a regular singular point and $p(x)=\frac{3}{2}$ and $q(x)=\frac{x}{2}$. Hence the equation in $r$ becomes $r(r-1)+\frac{3}{2} r=0$ which has two solutions: $r_{1}=0$ and $r_{2}=-\frac{1}{2}$.

Therefore, according to the Theorem 6.1.4 we must have two linearly independent solutions of the form (6.8). Working out the details of this we get

$$
y_{1}(x)=\sum_{k=0}^{\infty} \frac{x^{k}}{k!(2 k+1)!!},
$$

and

$$
y_{2}(x)=\frac{1}{\sqrt{x}}\left[1+\sum_{k=1}^{\infty} \frac{x^{k}}{k!(2 k-1)!!}\right] .
$$

## Homework:

Section 8.1, Problems pages 509-510, 23, 25 and 27;
Section 8.2, page 520, Problems 5, 6, 32, 35.
Section 8.3, page 5535, Problems 1-31, 35, 38 and 39.

### 6.2 Lecture XX

### 6.2.1 When $r_{1}-r_{2}$ is an integer

We remind the reader the type of differential equation to which we have applied the method of Frobenius:

$$
\begin{equation*}
y^{\prime \prime}+\frac{p(x)}{x} y^{\prime}+\frac{q(x)}{x^{2}} y=0 . \tag{6.9}
\end{equation*}
$$

where $x=0$ is a regular singular point of (6.11), i.e., the two functions $p$ and $q$ are analytic around $x=0$.

We are going to take an example from the text to study what may happen in the situation $r_{1}-r_{2}$ is a positive integer.

Problem 28, page 535: $x y^{\prime \prime}+2 y^{\prime}-4 x y=0$. In this particular case $p(x)=2$ and $q(x)=-4 x^{2}$. The equation for $r$ (indicitial equation ) becomes $r(r-1)+$
$2 r=0$ with solutions $r_{1}=0$ and $r_{2}=-1$. This makes $r_{1}-r_{2}$ an integer. We are going to study the existence of the second solution: $y=x^{-1} \sum_{k=0}^{\infty} a_{k} x^{k}$. Since $y^{\prime}=\sum_{k=0}^{\infty}(k-1) a_{k} x^{k-2}$ and $y^{\prime \prime}=\sum_{k=0}^{\infty}(k-1)(k-2) a_{k} x^{k-3}$, after we substitute in the given equation we get:

$$
\sum_{k=0}^{\infty}(k-1)(k-2) a_{k} x^{k-2}+\sum_{k=0}^{\infty} 2(k-1) a_{k} x^{k-2}-\sum_{k=0}^{\infty} 4 a_{k} x^{k}=0
$$

or

$$
\sum_{k=0}^{\infty}(k-1) k a_{k} x^{k-2}-\sum_{k=0}^{\infty} 4 a_{k} x^{k}=0
$$

Since the first two terms in the first summation are zero we obtain only one summation if we shift the index $(k-2 \rightarrow k)$ and then combine the two: $\sum_{k=0}^{\infty}[(k+$ 1) $\left.(k+2) a_{k+2}-4 a_{k}\right] x^{k}=0$.

This gives $a_{k+2}=4 \frac{a_{k}}{(k+1)(k+2)}$ for all $k \geq 0$.
From here we see that $a_{2 n}=\frac{4^{n}}{(2 n)!} a_{0}$ and $a_{2 n+1}=\frac{4^{n}}{(2 n+1)!} a_{1}$ for all $n \geq 0$.
Therefore a general solution of our equation is

$$
y(x)=a_{0} x^{-1} \sum_{n=0}^{\infty} \frac{4^{n} x^{2 n}}{(2 n)!}+a_{1} \sum_{n=0}^{\infty} \frac{4^{n} x^{2 n}}{(2 n+1)!} .
$$

Let us observe that we actually get an analytic solution and one which is unbounded around $x=0$. Using the functions sinh and cosh we can re-write the general solution as

$$
y(x)=a_{0} x^{-1} \cosh 2 x+a_{1} \frac{\sinh 2 x}{2 x}
$$

So, in this case we have two solutions in the form (6.8).
To show that there are cases in which there is only one solution of the form (6.8) let us take Problem 39, page 536:
(a) Show that the Bessel's equation of order 1,

$$
x^{2} y^{\prime \prime}+x y^{\prime}+\left(x^{2}-1\right) y=0
$$

has exponents $r_{1}=1$ and $r_{2}=-1$ at $x=0$, and the Frobenius series corresponding to $r=1$ is

$$
J_{1}(x)=\frac{x}{2} \sum_{n=0}^{\infty} \frac{(-1)^{n} x^{2 n}}{n!(n+1)!2^{2 n}}
$$

(b) Show that there is no Frobenius solution corresponding to the smaller exponent $r_{2}=-1$; that is, it is impossible to determine the coefficients in

$$
\begin{equation*}
y_{2}(x)=x^{-1} \sum_{n=0}^{\infty} c_{n} x^{n} \tag{6.10}
\end{equation*}
$$

Solution: Let us start by differentiating and substituting in the Bessel's equation with (6.10) as recommended (calculations will cover both cases): $y_{2}^{\prime}(x)=$ $\sum_{n=0}^{\infty} c_{n}(n-1) x^{n-2}$ and $y_{2}^{\prime \prime}(x)=\sum_{n=0}^{\infty} c_{n}(n-1)(n-2) x^{n-3}$.

The Bessel's equation becomes

$$
\sum_{n=0}^{\infty} c_{n}(n-1)(n-2) x^{n-1}+\sum_{n=0}^{\infty} c_{n}(n-1) x^{n-1}+\left(x^{2}-1\right) \sum_{n=0}^{\infty} c_{n} x^{n-1}=0
$$

The first two sums can be combined together and together with the last sum after multiplication by $x^{2}-1$ and distributing:

$$
\sum_{n=0}^{\infty} c_{n}\left[n^{2}-2 n\right] x^{n-1}+\sum_{n=0}^{\infty} c_{n} x^{n+1}=0
$$

Shifting the index of summation in the first sum we get

$$
\sum_{n=-2}^{\infty} c_{n+2}\left(n^{2}+2 n\right) x^{n+1}+\sum_{n=0}^{\infty} c_{n} x^{n+1}=0
$$

For $n=-2$ we get $c_{0} \times 0=0$ which is satisfied for every $c_{0}$. For $n=-1$ we obtain $c_{1}=0$. For $n=0$ we get $c_{2} \times 0+c_{0}=0$ which implies $c_{0}=0$. For $n \geq 1$ we have $c_{n+2}=-\frac{c_{n}}{n(n+2)}$. This implies $c_{2 n+1}=0$ for all $n \geq 0$ and

$$
c_{2 n+2}=-\frac{c_{2 n}}{2 n(2 n+2)}=-\frac{c_{2 n}}{n(n+1) 2^{2}}=\frac{c_{2 n-2}}{(n-1) n n(n+1) 2^{4}}=\ldots=\frac{(-1)^{n} c_{2}}{n!(n+1)!2^{2 n}},
$$

for all $n \geq 1$. The final form of $y_{2}$ is

$$
\begin{aligned}
& y_{2}(x)=x^{-1} \sum_{n=0}^{\infty} c_{n} x^{n}=x^{-1} \sum_{n=1}^{\infty} c_{2 n} x^{2 n}=x \sum_{n=1}^{\infty} c_{2 n} x^{2 n-2}= \\
& =x \sum_{n=0} c_{2 n+2}^{\infty} x^{2 n}=c_{2} x \sum_{n=0}^{\infty} \frac{(-1)^{n} x^{2 n}}{n!(n+1)!2^{2 n}}=2 c_{2} J_{2}(x)
\end{aligned}
$$

This shows both parts (a) and (b) of the problem.
In general the equation

$$
\begin{equation*}
x^{2} y^{\prime \prime}+x p(x) y^{\prime}+q(x) y=0 \tag{6.11}
\end{equation*}
$$

has a second solution which is described by the next theorem:
Theorem 6.2.1. [ The Exceptional Case] Assume $x=0$ is a regular singular point for (6.11) and $r_{1} \geq r_{2}$ are the two roots of $r^{2}+\left(p_{0}-1\right) r+q_{0}=0$.
(a) If $r_{1}=r_{2}$ then the equation (6.11) has two linearly independent solutions of the form:

$$
\begin{gathered}
y_{1}(x)=x^{r_{1}} \sum_{n=0}^{\infty} a_{n} x^{n},\left(a_{0} \neq 0\right), \\
y_{2}(x)=y_{1}(x) \ln x+x^{r_{1}+1} \sum_{n=0}^{\infty} b_{n} x^{n} .
\end{gathered}
$$

(b) If $r_{1}-r_{2}=N$ with $N \in \mathbb{N}$, then the equation (6.11) has two linearly independent solutions of the form:

$$
\begin{gathered}
y_{1}(x)=x^{r_{1}} \sum_{n=0}^{\infty} a_{n} x^{n},\left(a_{0} \neq 0\right), \\
y_{2}(x)=C y_{1}(x) \ln x+x^{r_{2}} \sum_{n=0}^{\infty} b_{n} x^{n} .
\end{gathered}
$$

## Homework:

Section 8.3, page 535, Problems 1-31, 35, 38 and 39.
Section 8.4, pages 551-552, Problems 1-8, 18, and 21.

## Chapter 7

## Fourier Series

### 7.1 Lecture XXI

Quotation: "Even fairly good students, when they have obtained the solution of the problem and written down neatly the argument, shut their books and look for something else. Doing so, they miss an important and instructive phase of the work. ... A good teacher should understand and impress on his students the view that no problem whatever is completely exhausted. " George Pólya

### 7.1.1 Fourier series, definition and examples

Another type of expansions for functions that can be helpful in computing solutions of differential equations is the Fourier series expansion. The method of using a different type of expantion works basically the same way as with power series: substitute in the given differential equation, find a recurrence for the coefficients and then use that to determine the coefficients and the function if possible. In general a function that has a Fourier expansion will have to be periodic. So, it is natural to work with periodic functions defined on $\mathbb{R}$ and we will take for simplicity the period to be $2 \pi$.
Definition 7.1.1. Assume $f$ is a piecewise continuous function of period $2 \pi$ defined on $\mathbb{R}$. The Fourier series of $f$ is

$$
\begin{equation*}
\frac{a_{0}}{2}+\sum_{n=1}^{\infty}\left(a_{n} \cos n t+b_{n} \sin n t\right) \tag{7.1}
\end{equation*}
$$

where $a_{n}=\frac{1}{\pi} \int_{-\pi}^{\pi} f(t) \cos n t d t$ for $n=0,1,2,3, \ldots$ and $b_{n}=\frac{1}{\pi} \int_{-\pi}^{\pi} f(t) \sin n t d t$ for $n=1,2,3, \ldots$ are called the Fourier coefficients.

Let us see an example. Suppose we take Example 21, page 580. The function is $f(t)=t^{2}$ for $t \in[-\pi, \pi]$. Then the Fourier coefficients of $f$ are $a_{0}=\frac{1}{\pi} \int_{-\pi}^{\pi} t^{2} d t=$ $2 \frac{\pi^{3}}{3 \pi}=\frac{2 \pi^{2}}{3}$,

$$
\begin{aligned}
& a_{n}=\frac{1}{\pi} \int_{-\pi}^{\pi} f(t) \cos n t d t=\frac{1}{\pi}\left(\left.t^{2} \frac{\sin n t}{n}\right|_{-\pi} ^{\pi}+\left.2 t \frac{\cos n t}{n^{2}}\right|_{-\pi} ^{\pi}+\left.2 \frac{\sin n t}{n^{3}}\right|_{-\pi} ^{\pi}\right) \\
& =\frac{4(-1)^{n}}{n^{2}}
\end{aligned}
$$

for $n=1,2,3, \ldots$ and

$$
b_{n}=\frac{1}{\pi} \int_{-\pi}^{\pi} f(t) \sin n t d t=\frac{1}{\pi}\left(-\left.t^{2} \frac{\cos n t}{n}\right|_{-\pi} ^{\pi}+\left.2 t \frac{\sin n t}{n^{2}}\right|_{-\pi} ^{\pi}+\left.2 \frac{\cos n t}{n^{3}}\right|_{-\pi} ^{\pi}\right)=0
$$

for $n=1,2,3, \ldots$. We will see later that this gives the following formula

$$
\begin{equation*}
t^{2}=\frac{\pi^{2}}{3}+\sum_{n=1}^{\infty} \frac{\cos 2 n t}{n^{2}}-4 \sum_{n=0}^{\infty} \frac{\cos (2 n+1) t}{(2 n+1)^{2}} \tag{7.2}
\end{equation*}
$$

which we will call the Fourier series expansion of $f$. We have to assign a meaning to the series in (7.2). As usual, we will understand by it the limit of the partial sums. If one plots the partial sums of (7.2) against $t \rightarrow t^{2}$ (in our plot on $[-3 \pi, 3 \pi]$ and taking only five terms in each sum) will get


Figure 1
This suggests that the series converges to actually the given function. This usually is the case if the function is more than continuous (not true for continuous functions only) and the convergence is uniform if the function has a derivative which is piecewise continuous.
Theorem 7.1.2. [Dirichlet] Suppose $f$ is a periodic function of period $2 \pi$ which is piecewise differentiable. The Fourier series converges
(a) to the value of $f(t)$ for every value $t$ where $f$ is continuous;
(b) to the value $\frac{1}{2}(f(t+0)+f(t-0))$ at each point of discontinuity

Let us take an example where the function is discontinuous. We consider the function

$$
g(t)=\left\{\begin{array}{l}
-1 \quad \text { for } t \in(-\pi, 0) \\
1 \quad \text { for } t \in[0, \pi]
\end{array}\right.
$$

extended by periodicity for all real axis. Then $a_{n}=0$ for all $n=0,1,2,3,4, \ldots$ and $b_{n}=\frac{2}{\pi} \int_{0}^{\pi} \sin n t d t=\frac{2\left(1-(-1)^{n}\right)}{n \pi}$ for all $n=1,2,3,4, \ldots$.
Then $g(t)=\frac{4}{\pi} \sum_{n=0}^{\infty} \frac{\sin (2 n+1) t}{2 n+1}$ for all $t \in(-\pi, \pi]\{0\}$. We can see that the part (b) of the Theorem 7.1.2 is satisfied. By taking $t=\frac{\pi}{2}$, we observe that this is a point of continuity for $g$ and $g\left(\frac{\pi}{2}\right)=1$ and hence the Theorem 7.1.2 implies that $1=\frac{4}{\pi} \sum_{n=0}^{\infty} \frac{(-1)^{n}}{2 n+1}$ or

$$
\begin{equation*}
\frac{\pi}{4}=\sum_{n=0}^{\infty} \frac{(-1)^{n}}{2 n+1} \tag{7.3}
\end{equation*}
$$

One of the important formulae that one needs in the calculation of the Fourier coefficients is given in Problem 22, page 587: show that if $p(t)$ is a polynomial of degree $n$, and $g$ is a continuous function,

$$
\begin{equation*}
\int p(t) g(t) d t=p(t) G_{1}(t)-p^{\prime}(t) G_{2}(t)+\ldots+(-1)^{n} p^{(n)}(t) G_{n+1}(t)+C \tag{7.4}
\end{equation*}
$$

where $G_{k+1}$ is the antiderivative of $G_{k}$ for all $k=0,1, \ldots, n$ and $G_{0}=g$.
This can be checked by differentiation:

$$
\begin{aligned}
& \frac{d}{d t}\left(p(t) G_{1}(t)-p^{\prime}(t) G_{2}(t)+\ldots+(-1)^{n} p^{(n)}(t) G_{n+1}(t)\right)= \\
& p^{\prime}(t) G_{1}(t)-p^{\prime \prime}(t) G_{2}(t)+\ldots+(-1)^{n} p^{(n+1)}(t) G_{n+1}(t)+ \\
& p(t) g(t)-p^{\prime}(t) G_{1}(t)+\ldots+(-1)^{n} p^{(n)}(t) G_{n}(t)=p(t) g(t)
\end{aligned}
$$

### 7.2 General Fourier Series

In general if we have a function which is periodic of period $2 L$ then we can still expand it in terms of trigonometric functions but we need to change the period.

Definition 7.2.1. If $f$ is a piecewise continuous function of period $2 L$ then the Fourier series of $f$ is

$$
\begin{equation*}
\frac{a_{0}}{2}+\sum_{n=1}^{\infty}\left(a_{n} \cos \frac{n \pi t}{L}+b_{n} \sin \frac{n \pi t}{L}\right) \tag{7.5}
\end{equation*}
$$

where $a_{n}=\frac{1}{L} \int_{-L}^{L} f(t) \cos \frac{n \pi t}{L} d t$ for $n=0,1,2,3, \ldots$ and $b_{n}=\frac{1}{\pi} \int_{-L}^{L} f(t) \sin \frac{n \pi t}{L} d t$ for $n=1,2,3, \ldots$ are called the Fourier coefficients of $f$ on $[0,2 L]$.

A similar theorem to Theorem 7.1.2 takes place in the case of periodic functions of period $2 L(L>0)$. Let us look at the Problem 17, page 587. The function $f$ is periodic of period 2 and defined by $f(t)=t$ for $t \in(0,2)$. We want to show that

$$
\begin{equation*}
f(t)=1-\frac{2}{\pi} \sum_{n=1}^{\infty} \frac{\sin n \pi t}{n} \tag{7.6}
\end{equation*}
$$

In this cases $L=1$. Let us compute first $a_{0}=\int_{0}^{2} t d t=\left.\frac{t^{2}}{2}\right|_{0} ^{2}=2$. For $n \geq 1$ we have $a_{n}=\int_{0}^{2} t \cos n \pi t d t$. Using formula (7.4) we get

$$
a_{n}=\left.t \frac{\sin n \pi t}{n \pi}\right|_{0} ^{2}+\left.\frac{\cos n \pi t}{n^{2} \pi^{2}}\right|_{0} ^{2}=0
$$

For $n \geq 1$ we have $b_{n}=\int_{0}^{2} t \sin n \pi t d t$. Similarly we get

$$
b_{n}=-\left.t \frac{\cos n \pi t}{n \pi}\right|_{0} ^{2}+\left.\frac{\sin n \pi t}{n^{2} \pi^{2}}\right|_{0} ^{2}=-\frac{2}{n \pi},
$$

and so (7.6) takes place. Substituting $t=1 / 2$ in (7.6) will give

$$
\frac{\pi}{4}=\sum_{n=0} \frac{(-1)^{n}}{2 n+1}
$$

which is nothing but the Leibniz's identity (series) (7.3).
From formula (7.2) let us derive another important series that is so common in mathematics. Denote by $x$ the sum of the series $\sum_{n=1}^{\infty} \frac{1}{n^{2}}$.
Substituting $t=\pi$ in (7.2) we get $\pi^{2}=\frac{\pi^{2}}{3}+4 x$ which will give $x=\frac{\pi^{2}}{6}$. Therefore

$$
\sum_{n=1}^{\infty} \frac{1}{n^{2}}=\frac{\pi^{2}}{6}
$$

a formula discovered by Euler. Euler and Leibnitz identities seem to be such curious facts in mathematics since they relate all the whole numbers with the number $\pi$.

## Homework:

Section 9.1, page 580, Problems 1-31.
Section 9.2, pages 586-587, Problems 1-25.

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[^0]:    1.1 The man and his dog

